High Statistics Quasielastic AntiNeutrino Scattering at MINERvA

Amit Bashyal For the MINERvA Collaboration 10th March, 2023



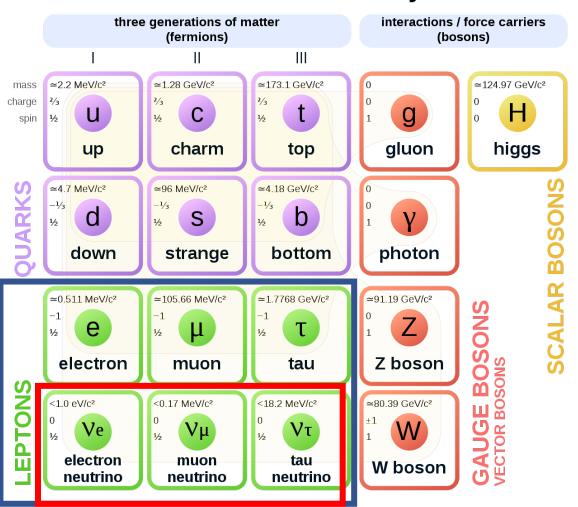
arxiv.org/abs/2211.10402



*Now at Argonne National Lab

Neutrinos In Standard Model

Standard Model of Elementary Particles



Neutrinos are electrically neutral leptons.

3 Generations of charged leptons —>3 generations of neutral massless neutrinos

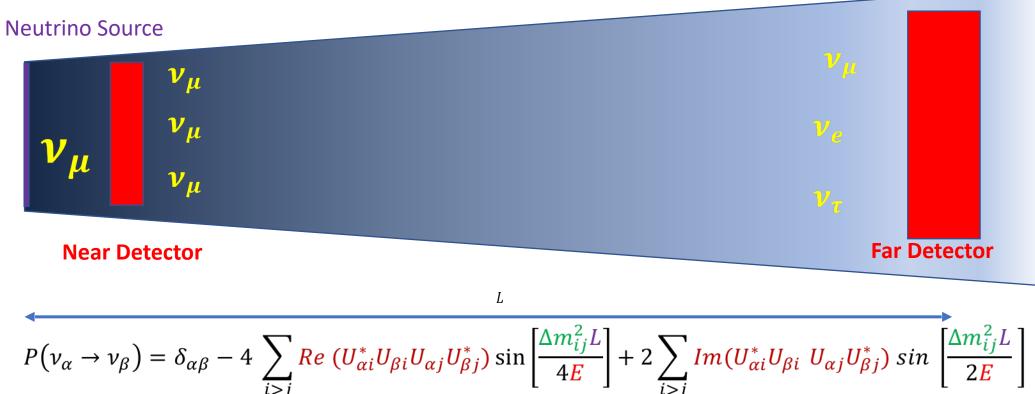
$$W^{+} \rightarrow e^{+} + \nu_{e} \qquad W^{-} \rightarrow e^{-} + \bar{\nu}_{e}$$

$$W^{+} \rightarrow \mu^{+} + \nu_{\mu} \qquad W^{+} \rightarrow \mu^{-} + \bar{\nu}_{\mu}$$

$$W^{+} \rightarrow \tau^{+} + \nu_{\tau} \qquad W^{+} \rightarrow \tau^{-} + \bar{\nu}_{\tau}$$

• ν_e , ν_μ , $\nu_ au$ are only interact through weak force.

Neutrino Oscillation Experiments



Decreasing flux intensity

L→ Distance Between neutrino source and far detector

 $E \rightarrow$ Energy of the neutrino

 $U_{\alpha i}, U_{\beta i} \rightarrow \text{Mixing matrix elements } (\alpha, \beta \rightarrow \nu_e, \nu_\mu, \nu_\tau \& i, j \rightarrow \nu_1, \nu_2, \nu_3)$

 $\Delta m_{ij}^2 = m_i^2 - m_i^2$ \rightarrow Difference of mass of eigenstates v_i, v_j

Neutrino Oscillation: Measurement

In a $u_{\mu}
ightarrow
u_{e}$ oscillation experiment, quantity we want to measure is:

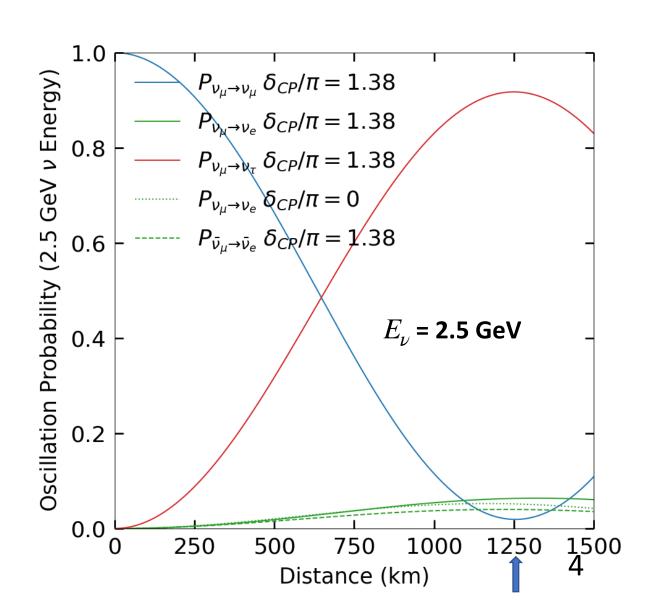
$$P_{\nu_{\mu}\to\nu_{e}}(E,L) \propto \frac{\Phi_{\nu_{e}}(E,L)}{\Phi_{\nu_{\mu}}(E,0)}$$

 $\Phi_{\nu_e}(E, L) \to \text{Electron neutrino flux measured}$ at distance L from the source

 $\Phi_{\nu_{\mu}}(E,0) \rightarrow$ Muon neutrino flux measured at the source

One of the goals of current oscillation experiments is to measure how much CP violation happens in lepton sector $\delta_{\text{CP}} = 0 \rightarrow \text{No CP Violation} \quad P(\nu_{\mu} \rightarrow \nu_{e}) = P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e})$ $\delta_{\text{CP}} \neq 0 \rightarrow \text{CP Violation} \quad P(\nu_{\mu} \rightarrow \nu_{e}) \neq P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e})$

Neutrino experiments don't measure flux directly. They measure the number of events N as a function of reconstructed neutrino energy E_{rec} .



Neutrino Oscillation: Measurement

$$[\sigma(E_{\nu})^* \theta(E_{\nu})] = \begin{bmatrix} \text{Cross section is convoluted by nuclear effects } (\theta(E_{\nu})) \\ \text{Ne}(E_{rec},L) \propto \sum_{\pmb{i}} \Phi_{\pmb{e}}(E,L) \sigma_{\pmb{i}}(E) f_{\pmb{\sigma}\pmb{i}}(E,Erec) dEdM \end{bmatrix}$$

Estimation of neutrino flux: MINERvA employs various methods to accurately estimate neutrino flux and constrain flux related uncertainties.

Neutrino Interaction Cross section: Cross section measurements improve our neutrino interaction models

(Accelerator neutrino) oscillation experiments like NOvA, DUNE, T2K,HK need accurate estimations of neutrino flux $(\Phi(E))$ and cross section $(\sigma(E))$ to extract the oscillation parameters accurately.

Nuclear effects complicate the cross-section measurements.

Smearing of true neutrino energy (E): Limitation on how accurately neutrino energy can be reconstructed (E_{rec})

Smearing of true (E) to reconstructed (E_{rec}) depends upon detector properties.

Neutrino Oscillation: Measurement

$$N_e(E_{rec}, L) \propto \sum \Phi_e(E, L) \sigma_i(E) f_{\sigma i}(E, Erec) dEdM$$

i : neutrino-nucleus interaction channel

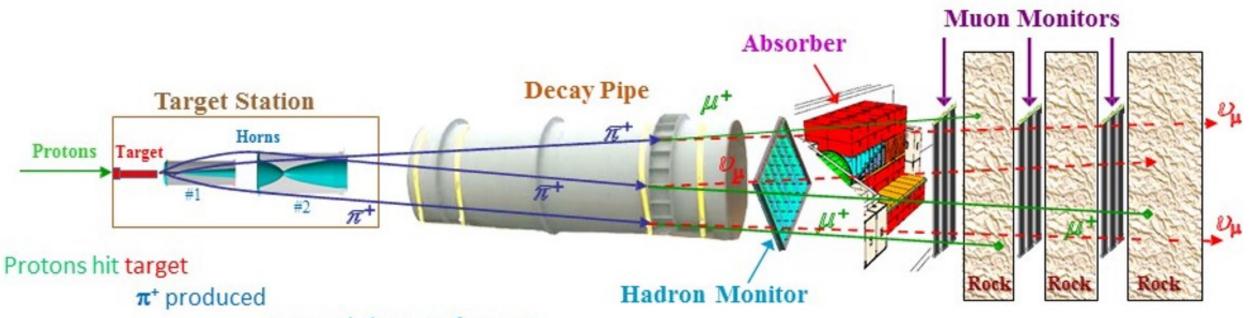
Estimation of neutrino flux: MINERvA employs various methods to accurately estimate neutrino flux and constrain flux related uncertainties.

Neutrino Interaction Cross section: Cross section measurements improve our neutrino interaction models

MINERvA was designed to study the poorly understood neutrino-heavy nucleus interactions

- ☐ MINERvA has developed a comprehensive flux strategy to constrain the neutrino flux uncertainties.
- □ MINERvA has provided cross section measurements of various neutrino scattering processes over a wide neutrino energy range.
- ☐ Measurements in various target materials has helped us to understand nuclear effects in neutrino-heavy nucleus interactions.
- Neutrino oscillation experiments often use heavy target material detectors. Understanding nuclear effects is crucial to reconstruct neutrino energy in near and far detectors.

Neutrino Production in NuMI Beamline



magnetic horn to focus π^*

 π^+ decay to $\mu^+ \nu$ in long evacuated pipe

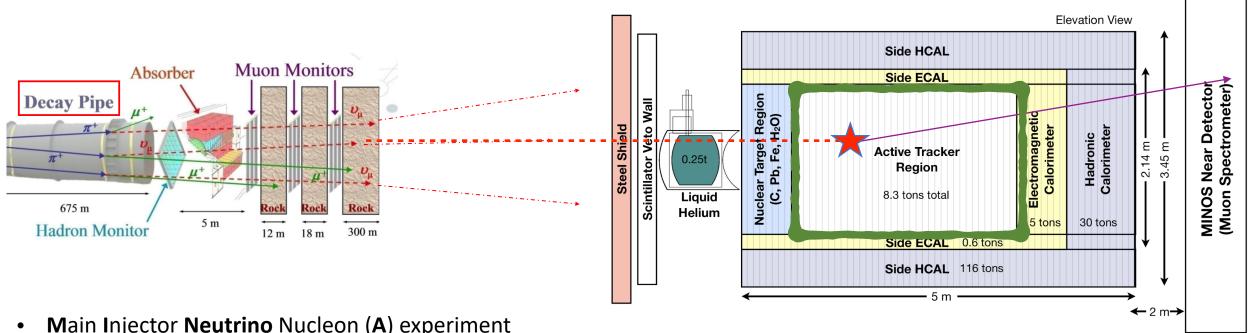
left-over hadrons shower in hadron absorber

rock shield ranges out μ*

v beam travels through earth to experiment

Image Credit: Fermilab

MINERvA Experiment



- Main Injector Neutrino Nucleon (A) experiment
- MINERvA is a neutrino cross-section measurement experiment
- Downstream of the NuMI beam line

- MINERvA detector → 1.032 km from the NuMI target
- MINOS near detector \rightarrow 1.04 km from the NuMI target
 - MINOS near detector is magnetized.
 - Identify charge of muon from interaction

MINERvA Experiment



Detector shutting down for the last time.

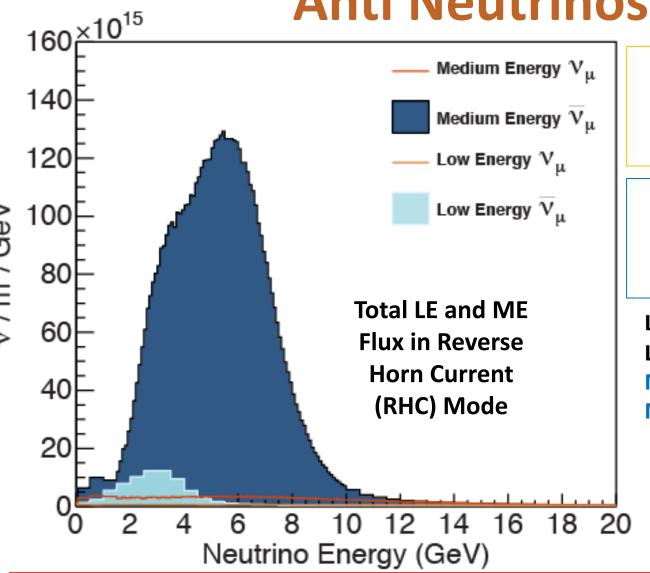
Front face of the MINERvA detector





Celebrating for collecting 30e20 POT

Anti Neutrinos in MINERVA



ME (Medium Energy) = 6
GeV Neutrino Energy
Focusing Peak Run

LE (Low Energy) = 3 GeV Neutrino energy Focusing Peak Run

LE ν POT: 4. 0×10^{20} LE $\overline{\nu}$ POT: 1. 7×10^{20}

$$\label{eq:local_potential} \begin{split} \text{ME} \, \nu \textit{POT} \, \sim \, 3 \times \textit{LE} \, \nu \, \textit{POT} \\ \text{ME} \, \overline{\nu} \, \textit{POT} \, \sim \, 7 \times \textit{LE} \, \overline{\nu} \, \textit{POT} \end{split}$$

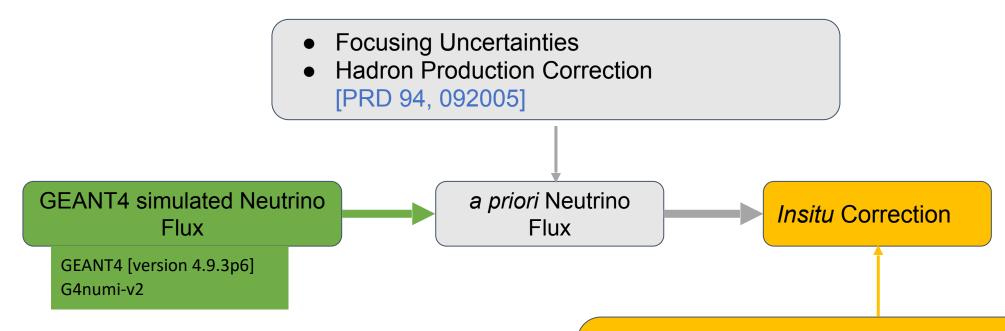
LE \rightarrow 2 years run ME \rightarrow 5.5 years run

ME data:

High Statistics+ High Energy

RHC: Focusing Horns magnetic field to focus π^- which in turn decay to give $\bar{\nu}_\mu$.

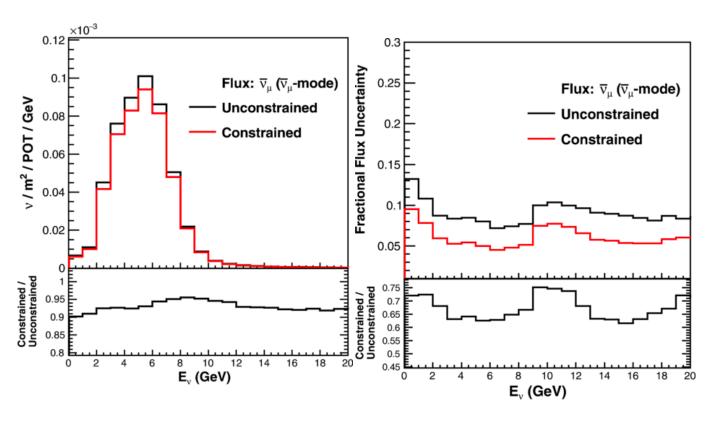
MINERvA Flux Strategy



- G4numi simulated Flux Using FTFP_BERT Hadronic Model
 - O GEANT4 model dependent
- Hadron Production Corrections using Thin Target Datasets [Na49, Barton]

- ν+e constraint [L. Zazutsa, PRD 107,012001]
 - \circ $\nu e \rightarrow \nu e$
- Inverse Muon Decay [D. Ruterbories, PRD 104,092010]
 - \circ ν_{μ} e $\rightarrow \mu^{-}\nu_{e}$
- Combined Fit of (FHC+RHC νe scattering)+ IMD Data
- Uncertainty reduced from 7.8% to 4.7% ($\bar{\nu}_{\mu}$ Flux)

Insitu correction on the Flux



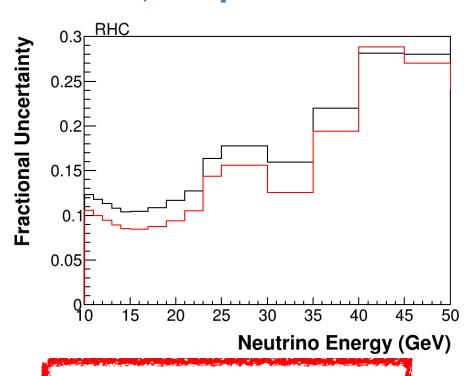
a priori anti neutrino Flux before and after constrained by ν e scattering data Uncertainty on predicted anti neutrino flux before and after constrained by ν e scattering data

Combined Fit of (FHC+RHC νe scattering) + IMD Data Uncertainty reduced from 7.8% to 4.7% ($\bar{\nu}_{\mu}$ Flux)

ν+e constraint [L. Zazutsa, PRD 107,012001]

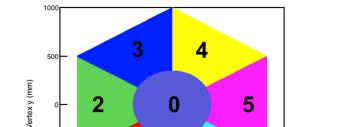
$$\circ$$
 $\nu e \rightarrow \nu e$

 Inverse Muon Decay [D. Ruterbories, PRD 104,092010]

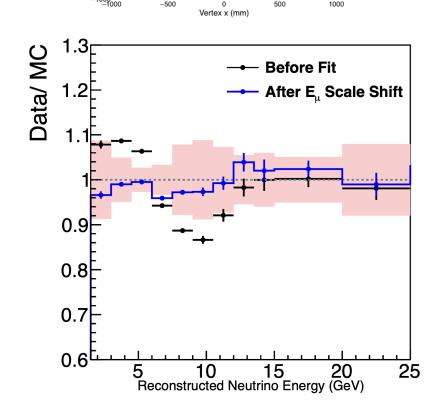


IMD constrains Flux with E_{ν} >10 GeV

Low nu Fit to resolve discrepancy



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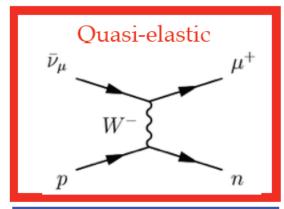
- MINERvA saw discrepancy between its data and simulated sample.
- Discrepancy due to shift in energy spectrum
- A multi parameter fit
 - Fit done with Low recoil sample
 - Cross-section independent of E_{ν} —> Shape depends on Flux
 - Focusing parameters and MINOS muon energy scale as fit parameters
- \bullet Shift of MINOS muon energy scale by 1.8 $\sigma\,$ to resolve discrepancy

[A. Bashyal,arXiv:2104.05769]

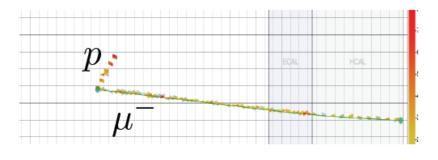
Parameter	Nominal	Best Fit (No Prior)	Best Fit (Prior)
Beam Position (X)	0.0 mm	$-0.3 \pm 0.3 \pm 0.1 \text{ mm}$	$-0.3 \pm 0.2 \pm 0.1 \text{ mm}$
Beam Position (Y)	0.0 mm	$0.8 \pm 0.3 \pm 0.3 \text{ mm}$	$0.7 \pm 0.2 \pm 0.2 \text{ mm}$
Target Position (X)	0.0 mm	$-0.8 \pm 0.3 \pm 0.1 \text{ mm}$	$-0.8 \pm 0.3 \pm 0.1 \text{ mm}$
Target Position (Y)	0.0 mm	$2.3 \pm 0.7 \pm 1.2 \text{ mm}$	$1.7 \pm 0.6 \pm 0.8 \mathrm{mm}$
Target Position (Z)	-1433 mm	$-1432.4 \pm 2.4 \pm 0.3 \text{ mm}$	$-1431 \pm 1.8 \pm 0.3 \text{ mm}$
Horn 1 Position (X)	0.0 mm	$-0.3 \pm 0.4 \pm 0.5 \text{ mm}$	$-0.1 \pm 0.3 \pm 0.1 \text{ mm}$
Horn 1 Position (Y)	0.0 mm	$0.1 \pm 0.5 \pm 0.5 \text{ mm}$	$0.0 \pm 0.3 \pm 0.3 \text{ mm}$
Beam Spot Size	1.5 mm	$1.41 \pm 0.09 \pm 0.03$ mm	$1.32 \pm 0.09 \pm 0.03$ mm
Horn Water Layer	1.0 mm	$1.2 \pm 0.3 \pm 0.05 \text{ mm}$	$1.3 \pm 0.25 \pm 0.1 \text{ mm}$
Horn Current	200 kA	$198.0 \pm 1.4 \pm 1.4 \text{ kA}$	$199.1 \pm 0.7 \pm 0.5 \text{ kA}$
Muon Energy Scale	1.0	$1.032 \pm 0.004 \pm 0.008$	$1.036 \pm 0.004 \pm 0.006$

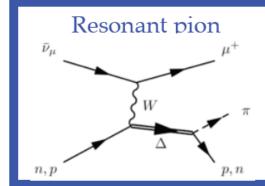
$\bar{\nu}_{\mu}$ Charged Current Quasi Elastic Cross-section Measurement in CH Target

General Neutrino Nucleon Interactions in MINERvA

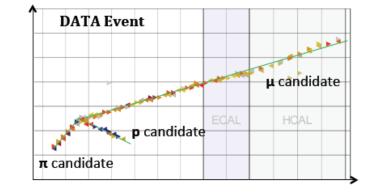


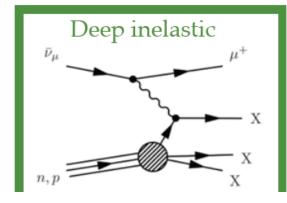
QE



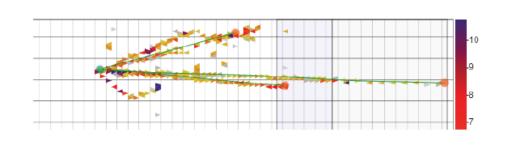


RES

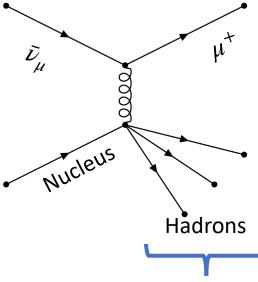




DIS



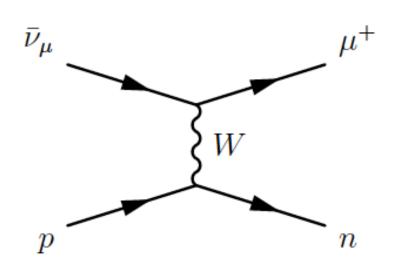
system



Recoil hadrons (Inv. mass W)

$$W^2 = M^2 + 2M\nu - Q^2$$

CCQE Interactions

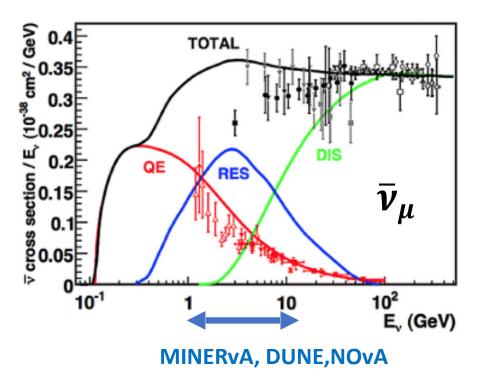


Feynman diagram of an anti muon neutrino CCQE interaction

CCQE Processes

$$\nu_l + n \rightarrow l^- + p$$

$$\bar{\nu_l} + p \to l^+ + n$$



Oscillation experiments like
NOvA and DUNE (will) see
CCQE interactions as one of the
major interactions.

by MINERvA will help oscillation experiment to understand their data.

Z.A Formaggio and Z.P. Zeller, Rev. Mod. Phys. 84, 1307-1341, 2012

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Anti-neutrino nucleon cross-section (per nucleon) for different interaction channels.

Red: QE cross-section

Neutrino Nucleon QE Scattering

- CCQE: Relatively clean process
- Assuming the nucleon is at rest, the energy of the incoming neutrino can be reconstructed using the kinematics of outgoing muon.

$$E_{\nu}^{QE} = \frac{m_n^2 - (m_p - E_b)^2 - m_{\mu}^2 + 2(m_p - E_b)E_{\mu}}{2(m_p - E_b - E_{\mu} + p_{\mu}\cos\theta_{\mu})}$$
$$Q_{QE}^2 = 2E_{\nu}^{QE}(E_{\mu} - p_{\mu}\cos\theta_{\mu}) - m_{\mu}^2$$

 $m_p \rightarrow mass\ of\ proton$

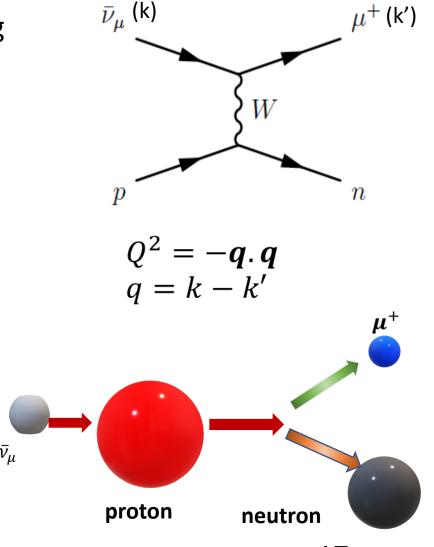
 $m_n \rightarrow mass\ of\ neutron$

 $E_b \rightarrow Binding \ Energy \ of \ the \ neutron \ in \ nucleus$

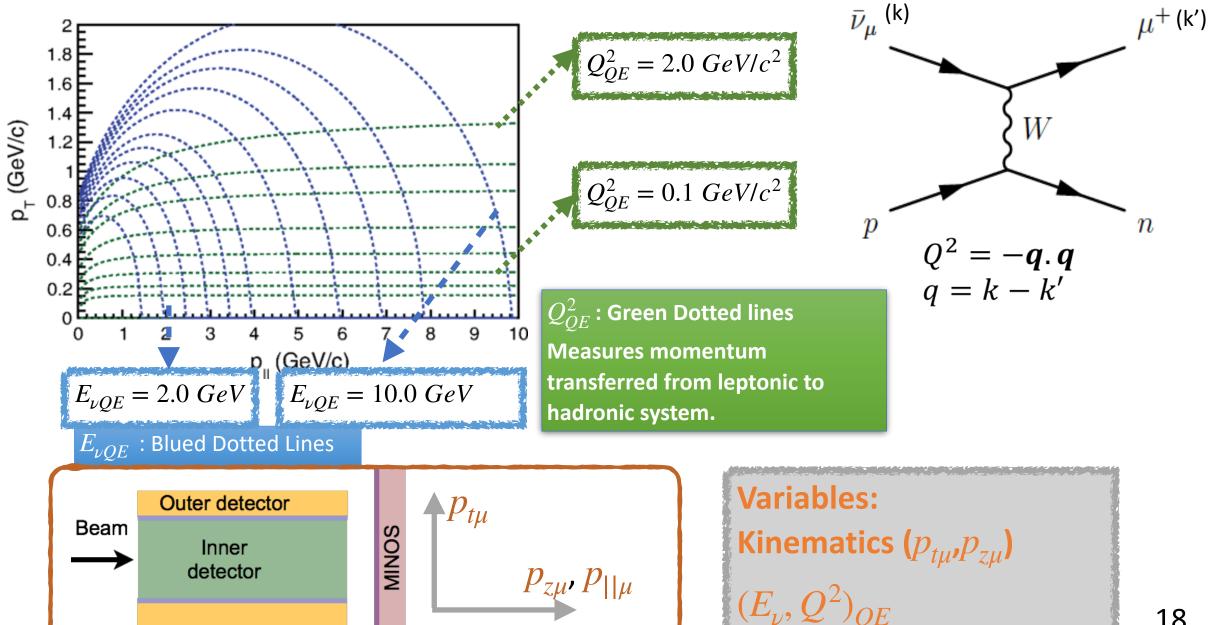
 $E_{\mu} \rightarrow Energy \ of \ the \ outgoing \ muon$

 $p_{\mu} \rightarrow Momentum \ of \ the \ outgoing \ muon$

 $Q^2 \rightarrow$ Four momentum transferred squared



Cross section Measurement Variables



$\bar{\nu}_{\mu}$ CCQELike (CH Target) Measurement in MINERvA

Measurement of Muon Antineutrino Quasielastic Scattering on a Hydrocarbon Target at Ev~3.5 GeV

MINERvA Collaboration • L. Fields (Northwestern U.) et al.

LE Beam

 $\overline{dQ^2}_{OR}$

Indication of multi-nuclear effects in CCQE processes

Measurement of the Muon Antineutrino Double-Differential Cross Section for Quasielastic-like

Scattering on Hydrocarbon at \$E \nu \sim 3.5\$GeV

Published in: Phys.Rev.Lett. 111 (2013) 2, 022501

LE Beam

MINERvA Collaboration • C.E. Patrick (Northwestern U.) et al.

Incorporation of 2p2h (multi nuclear) process

Published in: Phys.Rev.D 97 (**2018**) 5, 052002

High-Statistics Measurement of Antineutrino Quasielastic-like scattering at

<u>\$E \nu \sim\$ 6~GeV on a Hydrocarbon Target</u>

ME Beam

Higher Statistics, improved background constrain and measurement in previously unexplored region

MINERvA Collaboration • <u>A. Bashyal</u> (<u>Argonne</u> and <u>Oregon State U.</u>) et $a dp_z dp_t$, $(dE_\nu dQ^2)_{OE}$

e-Print: 2211.10402 [hep-ex] Submitted to the PRD (This Talk)

3 D $\bar{\nu}_u$ CCQELike Cross section Measurement

Ongoing Work

ME Beam

 $d^3\sigma$ $dRdp_z dp_t$

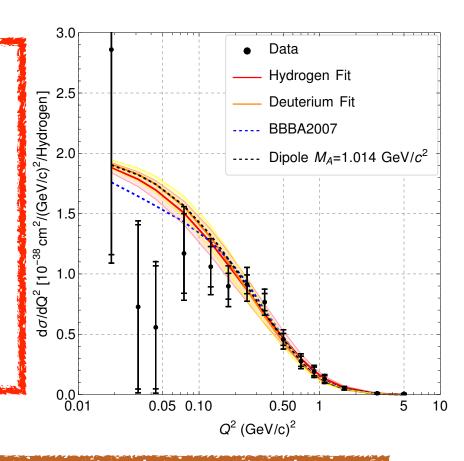
Measurement of visible energy (needed to measure E_{ν}^{reco}) in different μ^+ kinematic bins

Measurement of the axial vector form factor from antineutrino-proton scattering in $MINER\nu A$

Tejin Cai

University of Rochester York University

Joint Experimental-Theoretical Physics Seminar, Feb 1, 2023



Article | Open Access | Published: 01 February 2023

Measurement of the axial vector form factor from antineutrino-proton scattering

T. Cai , M. L. Moore, A. Olivier, S. Akhter, Z. Ahmad Dar, V. Ansari, M. V. Ascencio, A. Bashyal, A. Bercellie, M. Betancourt, A. Bodek, J. L. Bonilla, A. Bravar, H. Budd, G. Caceres, M. F. Carneiro, G. A. Díaz, H. da Motta, J. Felix, L. Fields, A. Filkins, R. Fine, A. M. Gago, H. Gallagher, ... L. Zazueta + Show authors

Nature 614, 48–53 (2023) | Cite this article 6757 Accesses | 158 Altmetric | Metrics

Signal of above analysis is subset of this analysis' Signal.

Simulation of Neutrino Nucleon Interaction in MINERvA

GENIE 2.12.6 [arXiv:1510.05494	[hep-ph]]is our neutrino	MC generator:
---------------------------------------	--------------------------	---------------

- Nuclear Model (initial state) → Relativistic Fermi Gas (RFG) Model for initial nuclear state with an additional high energy tail as prescribed by Bodek and Ritchie [A. Bodek and J. Ritchie, Phys. Rev. D 23, 1070 (1981)].
- ☐ Final State Interaction of hadrons →INTRANUKE h-A model
- ☐ QE Process → Llewellyn-Smith formalism + BBA05 with M_A = 0.99 GeV C.H. Llewellyn Smith, Phys. Rept. 3C, 261 (1972)
- \blacksquare Resonance π production \rightarrow Rein Sehgal model with $M_A = 1.12$ GeV
- Deep Inelastic Scattering → Bodek- Yang Model [A. Bodek,arXiv:hep-ph/0411202 [hep-ph].]

Full Detector Simulation (GEANT4) to simulate the response of the detector for particles that interact with the detector.

Simulated Neutrino Flux WE COVERED THIS

GENIE

Tunes

GENIE+Tunes → **MINERvA** Tune

Referred as MINERvA Tune v1 in this Talk

Various corrections to the GENIE generated events.

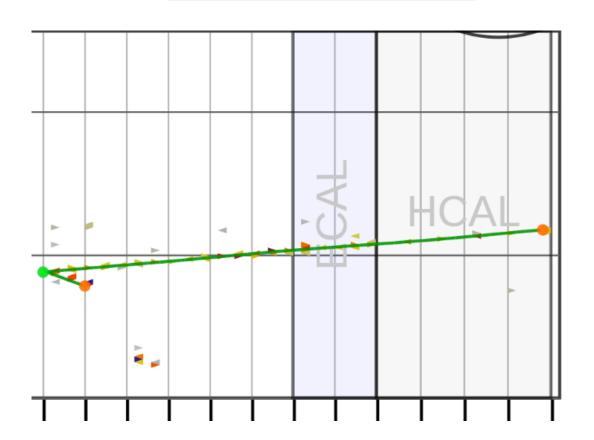
- ☐ QE events at Low Q² modified using Valencia RPA model
- Multi nuclear effects (2p2h) added based on Nieves Valencia model [Phys. Rev. D 88, 113007 (2013), arXiv:1307.8105 [hep-ph].
- 2p2h processes are enhanced based on a fit to MINERvA Low energy Inclusive data [P. Rodrigues et al. (MINERvA), Phys. Rev. Lett. 116, 071802 (2016)]
- Non resonant pion production suppressed by 40% based on re-analysis of bubble chamber data [P. Rodrigues, EPhys. J. C 76, 474 (2016), arXiv:1601.01888 [hep-ex]].

MINERVA Detector Simulation

GENIE Generated Neutrino
Interactions



Full Detector Simulation (GEANT4) to simulate the response of the detector for particles that interact with the detector.



- Data Overlay is done on the detector simulated events.
- Simulate pileup effect from other interactions (upstream and in the target) in the simulated interactions.

CCQE Cross-section

 CCQE Cross-section is generally expressed in Llewellyn Smith formalism:

$$\frac{d\sigma}{dQ^2} \begin{pmatrix} \nu n \to l^- p \\ \bar{\nu}p \to l^+ n \end{pmatrix} = \frac{M^2 G^2 \cos^2 \theta_c}{8\pi E_{\nu}^2} [A(Q^2) \mp B(Q^2)] \frac{(s-u)}{M^2} + \frac{C(Q^2)(s-u)^2}{M^4}]$$
(1)

$$M = \frac{m_p + m_n}{2}$$

$$s - u = 4ME - Q^2 - m_l^2$$

$$G_F^2 = 1.18 \times 10^{-5} GeV^{-2}$$
 (Fermi Coupling constant)
$$\theta_C = 0.974$$

CCQE Cross-section

$$\begin{split} A(Q^2) &= 4\frac{Q^2}{4M^2} |F_A|^2 (1 + \frac{Q^2}{4M^2}) - |F_V^1|^2 (1 - \frac{Q^2}{4M^2}) + \\ & |\xi F_V^2|^2 \frac{Q^2}{4M^2} (1 - \frac{Q^2}{4M^2}) + 4F_V^1 \xi F_V^2 \frac{Q^2}{4M^2}] \end{split} \tag{1}$$
 Valid for ν_μ, ν_e

$$B(Q^2) = 4\frac{Q^2}{4M^2} [F_A(F_V^1 + \xi F_V^2)] \tag{2}$$

$$C(Q^2) = \frac{1}{4} |F_A|^2 + |F_V^1|^2 + \frac{Q^2}{4M^2} |\xi F_V^2|^2$$
 (3)

$F_A \rightarrow$ Axial Form Factor

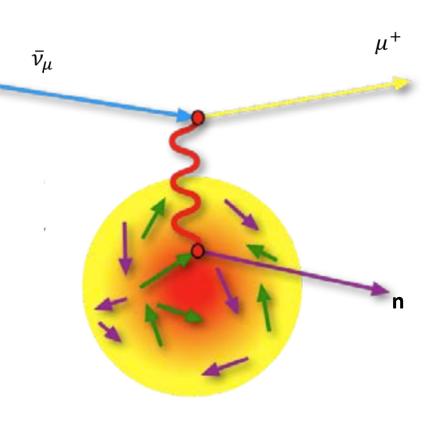
$$F_A(Q^2) = -\frac{g_A}{(1 + \frac{Q^2}{M_A^2})^2}$$



- Based on bubble chamber (hydrogen targets) measurements
- Measurements in heavier target report slightly higher axial mass
 - Nuclear Effects
- ullet Dipole Form Factor approximation breaks at high Q^2

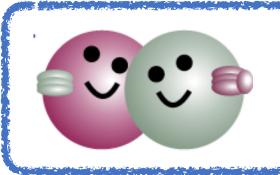
Neutrino Nucleon Cross-section Modeling

 CCQE cross-section model based on the Llewellyn Smith formalism.



- Cross-section models are based on neutrino interaction with the free nucleons.
- In heavy targets like carbon (this analysis), argon, etc., the final state particle (that exits the nucleus) can be changed due to nuclear effects that are not modeled by the Fermi Gas Model.

Nuclear Effects: nucleon-nucleon correlation

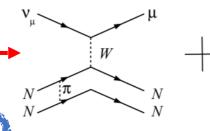


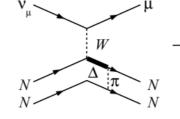
Short Range correlations modeled by Bodek Ritchie Tail added in RFG and Spectral Functions





modeled by Meson Exchange current and Transverse Enhancement model





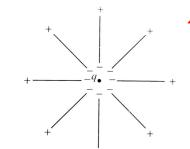




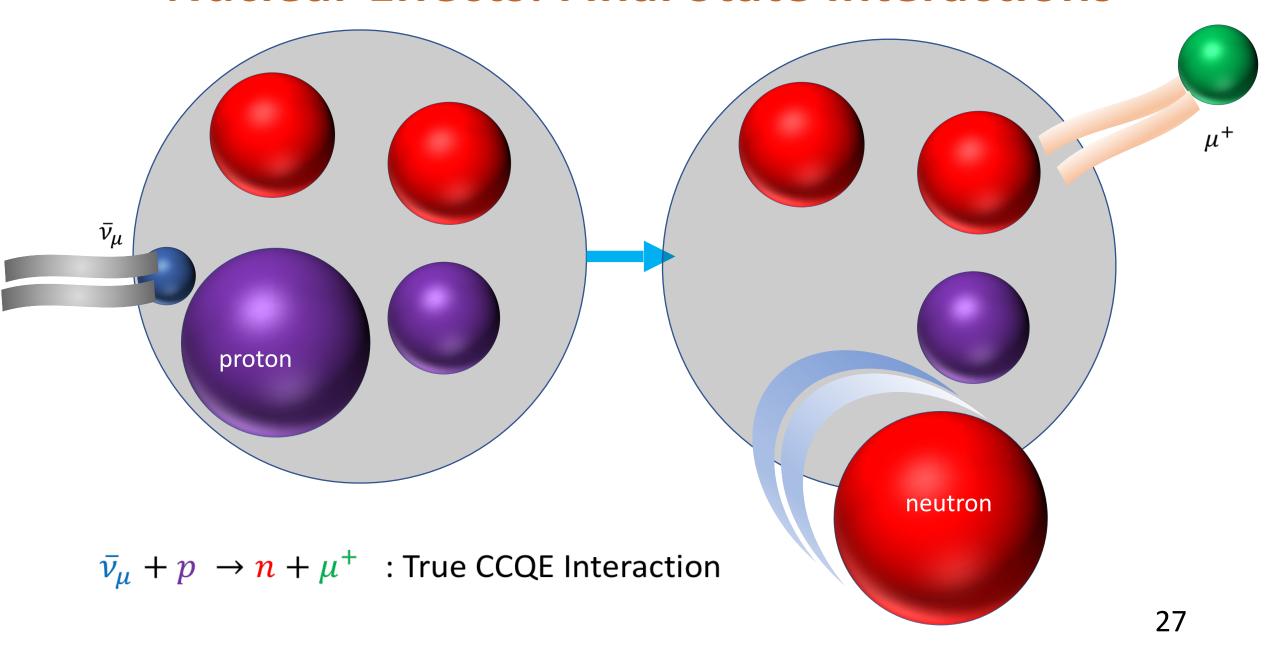
Long Range Correlations
Random Phase

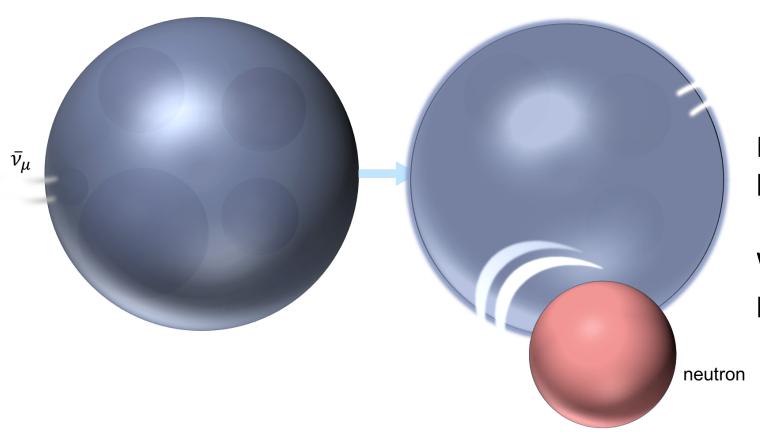
Approximation

Valencia 2p2h model for MEC effects



RPA: Like dielectric effect in EM interaction. Polarization of nucleus screens the coupling of W boson. Suppression of cross-section at low Q^2 .



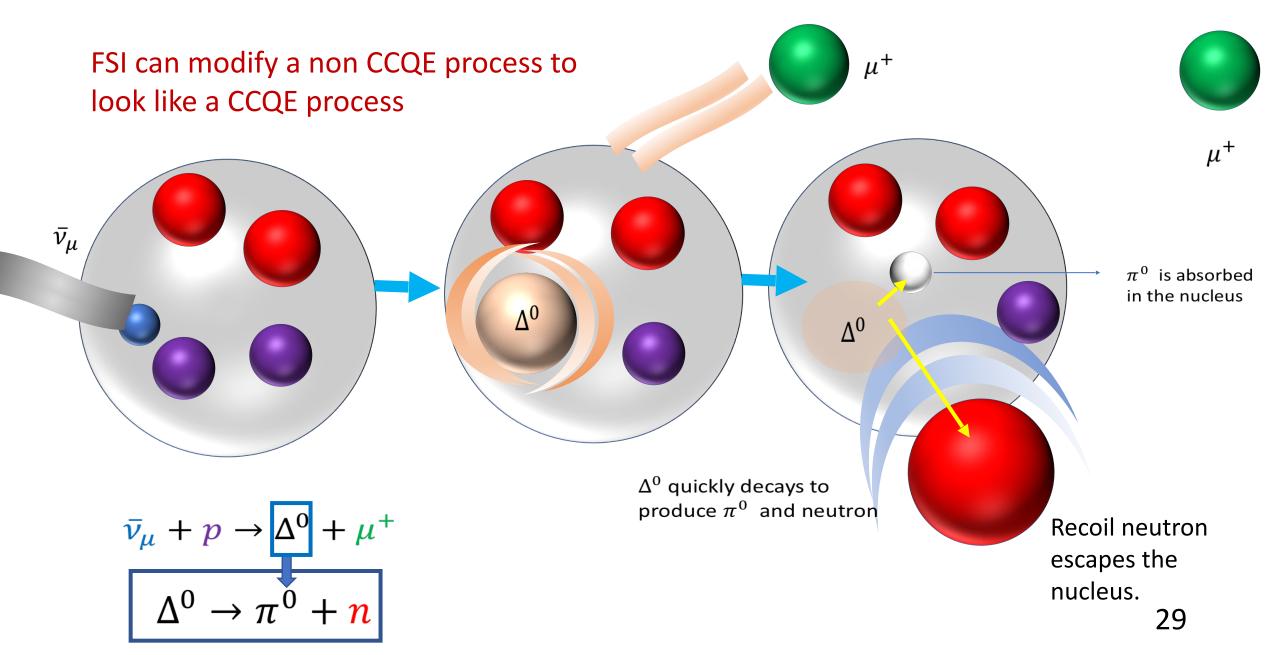


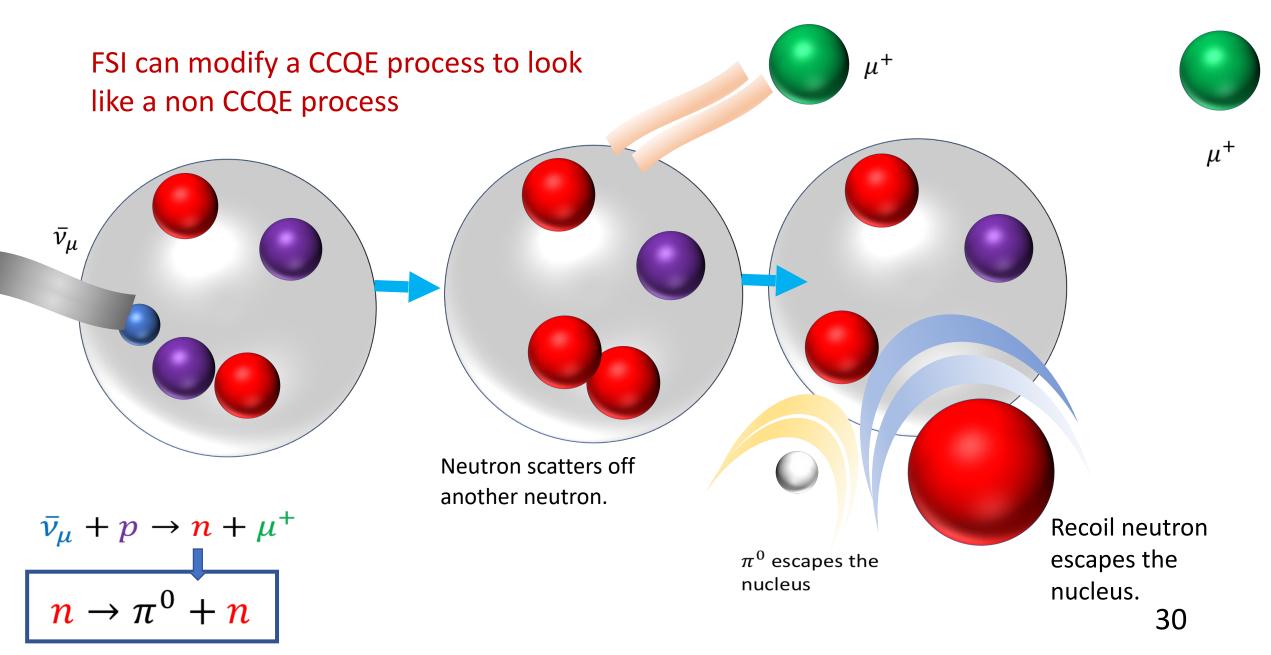


But we don't know what happens inside the nucleus.

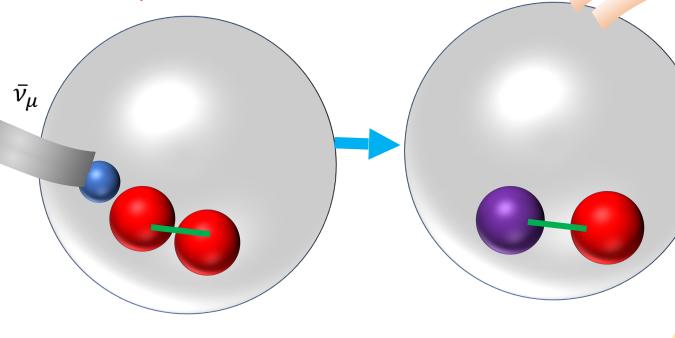
We can only see the final state particles in the detector.

$$\bar{\nu}_{\mu} + p \rightarrow n + \mu^{+}$$
: True CCQE Interaction



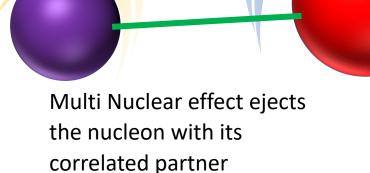


Multi-nuclear effects like 2p2h (2 particles 2 holes) makes the picture more complicated.



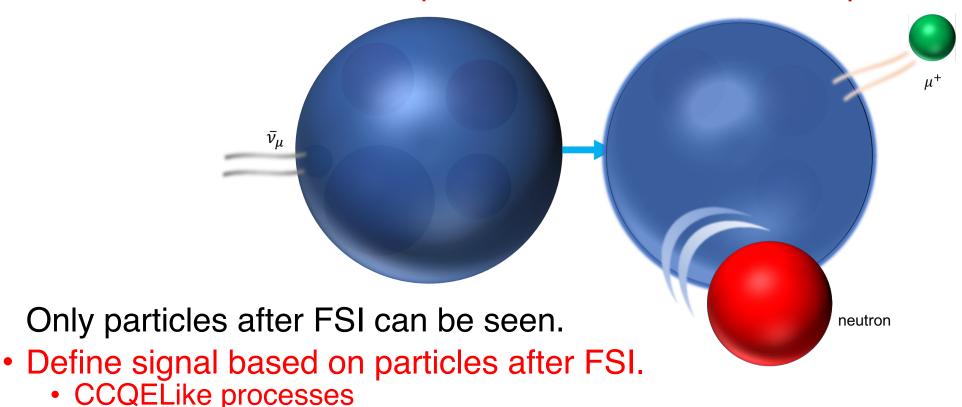
Few other things to consider:

- RPA and other multi-nuclear effects
- Nucleons are not at rest

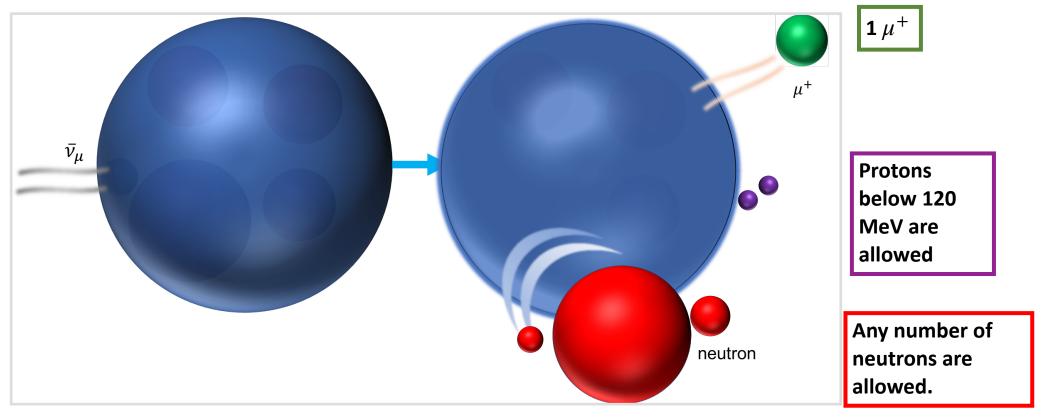


CCQE-Like Event

- FSI can fake a CCQE process to look like a non CCQE process.
- FSI can fake a non CCQE process to look like a CCQE process.



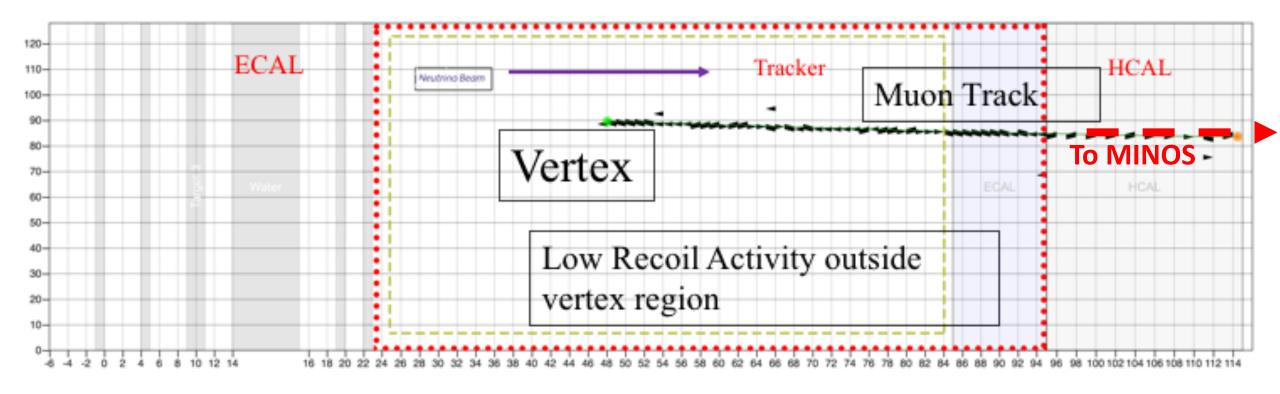
Signal Definition: CCQELike Process



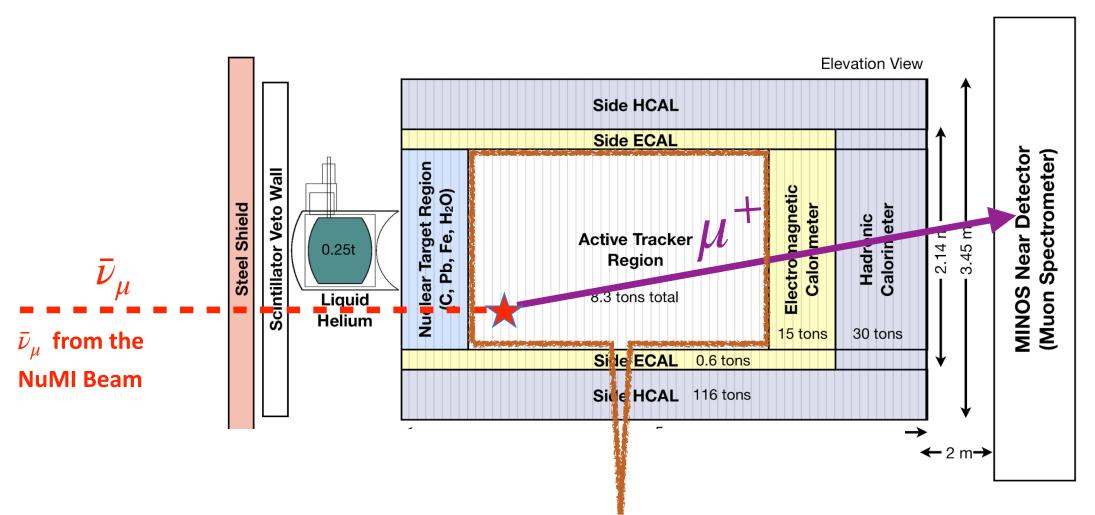
Signal Definition based on Final State Particles

- 1 positive muon (μ^+)
- Any number of neutrons
- Any number of Protons below 120 MeV Kinetic Energy
- No mesons (particles like π^{\pm} which are produced in Resonance processes)

Event Reconstruction: CCQELike Process

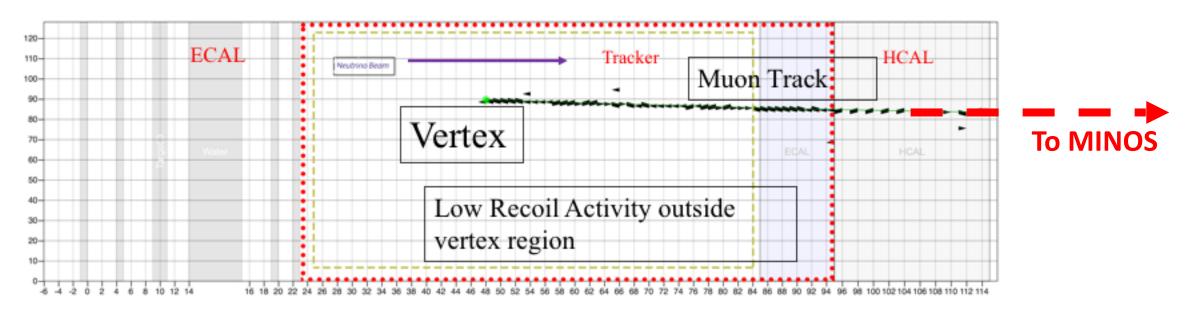


Event should be in the **Tracker (CH)** region of the detector. We are interested in the cross-section in CH target.



Tracker Region of of the MINERvA Detector

Event Reconstruction: CCQELike Process

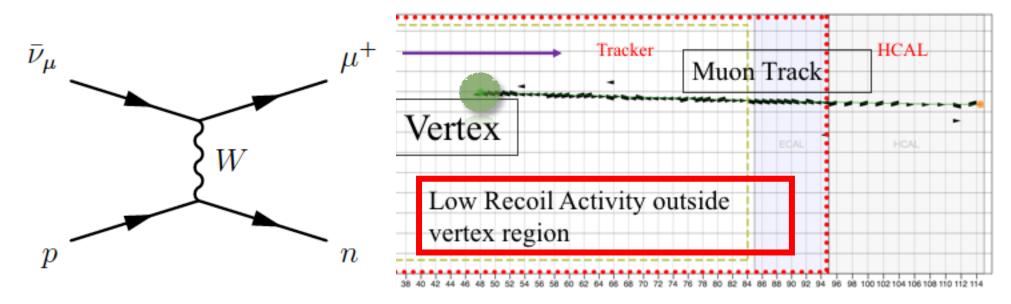


Signal Definition

Event should have 1 positive muon in the final state.

- 1 MINOS matched muon track (to identify the charge of the muon)
 - Apply a 20^o angle cut on muon track (with respect to the $\bar{\nu}_u$ beam)
- No Additional tracks (in next few slides)

Event Reconstruction: CCQELike Process



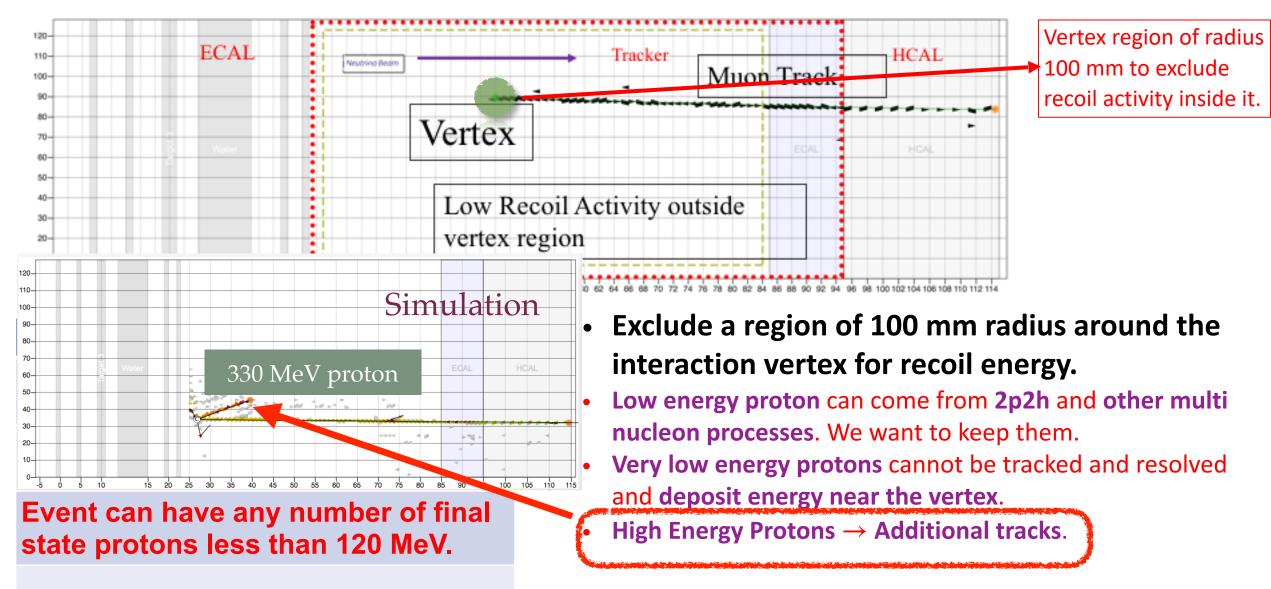
Signal Definition

Event should have 1 positive muon in the final state.

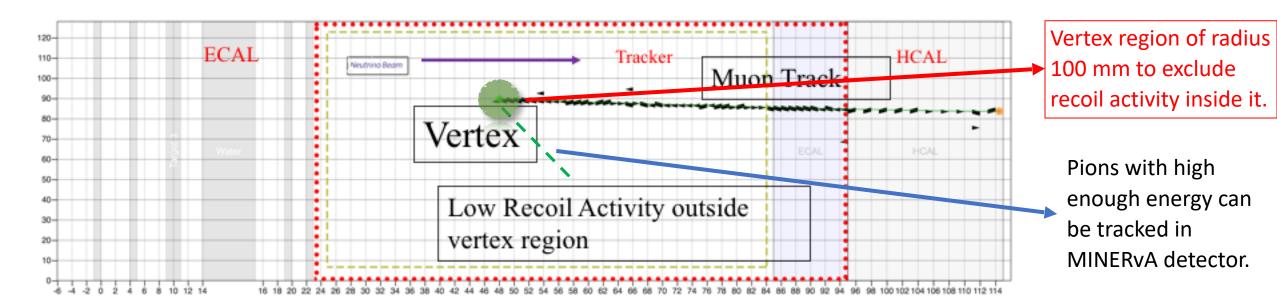
Event can have any number of neutrons in the final state.

- Low Recoil Activity outside the vertex region
- Most of the energy from the interaction carried away by the muon
- Recoil Activity → Isolated clusters outside the vertex region
- High recoil activity events are dominated by resonance and Deep Inelastic events.

Event Reconstruction: CCQELike Process



Event Reconstruction: CCQELike Process



Signal Definition

Event should have 1 positive muon in the final state.

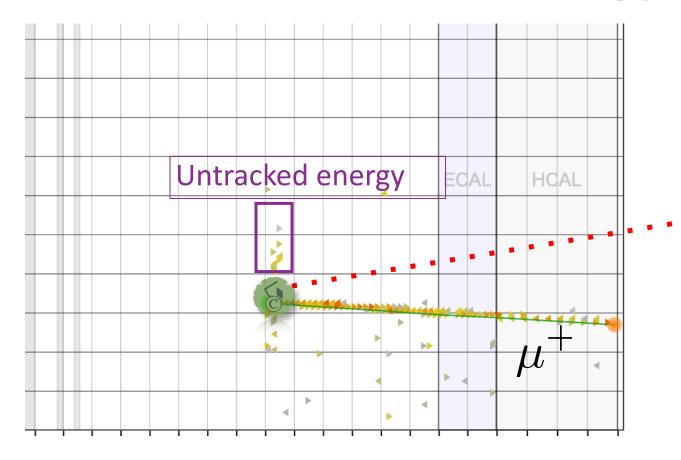
Event can have any number of neutrons in the final state.

Event can have any number of final state protons less than 120 MeV.

Event cannot have any mesons in the final state.

- Only 1 track events (track being muon track) are selected.
- Additional tracks → charged pions are rejected.
 - Remember protons with high Kinetic Energy also make tracks [and (>120 MeV) rejected]

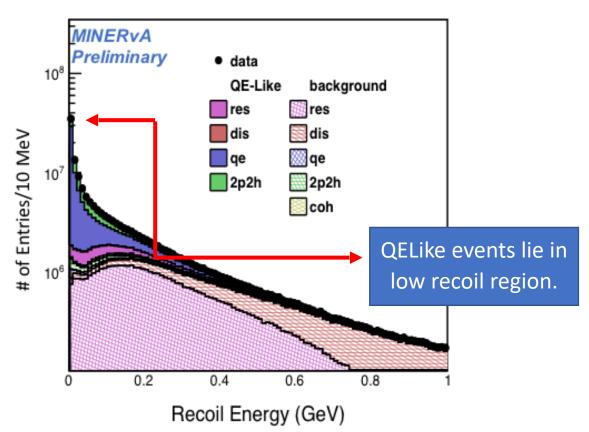
Recoil Energy Definition

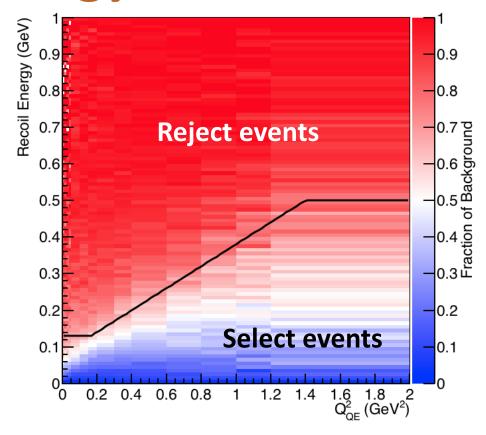


All the untracked energy outside the spherical blob of radius 100 mm outside the interaction vertex.

Untracked energy is the energy of the clusters that are not part of the μ^+ track.

Recoil Energy





- Distribution of Recoil Energy with Data and Various MC components.
- Recoil Energy cut based on the previous iteration of this analysis [Phys.Rev.D 97 (2018) 5, 052002]
- Optimized for signal selection efficiency+purity of selected sample.
- Loose cut at high Q_{OE}^2 region
 - Keep 2p2h events in this region that is not well-understood.

Extraction of Cross-section

$$\left(\frac{d^2\sigma}{dxdy}\right)_{ij} = \frac{\sum_{\alpha\beta ij} U_{\alpha\beta ij} (N_{data,\alpha\beta} - N_{data,\alpha\beta}^{bkg})}{\epsilon_{ij} \Phi T \Delta x_i \Delta y_j}$$

- This analysis measures cross section (σ) as a function of 2 variables x and y.
- Event reconstruction in (α, β) bins.
- Want to measure in true bins (i, j) bins.
- More on this later.

$$\begin{array}{c} \bullet \ (\frac{d^2\sigma}{dp_zdp_t}) \to \text{Cross section as a function of } \frac{\text{muon kinematics}}{E_{\nu}^{QE}} = \frac{m_n^2 - (m_p - E_b)^2 - m_{\mu}^2 + 2(m_p - E_b)E_{\mu}}{2(m_p - E_b - E_{\mu} + p_{\mu}\cos\theta_{\mu})} \\ \bullet \ (\frac{d^2\sigma}{dE_{\nu}dQ^2})_{QE} \to \text{Cross section as a function of } \frac{E_{\nu}}{e} \text{ and four} \end{array}$$

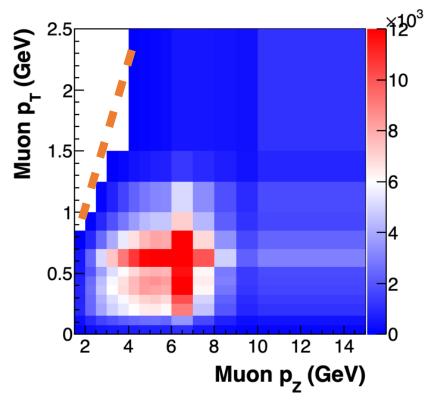
momentum transferred (Q^2) based on QE hypothesis

Event Reconstruction

$$\left(\frac{d^2\sigma}{dxdy}\right)_{ij} = \frac{\sum_{\alpha\beta ij} U_{\alpha\beta ij} \left(N_{data,\alpha\beta} - N_{data,\alpha\beta}^{bkg}\right)}{\epsilon_{ij} \Phi T \Delta x_i \Delta y_j}$$

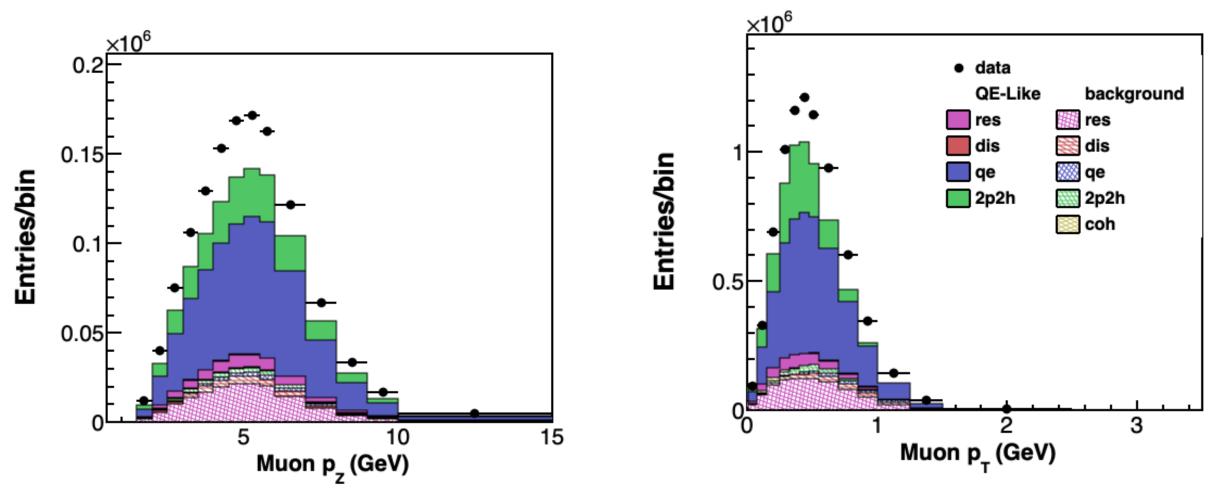
Raw Event Reconstruction

$$\left(\frac{d^2\sigma}{dxdy}\right)_{ij} = \frac{\sum_{\alpha\beta ij} U_{\alpha\beta ij} \left(N_{data,\alpha\beta} - N_{data,\alpha\beta}^{bkg}\right)}{\epsilon_{ij} \Phi T \Delta x_i \Delta y_j}$$



- Apply the selection conditions on the data and simulated sample to select the CCQE Like candidate events.
- Requiring muon track angle to be less than 20 degrees rejects events at high p_T and low p_Z phase space.

Raw Event Selection

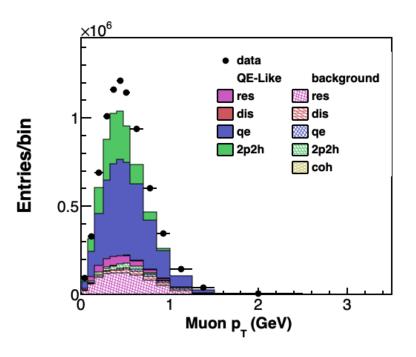


We look at our simulated sample to estimate the types of events we select in our data.

*1 D are the projections from 2 D distributions.

Muon p_z || Neutrino Beam

Selected Events (Signal)

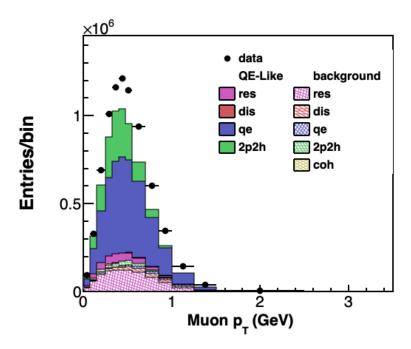


Signal Components	Fraction of Total MC Events
QE	0.54
RES	0.05
DIS	0.003
2p2h	0.193
Total	0.786

Signal (QE-Like)

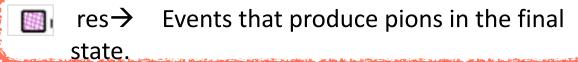
- res Events that produced pions initially but go through FSI to produce neutron in the final state.
- dis→ Events where the neutrino completely breaks the nucleon and creates high recoil activity but only neutron is the Final State Particle (very rare)
- qe → True QE events. Recoil neutrons do not go through FSI and escape the nucleus.
- 2p2h → Events in which the neutrino interacts with a correlated pair of nucleons and both nucleons exit the nucleus.

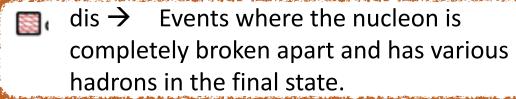
Selected Events (Background)



Background Components	Fraction of Total MC Events
RES	0.151
DIS	0.029
QE	0.012
2p2h	0.016
СОН	0.006
Total	0.214

Background





qe→ QE events where the neutron goes through FSI and produce other particles (like pions, or protons that can be tracked).

2p2h → correlated pair of nucleons are produced initially but go through FSI to produce mesons.

coh (coherent) → pion is produced through coherent process (initial state of the nucleon is not modified).

$$\bar{\nu}_{\mu} + A \rightarrow \mu^{+} + \pi^{-} + A$$

A is the target nucleus.

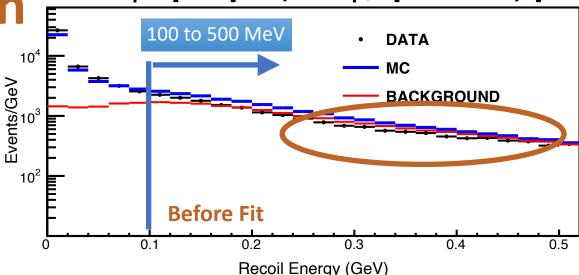
Background

$$\left(\frac{d^2\sigma}{dxdy}\right)_{ij} = \frac{\sum_{\alpha\beta ij} U_{\alpha\beta ij} (N_{data,\alpha\beta} - N_{data,\alpha\beta}^{bkg})}{\epsilon_{ij} \Phi T \Delta x_i \Delta y_j}$$

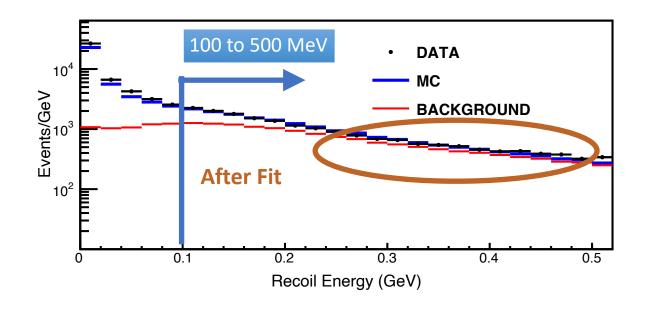
- Our selected data has both signal and background contribution.
- We need to subtract the background from our selected sample.
- Cannot rely on MC completely to estimate the background in the data.
 - Pion production may be over predicted by our simulation.
 - Nuclear effects have significant uncertainties.
 - We want to improve simulation models from our data after all.
- Use Data driven method to subtract the background from the data.

Background Subtraction

- We look at our recoil energy distribution in 14 different bins of muon p_T and p_Z .
- Fraction fit of data with simulated signal and background recoil energy distribution shapes.
 - ROOT::TFractionFitter [root.cern.ch]
 - Fit done between 100 to 500 MeV recoil energy region.
 - Background rich region
- Fit gives the best estimation of signal and background fraction in our data.
 - Efficiency correction to get the signal fraction in each fit region.

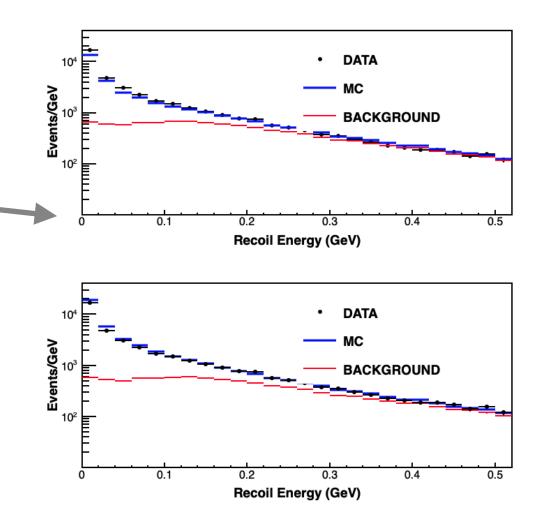


 $pz = [1.5-5] GeV/c and p_t = [0.2-0.4 GeV/c]$



Muon P_T(GeV/c) 1.5⊢ 0.5 15 10 Muon P_z(GeV/c) DATA Events/GeV Recoil Energy (GeV) DATA Events/GeV **BACKGROUND** Recoil Energy (GeV)

Background Subtraction



Fit Done in all 14 bins to extract the signal fraction in the data.

$635,592\pm1,251$ (stat.) $\pm 13,850$ (syst.) events after background subtraction.

$p_T \mathrm{GeV/c}$	$p_z \mathrm{Gev/c}$	Signal Before	Signal After	Signal Efficiency	Efficiency Corrected Signal	χ^2/NDF
0.0 - 0.2	1.5 - 5.0	0.56	0.610 ± 0.027	0.901	0.78	18.12/19
0.2 - 0.4	1.5 - 5.0	0.62	0.700 ± 0.011	0.905	0.86	26.82/19
0.4 - 0.65	1.5 - 5.0	0.61	0.673 ± 0.008	0.901	0.84	16.08/19
0.65 - 0.82	1.5 - 5.0	0.59	0.622 ± 0.012	0.922	0.75	5.96/19
0.82 - 1.0	1.5 - 5.0	0.59	0.582 ± 0.0188	0.951	0.66	14.74/19
1.0 - 2.5	1.5 - 5.0	0.0.59	0.597 ± 0.0373	0.988	0.61	16.44/19
0.0 - 0.2	5.0 - 8.0	0.62	0.788 ± 0.0367	0.923	0.89	11.43/19
0.2 - 0.4	5.0 - 8.0	0.67	0.772 ± 0.015	0.922	0.89	16.64/19
0.4 - 0.65	5.0 - 8.0	0.67	0.719 ± 0.010	0.918	0.85	26.45/19
0.65 - 0.82	5.0 - 8.0	0.65	0.700 ± 0.0119	0.930	0.80	21.69/19
0.82 - 1.0	5.0 - 8.0	0.64	0.638 ± 0.0166	0.952	0.71	11.33/19
1.0 - 2.5	5.0 - 8.0	0.62	0.615 ± 0.026	0.983	0.63	19.48/19
0.0 - 0.5	8.0 - 15.0	0.69	0.778 ± 0.026	0.926	0.89	18.86/19
0.0 - 0.5	8.0 - 15.0	0.66	0.69 ± 0.0156	0.950	0.77	11.87/19
				•		
		- 1				
		↓			\	

Signal fraction in the simulated sample (100 to 500 MeV recoil energy)

Signal fraction predicted by the fit (100 to 500 MeV recoil energy)

Fraction of signal in the data sample that passes the recoil cut

51

Unsmearing the reconstructed events

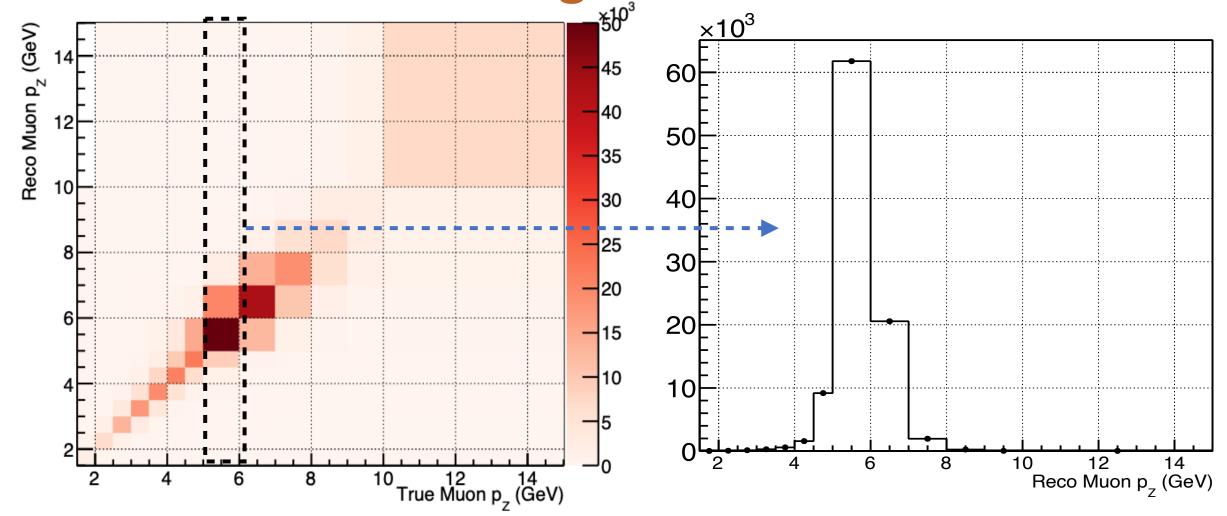
$$\left(\frac{d^2\sigma}{dxdy}\right)_{ij} = \frac{\sum_{\alpha\beta ij} U_{\alpha\beta ij} (N_{data,\alpha\beta} - N_{data,\alpha\beta}^{bkg})}{\epsilon_{ij} \Phi T \Delta x_i \Delta y_j}$$

- We look in our data as a function of reconstructed quantities $[(p_z,p_t)_\mu,(E_\nu,Q^2)_{QE}]$.
- Cannot measure these quantities perfectly
 - Limitation of our detector resolution
 - Reconstruction algorithms

Example

- A μ^+ whose actual (true) momentum is **5 GeV** could be reconstructed as 4 GeV sometimes or 3 GeV sometimes.
- Called the smearing of the events from true bin to the reconstructed bins.
- We want to test models against our measurements
 - Models are based on True Variables
- Need to correct our reconstructed events to their true phase space.
- $U_{\alpha\beta ij} \rightarrow$ Matrix that contain the smearing information of events in true bins ij to reconstructed bins $\alpha\beta$
- We use Iterative Bayesian Unfolding method [cite] to unfold our reconstructed data into the true bins.

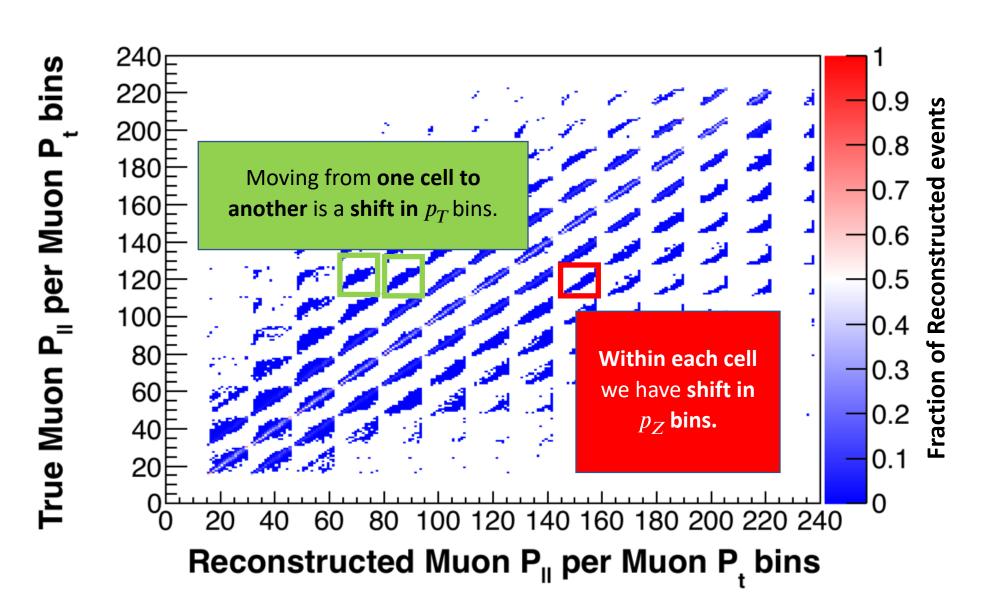
Smearing of Events

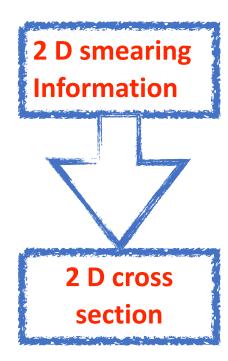


Smearing of events from their true p_z bins to reconstructed p_z bins.

Smearing of 5 to 6 GeV (true) events in different reconstructed bins

Unfolding Matrix





Unfolding of the Event

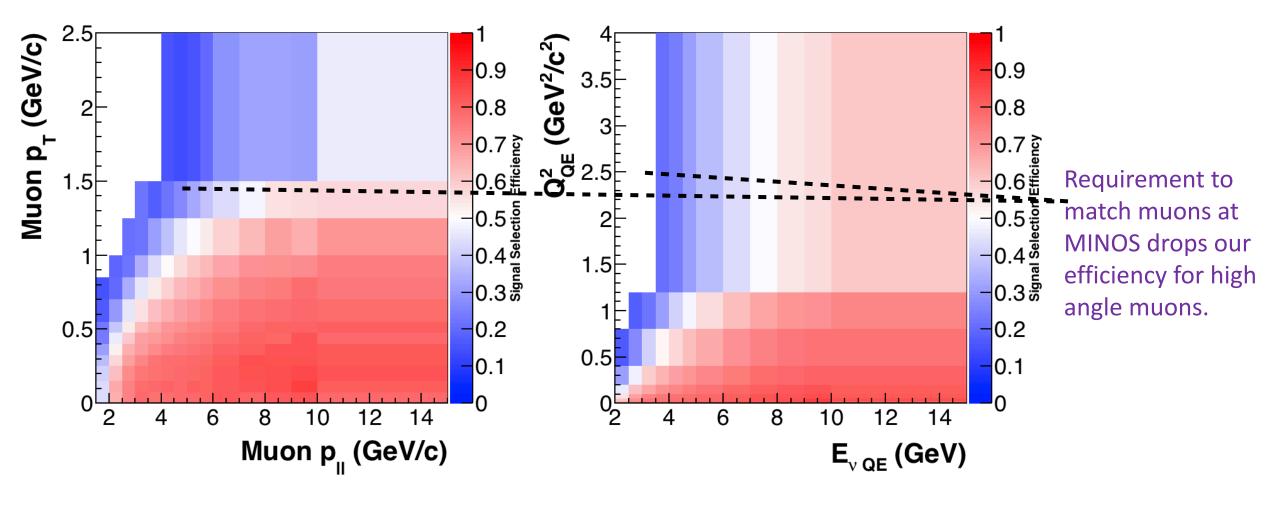
- Background subtracted sample is unfolded to the true kinematic variables.
 - Iterative Bayesian Unfolding [G. D' Agostini, Nuclear Instrument Method]
 - Based on RooUnfold Algorithm.
- Unfolding studies done with various model predictions.
 - Find Optimum number of iteration needed to unfold
 - Test the stability of the unfolding matrix
- $(p_z, p_t)_u = 4$ iterations
- $(E_{\nu},Q^2)_{QE}$ = 8 iterations

Signal Selection Efficiency Correction

$$\left(\frac{d^2\sigma}{dxdy}\right)_{ij} = \frac{\sum_{\alpha\beta ij} U_{\alpha\beta ij} (N_{data,\alpha\beta} - N_{data,\alpha\beta}^{bkg})}{\epsilon_{ij} \Phi T \Delta x_i \Delta y_j}$$

$$\epsilon = \frac{N_{QELIKE}^{Pass\ Cut}}{N_{QELIKE}^{Total\ Events}}$$

- We fail to reconstruct some signal events due to:
 - Detector acceptance
 - Remember MINOS acceptance requirement?
 - Reconstruction Efficiency
 - Our algorithms are not 100% perfect.
- Outer detector Correct for the fraction of events that we failed to reconstruct. Inner detector

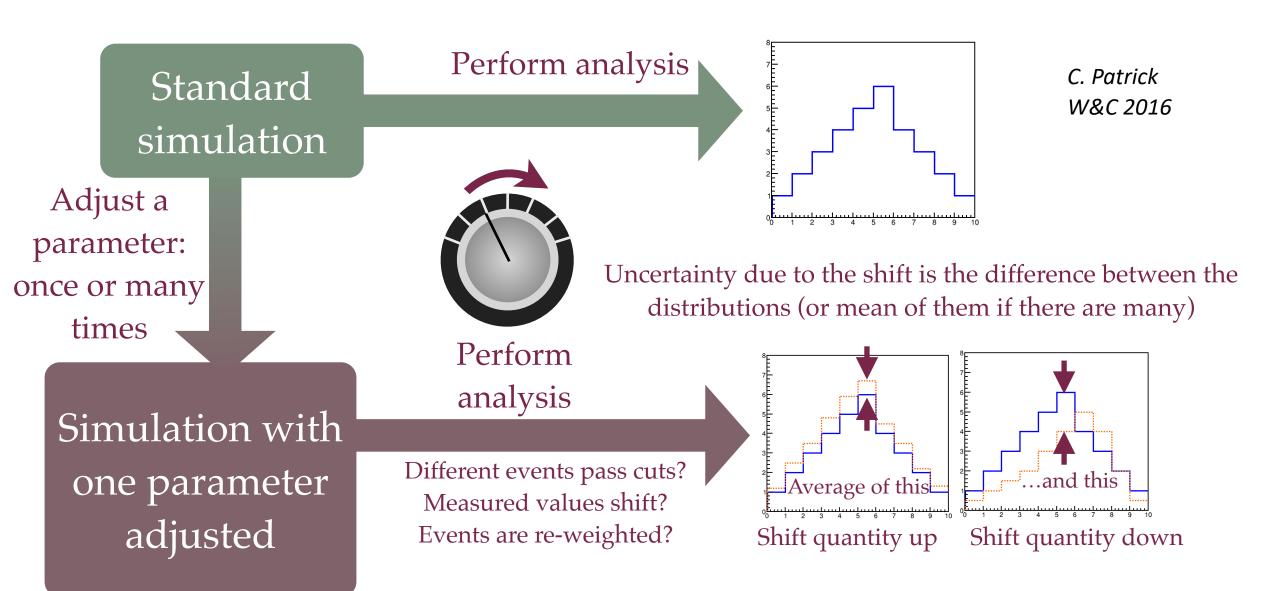


Flux And Target Normalization

$$\left(\frac{d^2\sigma}{dxdy}\right)_{ij} = \frac{\sum_{\alpha\beta ij} U_{\alpha\beta ij} (N_{data,\alpha\beta} - N_{data,\alpha\beta}^{bkg})}{\epsilon_{ij} \Phi T \Delta x_i \Delta y_j}$$

- Muon $P_z vs$. P_T : We use the integrated neutrino flux for normalization.
- $E_{\nu} \, vs$. Q^2 : distribution is normalized by neutrino flux of corresponding E_{ν} to get the neutrino cross-section independent of the shape of the flux.
- T \rightarrow Total number of nucleons in the tracker region.
- 3.23×10^{30} nucleons (protons+neutrons) in the tracker region.
- Φ (Flux) integrated from 0 to 120 GeV to obtain the differential cross section

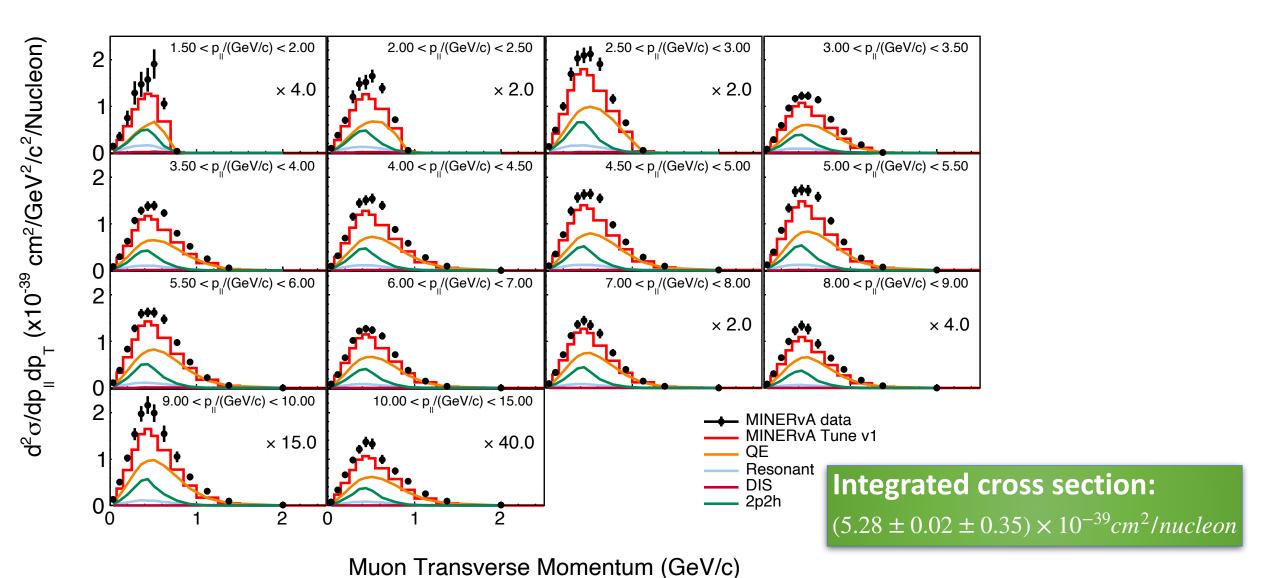
Systematic Uncertainties



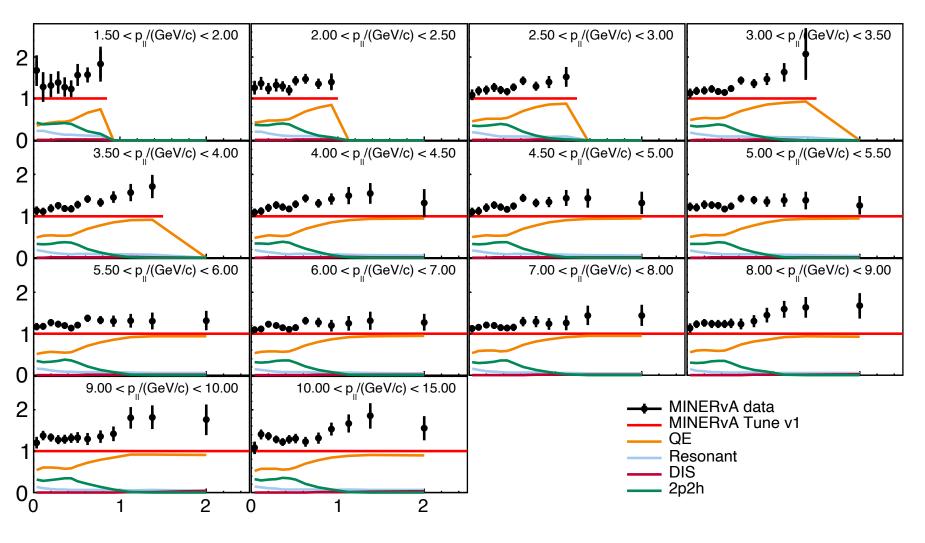
Examples: increase resonant cross section by 10% OR 500 "universes" of flux changes

Cross section Results

p_t cross section in the bins of p_z



Cross section Ratio (p_t in the bins of p_z)

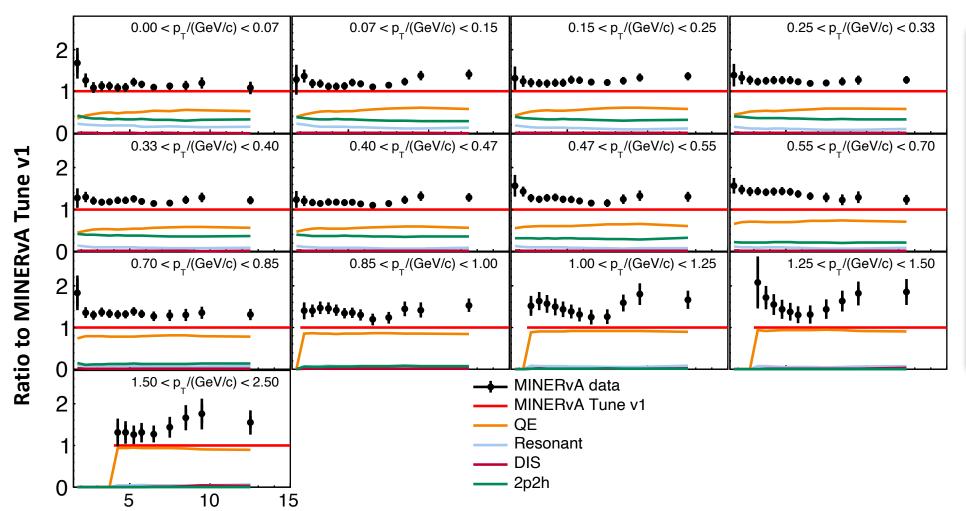


Across all bins the model under predicts our measured cross section.

CCQE Like Cross section dominated by QE processes followed by 2p2h.

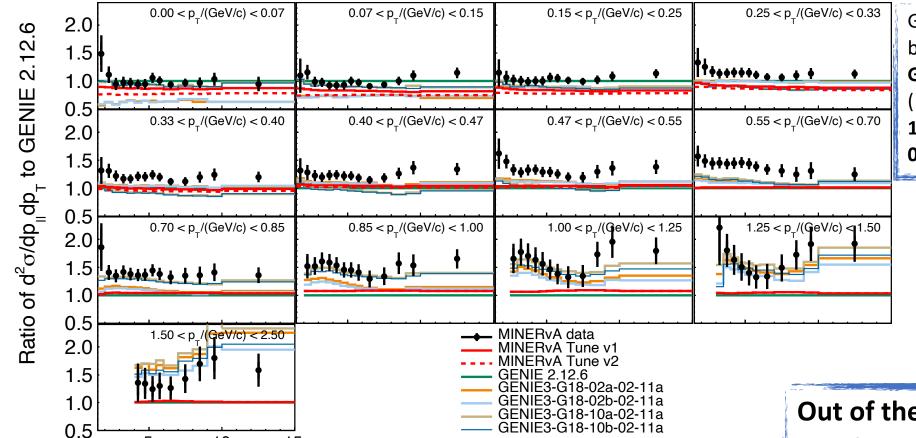
Muon Transverse Momentum (GeV/c)

Cross section ratio (p_z in the bins of p_t)



QE (1p1h) and 2p2h processes dominate the low transverse region whereas the higher transverse region is dominated by QE processes entirely.

Comparison with GENIE 3 models



GENIE 3 model **G18-10x-02-11a** agrees better with our data compared with **G18-02x-02-11a**

(x = a,b)

10x: incorporates Valencia model.

02x: GENIE ver2 2p2h model

GENIE 3 : GENIE 3.0.6

Muon Longitudinal Momentum (GeV/c)

MnvTune v2 = MnvTune v1+ Pion Production suppression Low Q^2

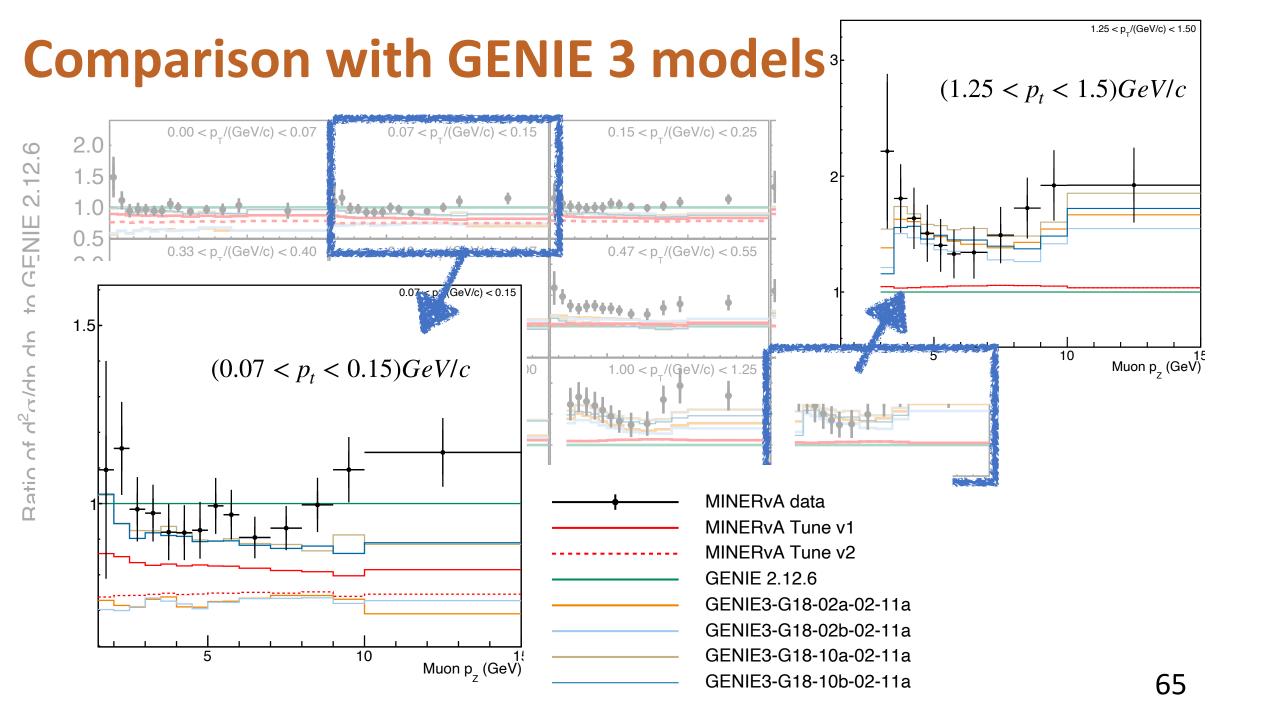
[*P. Stowell,* arXiv:1903.01558.]

5

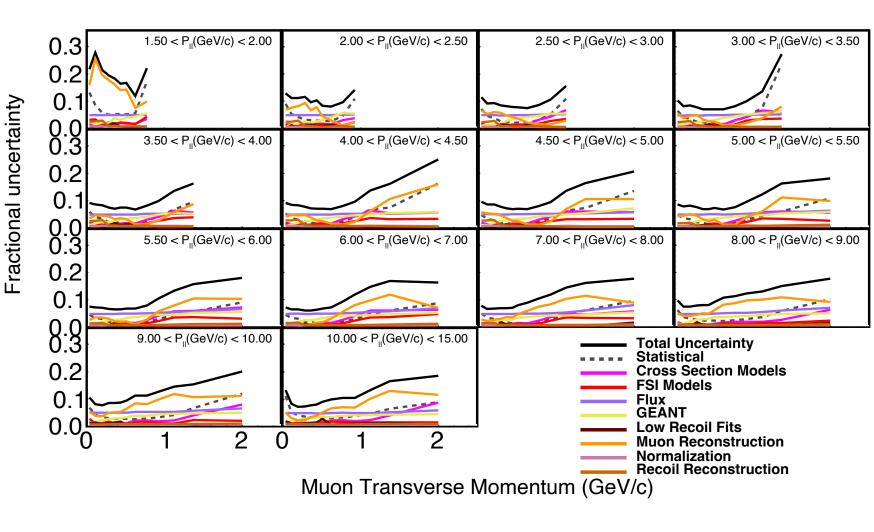
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15

Out of the box GENIE (2.12.6) does not describe our data well. We need more than our LE tunes to describe our ME data.



Error Summary on the data



- Uncertainty is dominated by the muon reconstruction.
- Flux contribution is around
 5%.
- Neutron related error dominates the GEANT category and contribute unto 5% in some regions.
- GENIE cross section models and FSI contribute contribute less than 5% overall.

χ^2 comparison

Model	χ^2 - linear	χ^2 - \log
GENIE 2.12.6 Tunes		
MINERvA Tune v1	362.6	580.4
MINERVA Tune v2	364.4	601.4
GENIE w/o 2p2h	226.5	473.2
GENIE (Default)	346.4	550.6
ĢENIE ∺ πtúne	354,3	568.5
GENIE+RPA	230.0	406.7
GENIE+RPA+ π tune	231.7	414.6
GENIE+Low Recoil Tune	755.4	1059.4
GENIE+Low Recoil Tune+RPA	361.2	570.0
GENIE+Low Recoil Tune+ π tune	760.6	1081.8
GENIE 3.0.6 Tunes		
GENIE 3.0.6 G18_02a_02_11a	602.9	865.0
GENIE 3.0.6 G18_02b_02_11a	586.9	878.3
GENIE 3.0.6 G18_10a_02_11a	353.1	447.5
GENIE 3.0.6 G18 ₋ 10b ₋ 02 ₋ 11a	312.8	421.7

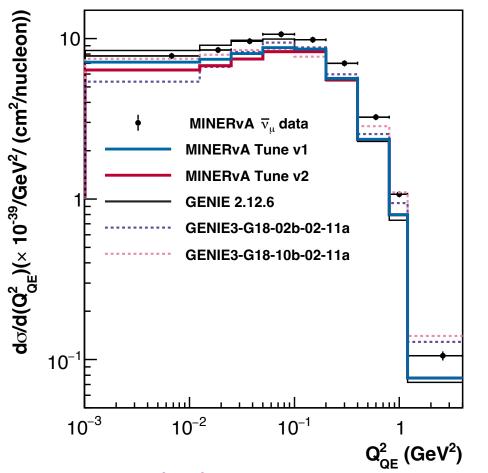
TABLE II. $p_{\parallel} - p_{\perp} \chi^2$ between data and model variants derived from GENIE. The number of degrees of freedom is 171. Both the χ^2 between the values and between the logs of the values are listed.

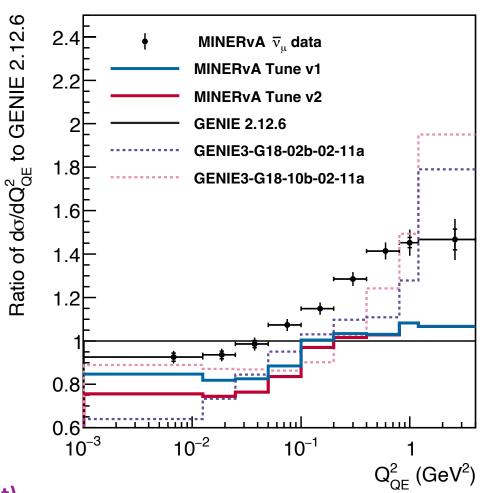
- Full systematic and bin to bin correlations are treated to calculate the χ^2 between data and models.
- χ^2 tells us that GENIE 2 variations that were based on low energy data are not sufficient to describe our Medium Energy data.
- GENIE 3 models with Nieves 2p2h
 implementation performs better than GENIE 2
 like 2p2h models.

1 D projections from $(E_{\nu}, \mathcal{Q}^2)_{QE} \text{ measurements}$

Cross section (Q_{OE}^2) : MINERvA Tune and GENIE 3

MINERvA Tune v2 = MnvTune v1+Pion Production suppression Low Q^2 based on P. Stowell, arXiv:1903.01558.

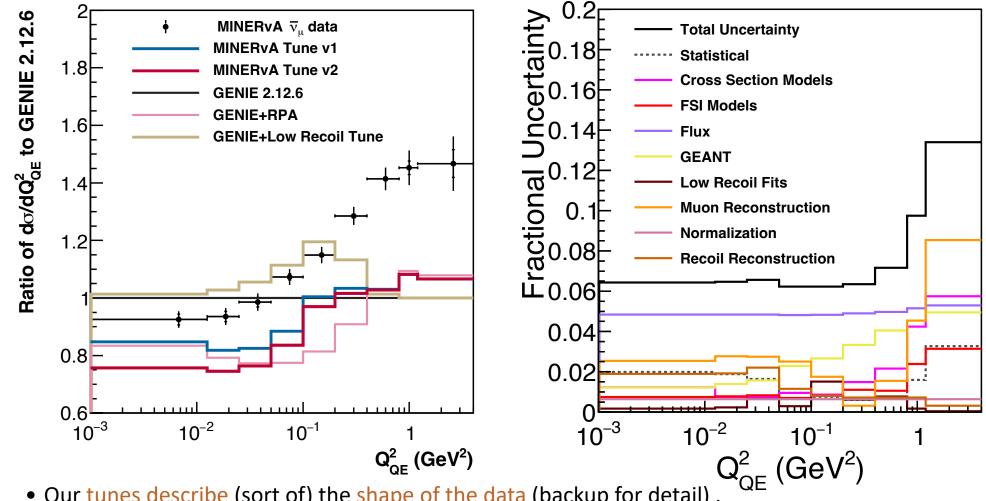




Cross section (left) and comparison of data and models (right).

GENIE 3 model with Nieves and hN (G18-10b series) describes data better at high Q_{QE}^2 than older models.

Cross section (Q $\binom{2}{OF}$: MINERVA Tunes

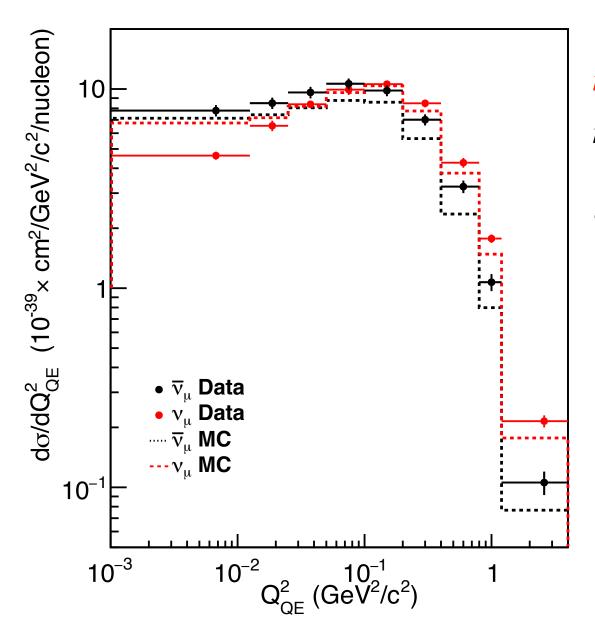


MnvTune v2 = MnvTune v1 **Pion Production**

suppression Low Q^2

- Our tunes describe (sort of) the shape of the data (backup for detail) .
 - But needs more than LE tunes to describe our ME data
- ullet Uncertainties on Q^2_{OE} cross section (right) are dominated by Flux, Muon reconstruction.
- GEANT4 uncertainties are dominated by neutron related interactions.

Comparison with MINERvA ME ν_{μ} CC0pi measurement



$$u_{\mu} + n \rightarrow \mu^{-} + p$$
 [M. Carniero, PRL 124,121801]
$$\bar{\nu}_{\mu} + p \rightarrow \mu^{+} + n$$
 [This Talk]

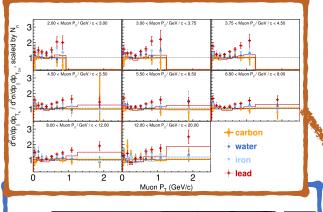
 Not one to one comparison but can show the agreement with data and our model for CCQELike cross section measurements for both neutrinos (black) and anti neutrinos (red)

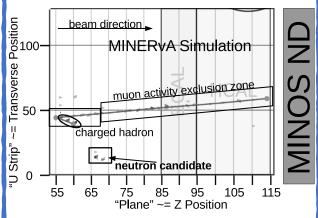
Conclusions

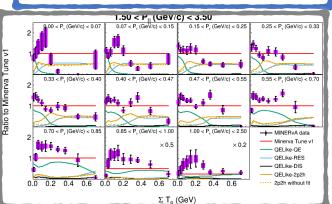
Conclusions

- This analysis provides high statistics cross section measurement of anti neutrinos CCQE like process.
- Extends measurements to previously unexplored kinematic regions.
- More than LE tunes are needed to describe our ME data.
- Models under predict the data
 - Similar to the LE era analysis
- Higher Statistics, better constrained flux systematics
 - Valuable information for upcoming oscillation experiments.

CCQELike LandScape • High Statistics results ν_{μ} [M. Carniero, PhysRevLett.124.121801]







and $\bar{\nu}_{\mu}$ CCQELike are published.

 $(rac{\sigma_{
u_{\mu}}}{\sigma_{ar{
u}_{\mu}}})_{CH}$ analysis with full treatment of systematics and

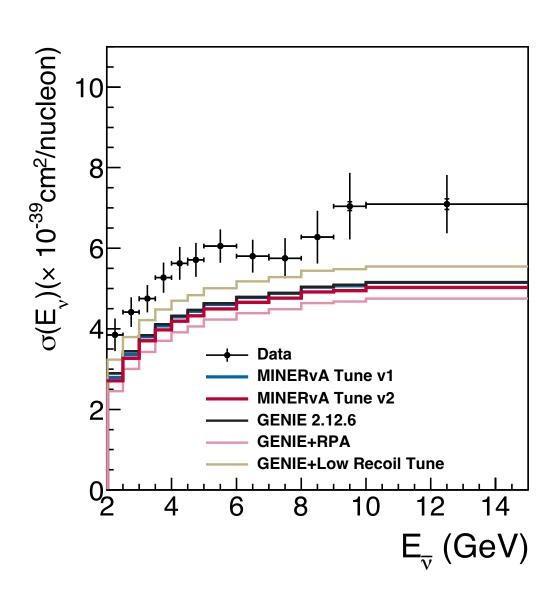
correlations is ongoing.

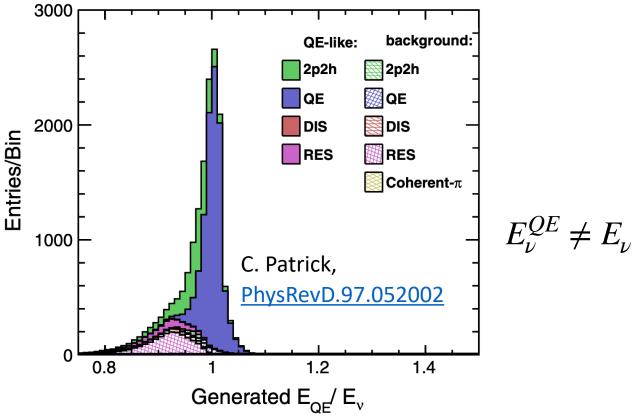
- $(\sigma_{\nu_{\mu}})_A, (\sigma_{\nu_{\mu}})_{CH}$ and ratios where **A** is C, H_20, Fe, Pb [J, Kleyklamp, arxiv: 2301.02272]
- Neutrons are the Final State particles of this analysis.
- Paper being prepared on cross section with 2 or more neutrons in final state.
- $\bullet \, \bar{\nu}_{\mu} \, \text{CCQELike}$ cross section measurement in heavier target (ongoing)
- \bullet $\bar{\nu}_{\mu}$ CCQELike cross section as a function of recoil energy in muon kinematic bins (ongoing)
 - \bullet ν_{μ} CCQELike version of this analysis published [*D*.

Ruterbories, arxiv 2203.08022]

Back up

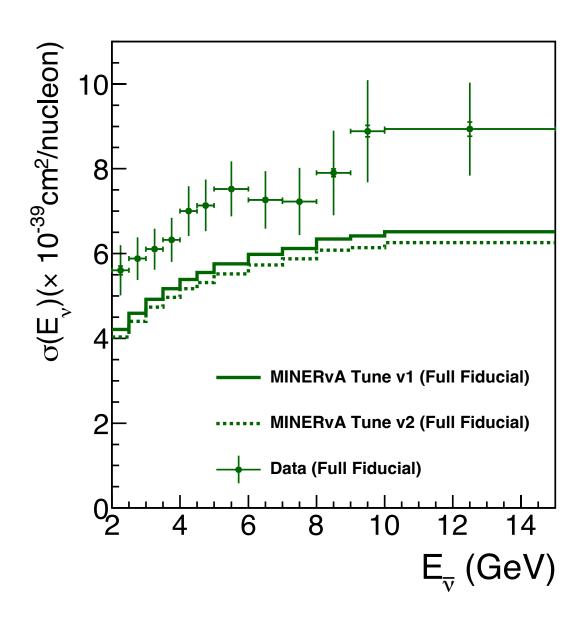
Cross section (E_{ν})





- Since the flux is estimated in the bins of true neutrino energy, $E_{\nu QE}$ cross section is corrected to E_{ν} (true neutrino energy).
- Correction introduces model dependency but allows (qualitative) comparison with other results and theoretical models.

Full Fiducial Cross section (E_{ν})



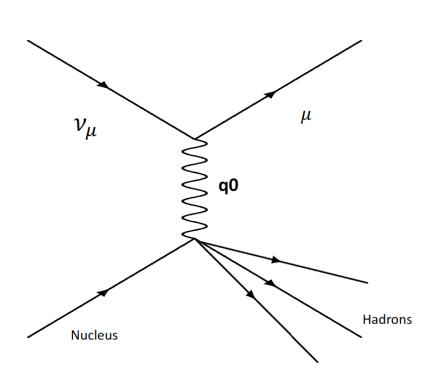
Signal Definition includes events with proton KE less than 120 MeV.

Remember that our selection criteria requires a 20% angle cut.

Removing the proton KE threshold and correcting for the angle cut gives the full fiducial cross section.

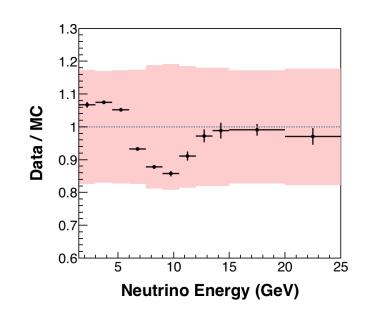
Allows closer comparison with other results.

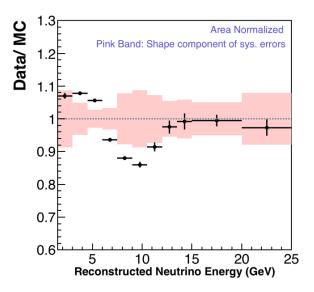
Low nu Fit to resolve Data/MC discrepancy



Charged Current Inclusive Event $u_{\mu} + N \rightarrow \mu^{+} + Hadrons$

- ν (nu) → Energy transferred (q₀) to the recoil system
- Low nu Events $\rightarrow q_o \ll E_{\nu}$
- Cross section Independent of incoming neutrino energy →Shape of Low nu distribution depends on flux shape only.



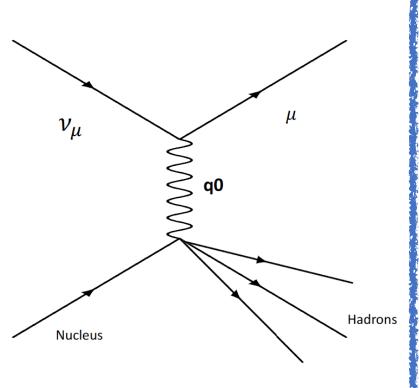


Ratio between MINERvA low nu data and simulated sample. The pink shade shows shape+normalization coverage (left) and shape coverage (right) by the systematic errors. This excludes cross section mismodeling as a candidate cause of discrepancy.

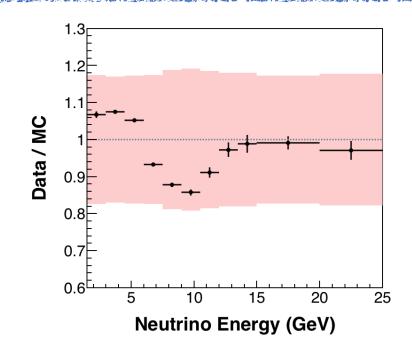
$$\frac{d\sigma}{d\nu} = \frac{G_F^2 M}{\pi} \int_0^1 (F_2 - \frac{\nu}{E_\nu} [F_2 \pm x F_3] + \frac{\nu}{2E_\nu^2} [\frac{Mx(1 - R_L)}{1 + R_L} F_2] + \frac{\nu^2}{2E_\nu^2} [\frac{F_2}{1 + R_L} \pm x F_3]) dx$$

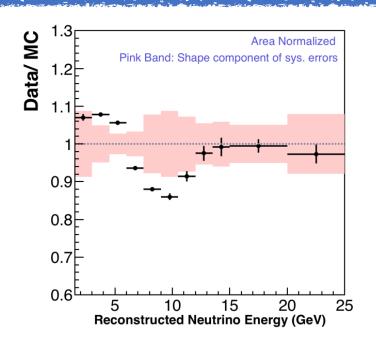
$$\frac{d\sigma}{d\nu} \approx \frac{G_F^2 M}{\pi} \int_0^1 (F_2) dx$$

Low nu Fit to resolve Data/MC discrepancy



Charged Current Inclusive Event $u_{\mu} + N
ightarrow \mu^{+} + Hadrons$





- Ratio between MINERvA low nu data and simulated sample.
- **Left**: Pink shade shows shape+normalization coverage by systematic errors.
- Right: Pink shade shows shape coverage by systematic errors.

- Low nu Events $\rightarrow q_o \ll E_{\nu}$
- Cross section Independent of incoming neutrino energy →Shape of Low nu distribution depends on flux shape only.

$$\frac{d\sigma}{d\nu} \approx \frac{G_F^2 M}{\pi} \int_0^1 (F_2) dx$$

CCQE Cross-section

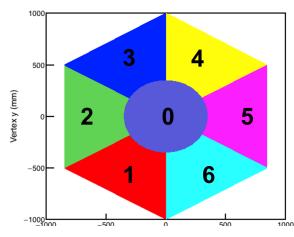
$$\frac{d\sigma}{dQ^2} \begin{pmatrix} \nu n \to l^- p \\ \bar{\nu}p \to l^+ n \end{pmatrix} = \frac{M^2 G^2 \cos^2 \theta_c}{8\pi E_{\nu}^2} [A(Q^2) \mp B(Q^2) \frac{(s-u)}{M^2} + \frac{C(Q^2)(s-u)^2}{M^4}]$$
(1)

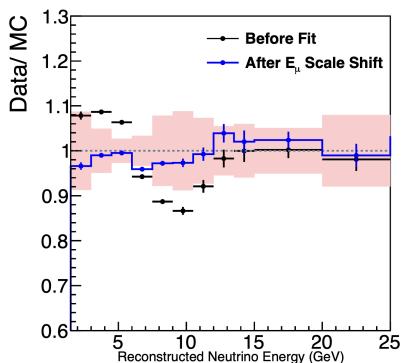
$$A(Q^{2}) = 4\frac{Q^{2}}{4M^{2}}[|F_{A}|^{2}(1 + \frac{Q^{2}}{4M^{2}}) - |F_{V}^{1}|^{2}(1 - \frac{Q^{2}}{4M^{2}}) + \frac{Q^{2}}{4M^{2}}) + \frac{Q^{2}}{4M^{2}}(1 - \frac{Q^{2}}{4M^{2}}) + 4F_{V}^{1}\xi F_{V}^{2}\frac{Q^{2}}{4M^{2}}]$$
(1)

$$B(Q^2) = 4\frac{Q^2}{4M^2} [F_A(F_V^1 + \xi F_V^2)] \tag{2}$$

$$C(Q^{2}) = \frac{1}{4} [|F_{A}|^{2} + |F_{V}^{1}|^{2} + \frac{Q^{2}}{4M^{2}} |\xi F_{V}^{2}|^{2}]$$
(3)

Low nu Fit to resolve discrepancy





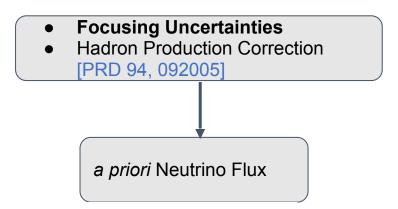
- Transverse face of the MINERvA detector was divided into 7 different areas and a multi parameter fit with the low nu sample was performed.
- Focusing + MINOS Muon energy scale as fit parameters
- Fit with and without prior uncertainty taken into account on each parameter
- Fit with (out) preferred a 3.2 (3.6)% shift in muon energy scale.
- MINERvA shifted muon energy scale by 3.6% for all sample (bottom left plot)

Parameter	Nominal	Best Fit (No Prior)	Best Fit (Prior)
Beam Position (X)	0.0 mm	$-0.3 \pm 0.3 \pm 0.1 \text{ mm}$	$-0.3 \pm 0.2 \pm 0.1 \text{ mm}$
Beam Position (Y)	0.0 mm	$0.8 \pm 0.3 \pm 0.3 \text{ mm}$	$0.7 \pm 0.2 \pm 0.2 \text{ mm}$
Target Position (X)	0.0 mm	$-0.8 \pm 0.3 \pm 0.1 \text{ mm}$	$-0.8 \pm 0.3 \pm 0.1 \text{ mm}$
Target Position (Y)	0.0 mm	$2.3 \pm 0.7 \pm 1.2 \text{ mm}$	$1.7 \pm 0.6 \pm 0.8 \text{ mm}$
Target Position (Z)	-1433 mm	$-1432.4 \pm 2.4 \pm 0.3 \text{ mm}$	$-1431 \pm 1.8 \pm 0.3$ mm
Horn 1 Position (X)	0.0 mm	$-0.3 \pm 0.4 \pm 0.5 \text{ mm}$	$-0.1 \pm 0.3 \pm 0.1 \text{ mm}$
Horn 1 Position (Y)	0.0 mm	$0.1 \pm 0.5 \pm 0.5 \text{ mm}$	$0.0 \pm 0.3 \pm 0.3 \text{ mm}$
Beam Spot Size	1.5 mm	$1.41 \pm 0.09 \pm 0.03$ mm	$1.32 \pm 0.09 \pm 0.03$ mm
Horn Water Layer	1.0 mm	$1.2 \pm 0.3 \pm 0.05 \text{ mm}$	$1.3 \pm 0.25 \pm 0.1 \text{ mm}$
Horn Current	200 kA	$198.0 \pm 1.4 \pm 1.4 \text{ kA}$	$199.1 \pm 0.7 \pm 0.5 \text{ kA}$
Muon Energy Scale	1.0	$1.032 \pm 0.004 \pm 0.008$	$1.036 \pm 0.004 \pm 0.006$

Low-nu Fit with Focusing Parameters only

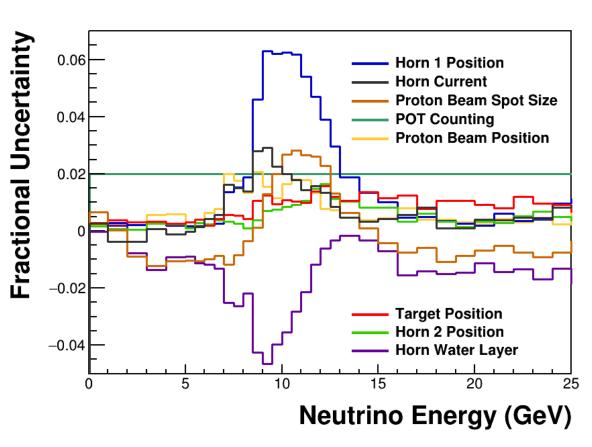
Parameter	Nominal Value	New Value
Beam Position (X)	0 mm	$-0.2 \pm 0.12 \text{ mm}$
Beam Position (Y)	0 mm	-0.53 ± 0.14
Beam Spot Size	1.5 mm	$1.22 \pm 0.14 \text{ mm}$
Horn Water Layer	1 mm	$0.895 \pm 0.16 \text{ mm}$
Horn Current	200 kA	$197.41 \pm 0.76 \mathrm{kA}$
Horn 1 Position (X)	0 mm	$0. \pm 0.17 \text{ mm}$
Horn 1 Position (Y)	0 mm	$-0.39 \pm 0.17 \text{ mm}$
Target Position (X)	0 mm	$-0.32 \pm 0.17 \text{ mm}$
Target Position (Y)	0 mm	$1.65 \pm 0.5 \text{ mm}$
Target Position (Z)	-1433 mm	-1419.44 ± 1.83 mm

MINERvA Flux Strategy



Parameter	Nominal Value	Final 1 σ shifts used
		in MINERvA analyses
Beam Position (X)	0 mm	0.4 mm
Beam Position (Y)	0 mm	0.4 mm
Beam Spot Size	1.5 mm	0.3 mm
Horn Water Layer	1.0 mm	0.5 mm
Horn Current	200 kA	1 kA
Horn 1 Position (X)	0 mm	1 mm
Horn 1 Position (Y)	0 mm	1 mm
Horn 1 Position (Z)	30 mm	2 mm
Horn 2 Position (X)	0 mm	1 mm
Horn 2 Position (Y)	0 mm	1 mm
Target Position (X)	0 mm	1 mm
Target Position (Y)	0 mm	1 mm
Target Position (Z)	-1433 mm	1 mm
POT Counting	0	2% of Total POT
Baffle Scraping	0	0.25% of POT

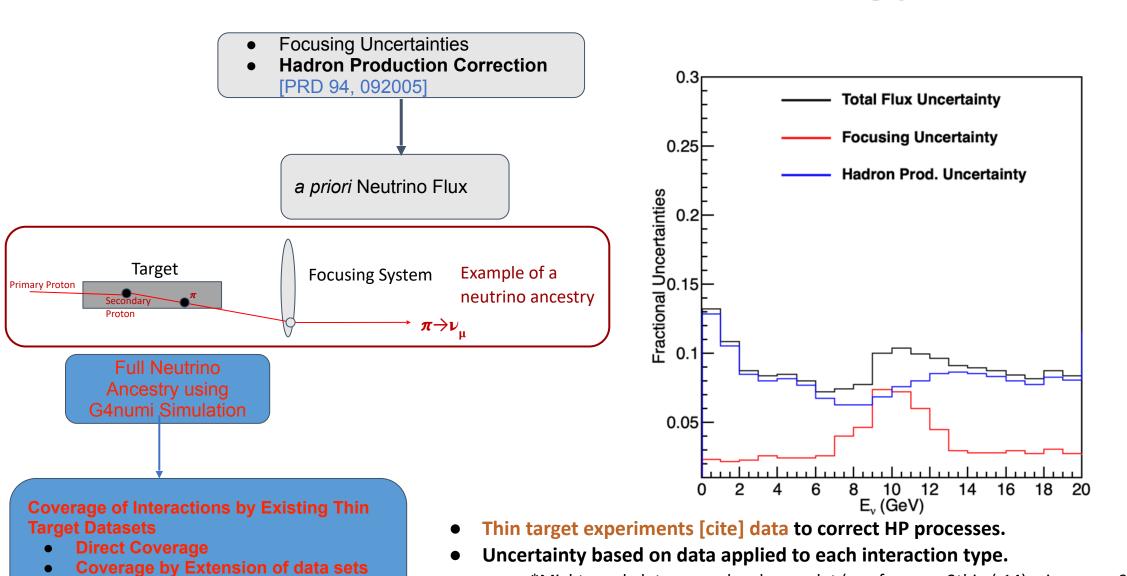
Table of Beam Parameters and their values at their nominal and shifted position.



Focusing Uncertainties in the Medium Energy neutrino Flux. Each uncertainty is the ratio of neutrino flux due to shifted beam parameter (by $+1 \sigma$) to the nominal neutrino flux.

MINERvA Flux Strategy

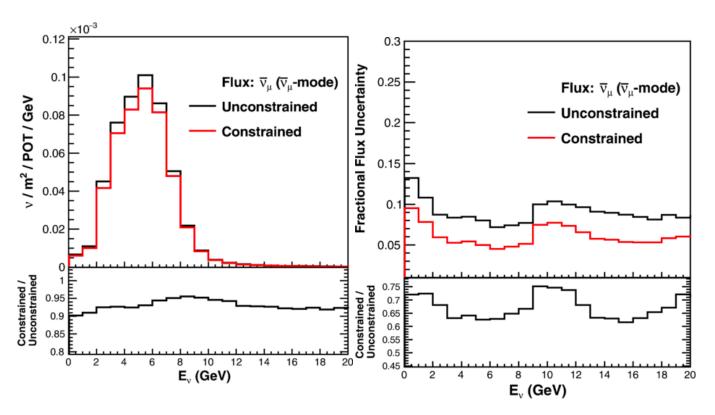
*Might need plot approval on lower plot (unc from gen2thin (-14) minervame6A)



No coverage at all

84

MINERvA Flux Strategy



a priori anti neutrino Fluxbefore and after constrainedby ve scattering data

Uncertainty on predicted anti neutrino flux before and after constrained by νe scattering data

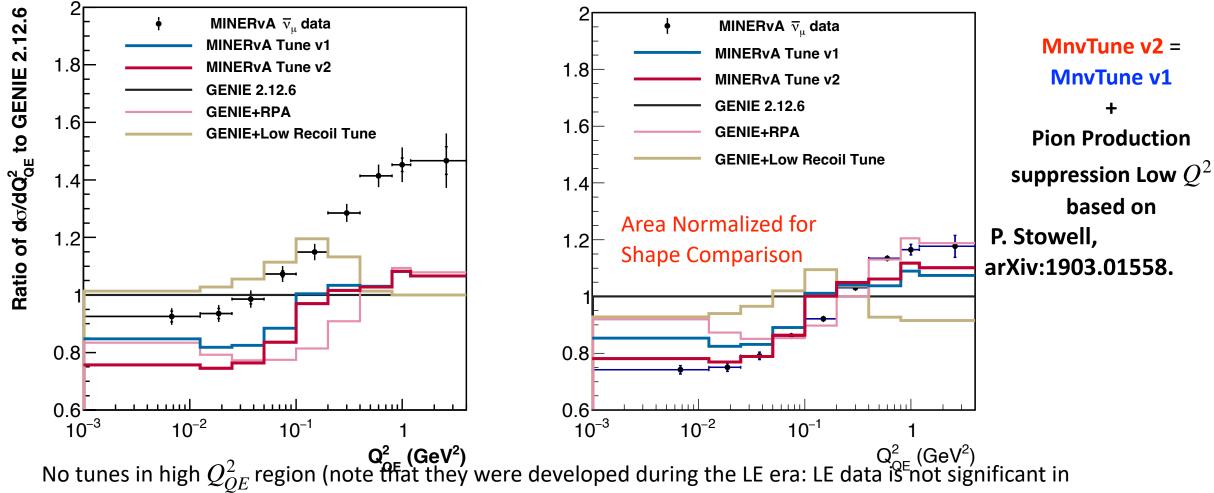
- ν +e constraint [PRD 107,012001]
 - $\circ \nu e \rightarrow \nu e$
- Inverse Muon Decay [PRD 104,092010]

$$\circ$$
 ν_{μ} e $\rightarrow \mu^{-}\nu_{e}$

Neutrino Flux (left) and uncertainty before and after IMD constraint.

Because of $\mathbf{E}_{\nu} \cong \mathbf{11}$ GeV, only constrain high energy region.

Cross section (Q_{OE}^2): MINERvA Tunes



this region).

Plots show that our tunes predict the shape of the distribution relatively well (right) compared to absolute distribution.

Systematic Uncertainties

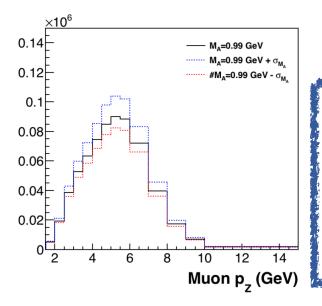
Reconstruction Related Uncertainties:

Uncertainties related to reconstruction of blobs, tracks, PID etc

GENIE Model Related Uncertainties

Uncertainties coming from GENIE models to model neutrino interactions.

They get propagated to our data during background subtraction, efficiency correction and unfolding.



Two Universe Method

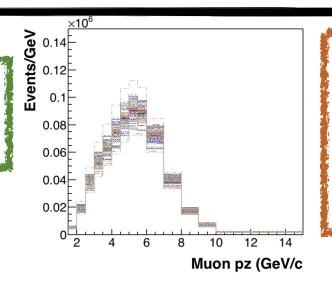
- ullet Shift the **parameter by** $\pm 1\sigma$
- Get the alternate distributions (universe).
- Calculate uncertainty from the spread relative to central value (CV) universe.

A GENIE M_A related systematic

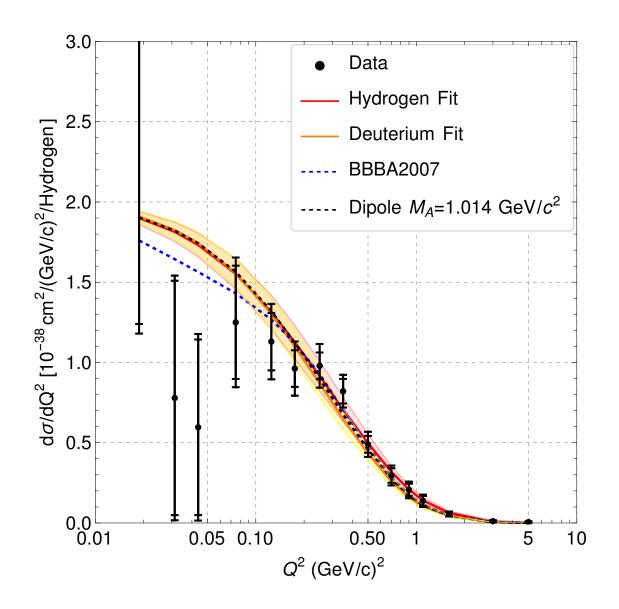
Flux Uncertainties (Already covered):

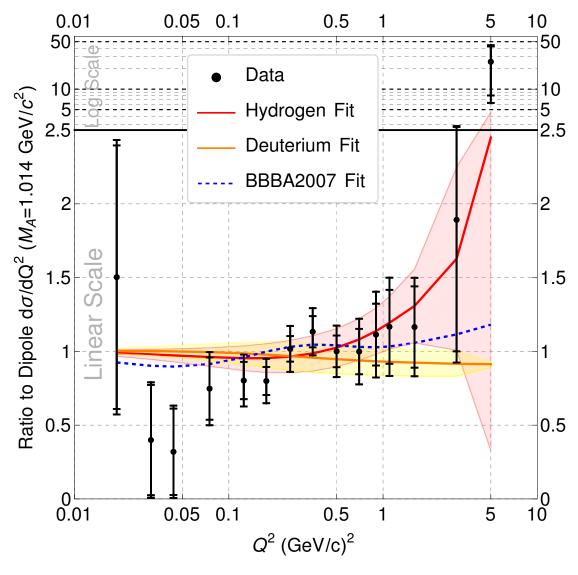
- Focusing Uncertainties
- Hadron Production Uncertainties

Multi Universe Method



- Many parameters (correlated)
- Shift parameters within $\pm 1\sigma$
- Generate N universes (500 in this analysis)
- Calculate uncertainty from the spread of the distribution.





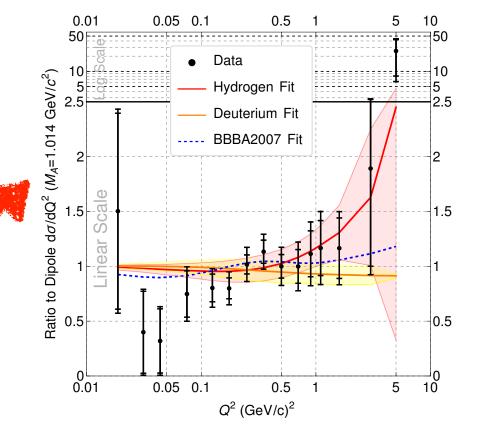
$\bar{\nu}_{\mu}$ CCQE (H) Cross-section

Article Open Access Published: 01 February 2023

Measurement of the axial vector form factor from antineutrino-proton scattering

T. Cai , M. L. Moore, A. Olivier, S. Akhter, Z. Ahmad Dar, V. Ansari, M. V. Ascencio, A. Bashyal, A. Bercellie, M. Betancourt, A. Bodek, J. L. Bonilla, A. Bravar, H. Budd, G. Caceres, M. F. Carneiro, G. A. Díaz, H. da Motta, J. Felix, L. Fields, A. Filkins, R. Fine, A. M. Gago, H. Gallagher, ... L. Zazueta + Show authors

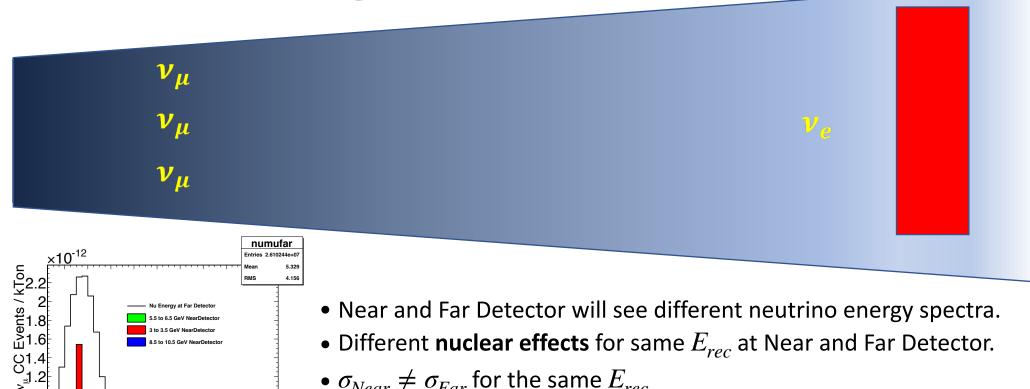
Recent W&C Seminar by T. Cai on ${\cal F}_A$ extraction from Hydrogen atoms (Free Nucleon)



- $F_A \rightarrow$ Axial Form Factor
- $F_A(Q^2) = -\frac{g_A}{(1 + \frac{Q^2}{M_A^2})^2}$

- Based on bubble chamber (hydrogen targets) measurements
- Measurements in **heavier target** report slightly **higher axial mass**
 - Nuclear Effects
- ullet Dipole Form Factor approximation breaks at high Q^2

$N_e(E_{rec}, L) \propto \sum \Phi_e(E, L) \sigma_i(E) f_{\sigma i}(E, Erec) dEdM$



- Near and Far Detector will see different neutrino energy spectra.
- ullet Different **nuclear effects** for same E_{rec} at Near and Far Detector.
- $\sigma_{Near} \neq \sigma_{Far}$ for the same E_{rec}

0.8 0.6 0.4

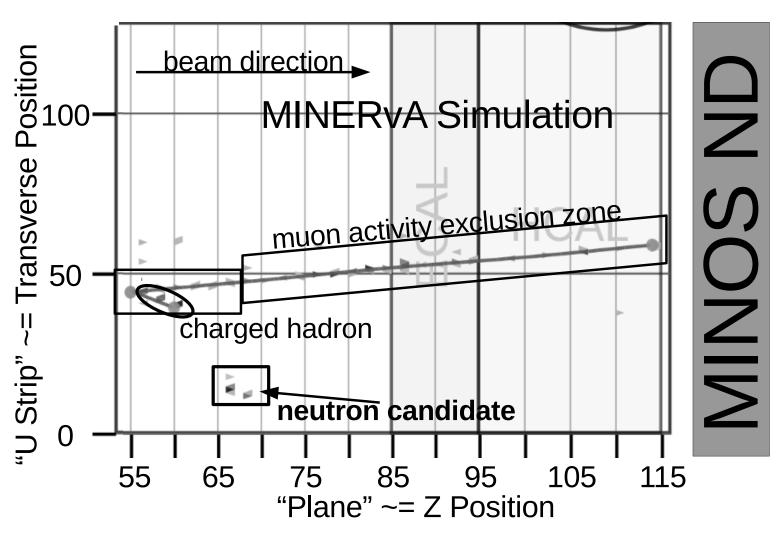
6 8 10 12 14 16 18 20

Far Detector Energy(GeV)

Need to understand the nuclear effects in heavy nucleus.

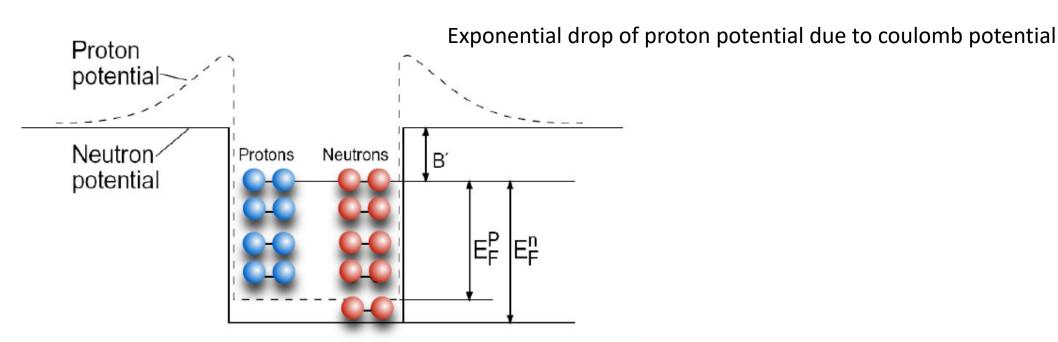
MINERvA was designed to study the poorly understood neutrino-heavy nucleus interactions

Neutron Reconstruction in MINERvA



Paper in Preparation to submit to the PRD.

Nucleus and Nucleons: Relativistic Fermi Gas Model



Relativistic Fermi Gas:

- Nucleons as independent particles with some fermi momentum in a mean field generated by the rest of the nucleus.
- Nucleons can interact with other nucleons.

