



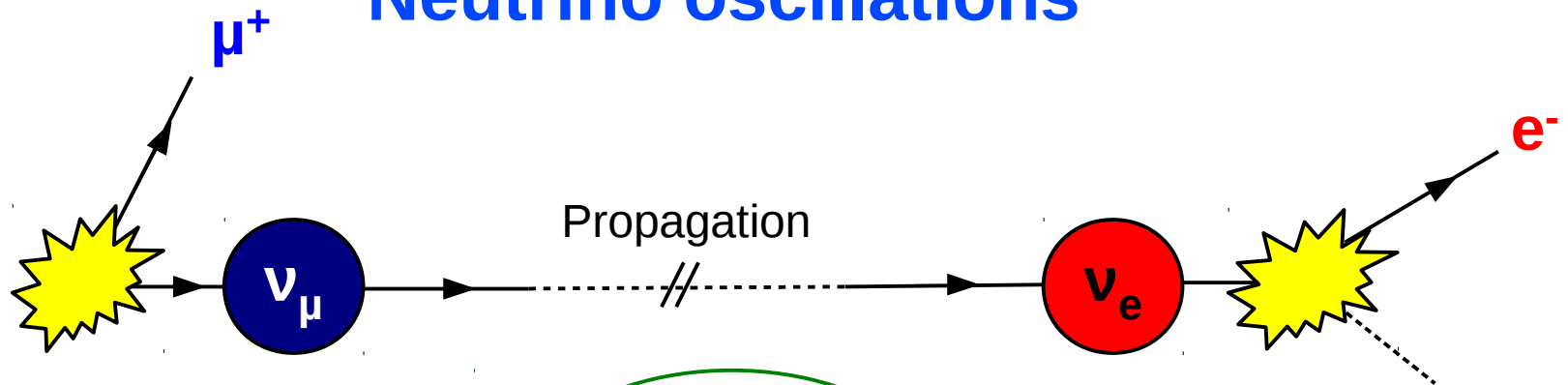
# Neutrino oscillation results from T2K

C. Bronner on behalf of the T2K collaboration

- Neutrino oscillations and long-baseline experiments
- The Tokai to Kamioka experience
- Description of analysis chain
- Systematic uncertainties
- Dataset
- Oscillation analysis results
- Perspective for the future

# Neutrino oscillations and long-baseline experiments

# Neutrino oscillations



Flavor eigenstates  
(interaction)

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix}$$

$$= \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \times$$

$$\begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Mass eigenstates  
(propagation)

Mixing (or Pontecorvo-Maki-Nagawa-Sakata) matrix  
link between the two sets of eigenstates

$P(\nu_\alpha \rightarrow \nu_\beta)$  oscillates as a function of distance  $L$  traveled by the neutrino

- Amplitude of oscillations depends on the mixing matrix  $U$
- Phase of the oscillation depends on energy and difference of mass squared:  $\Delta m^2_{ij} L/E$

$$(\Delta m^2_{ij} = m^2_i - m^2_j)$$



# Neutrino oscillations Parameters

In practice, for neutrino oscillations:

$$U = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix}}_{\text{“Atmospheric”}} \underbrace{\begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix}}_{\text{“Reactor”}} \underbrace{\begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\text{“Solar”}}$$

“Atmospheric”

“Reactor”

“Solar”

$(c_{ij} = \cos(\theta_{ij}), s_{ij} = \sin(\theta_{ij}))$

$P(\nu_\alpha \rightarrow \nu_\beta)$  depends on **6 parameters**:

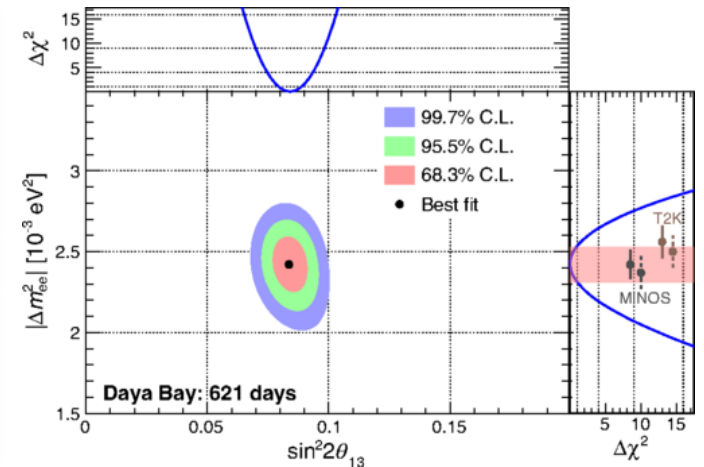
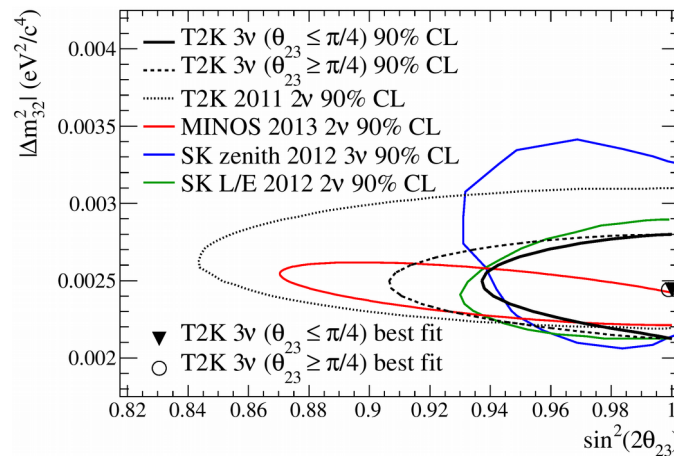
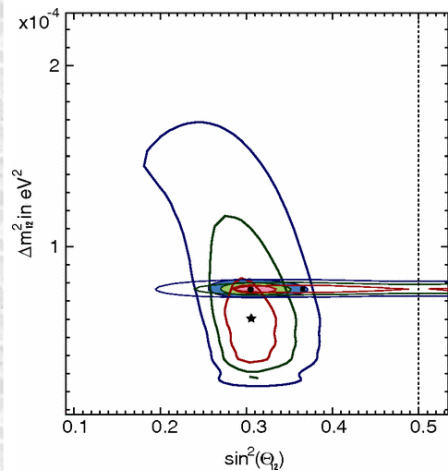
- 3 mixing angles  $\theta_{12}, \theta_{23}, \theta_{13}$
- 2 independent mass splittings  $\Delta m^2_{ij}$
- 1 complex phase, the **CP phase  $\delta$**

# Neutrino oscillations measurements Status

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**All mass splittings and mixing angles have been measured to be non-zero**

- $\Delta m_{21}^2$  and  $\theta_{21}$  from solar neutrinos and KamLAND
- $\Delta m_{32/31}^2$  and  $\theta_{23}$  from atmospheric neutrinos and later beam neutrinos
- $\theta_{13}$  from reactor  $\bar{\nu}_e$  disappearance and beam  $\nu_e$  appearance



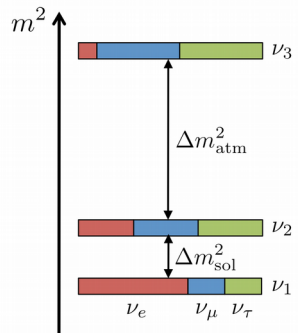
Observed **both the disappearance of neutrinos** (atmospheric, solar and reactor neutrinos) of a certain flavor, and **appearance of a different flavor** of neutrino (T2K, OPERA, NOvA)

# Neutrino oscillations measurements

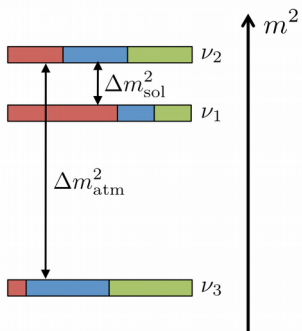
## What are we still looking for?

Mass hierarchy:  
 $m_3 > m_2, m_1$ ?

normal hierarchy (NH)



inverted hierarchy (IH)



PDG 2016 summary table

Parameter	best-fit	$3\sigma$
$\Delta m_{21}^2$ [ $10^{-5}$ eV <sup>2</sup> ]	7.37	6.93 – 7.97
$ \Delta m^2 $ [ $10^{-3}$ eV <sup>2</sup> ]	2.50 (2.46)	2.37 – 2.63 (2.33 – 2.60)
$\sin^2 \theta_{12}$	0.297	0.250 – 0.354
$\sin^2 \theta_{23}, \Delta m^2 > 0$	0.437	0.379 – 0.616
$\sin^2 \theta_{23}, \Delta m^2 < 0$	0.569	0.383 – 0.637
$\sin^2 \theta_{13}, \Delta m^2 > 0$	0.0214	0.0185 – 0.0246
$\sin^2 \theta_{13}, \Delta m^2 < 0$	0.0218	0.0186 – 0.0248
$\delta/\pi$	1.35 (1.32)	(0.92 – 1.99) ((0.83 – 1.99))

Octant of  $\theta_{23}$ :

$\theta_{23} > \pi/4$ ?

$\theta_{23} < \pi/4$ ?

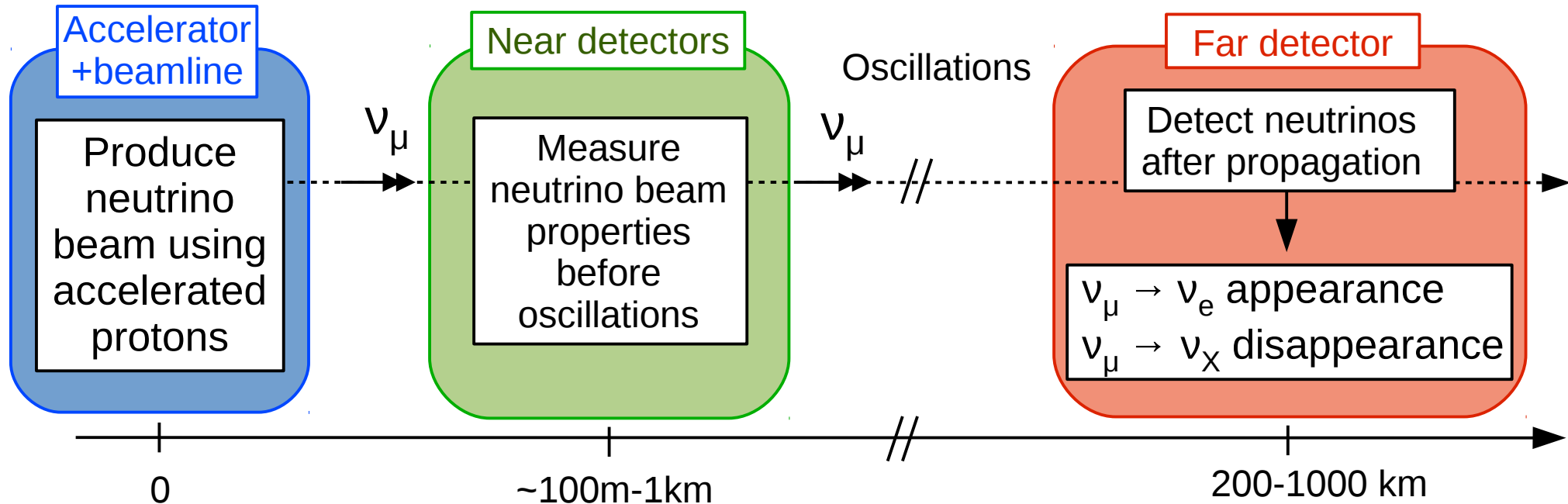
Violation of CP symmetry in neutrino oscillations?

+ more precise tests of the PMNS model via different channels

# Long baseline experiments Concept

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Man-made neutrino beam produced by an accelerator



## Several advantages:

- Better knowledge and control of neutrino flux
- Can select neutrino energy range
- Can use near detectors to reduce uncertainties
- Know direction of neutrinos reaching far detector
- Can produce either neutrino or anti-neutrino beam (compare oscillations of neutrinos and anti-neutrinos)

# Long-baseline experiments

## First measurements

In first approximation LBL experiments can measure some of the PMNS parameters through exclusive channels:

### Far detector $\nu_\mu$ events

$\nu_\mu \rightarrow \nu_\mu$  disappearance

$$P(\nu_\mu \rightarrow \nu_\mu) \approx 1 - \sin^2(2\theta_{23}) \sin^2\left(1.27 \frac{\Delta m^2 \times L}{E}\right)$$

Precise measurement of  $\theta_{23}$  and  $|\Delta m^2|$

### Far detector $\nu_e$ events

$\nu_\mu \rightarrow \nu_e$  appearance

$$P(\nu_\mu \rightarrow \nu_e) \approx \sin^2(\theta_{23}) \sin^2(2\theta_{13}) \sin^2\left(1.27 \frac{\Delta m^2 \times L}{E}\right)$$

- Observation of  $\nu_e$  appearance
- Measurement of  $\theta_{13}$

And similar measurements for anti-neutrinos

# Long baseline experiments

## Main current physics goals

Look for more subtle effects by comparing  $P(\nu_\mu \rightarrow \nu_e)$  and  $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$

- CP violation: is  $\sin(\delta) \neq 0$ ?
- Mass hierarchy: sign of  $\Delta m_{32}^2$  ?

### Full probability in vacuum:

$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_e) = & 4c_{13}^2 s_{13}^2 s_{23}^2 \sin^2 \Delta_{31} \\
 & + 8c_{13}^2 s_{12} s_{13} s_{23} (c_{12} c_{23} \cos \delta - s_{12} s_{13} s_{23}) \cos \Delta_{32} \sin \Delta_{31} \sin \Delta_{21} \\
 & - 8c_{13}^2 c_{12} c_{23} s_{12} s_{13} s_{23} \sin \delta \sin \Delta_{32} \sin \Delta_{31} \sin \Delta_{21} \\
 & + 4s_{12}^2 c_{13}^2 (c_{12}^2 c_{23}^2 + s_{12}^2 s_{23}^2 s_{13}^2 - 2c_{12} c_{23} s_{12} s_{23} s_{13} \cos \delta) \sin^2 \Delta_{21}
 \end{aligned}$$

$$\sin^2 \Delta_{ij} = \sin^2(1.27 \Delta m_{ij}^2 \times L/E)$$

### In matter leading term

$$P(\nu_\mu \rightarrow \nu_e) \approx \sin^2(\theta_{23}) \sin^2(2\theta_{13}) \sin^2\left(1.27 \frac{\Delta m^2 \times L}{E}\right)$$

$$\text{Multiplied by } 1 + \frac{2a}{\Delta m_{31}^2} (1 - 2\sin^2(\theta_{13}))$$

$$(a \equiv 2\sqrt{2} G_F n_e E)$$

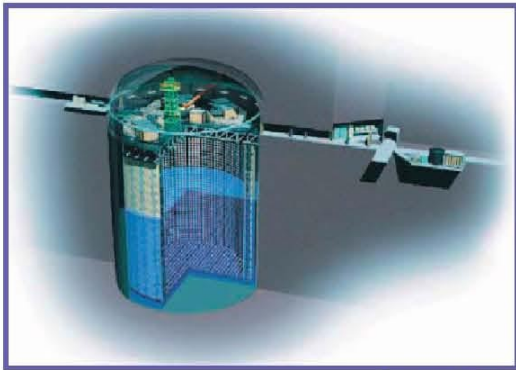
$$\begin{array}{l}
 \nu \rightarrow \bar{\nu} \\
 \delta \rightarrow -\delta \\
 a \rightarrow -a
 \end{array}$$

Not too long baseline (~300km):  
Mainly effect of  $\delta$ : **T2K** (~<27% vs ~10%)

Very long baseline: effect of  $\delta$  and  
matter effect: **NOvA**



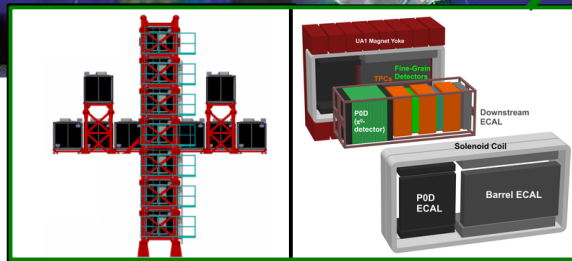
# The Tokai to Kamioka experiment



**Super-Kamiokande**  
(ICRR, Univ. Tokyo)



**J-PARC Main Ring**  
(KEK-JAEA, Tokai)



# The T2K experiment

## The collaboration



~ 500 members, 62 Institutes, 11 countries

### Canada

TRIUMF  
U. B. Columbia  
U. Regina  
U. Toronto  
U. Victoria  
U. Winnipeg  
York U.

### Italy

INFN, U. Bari  
INFN, U. Napoli  
INFN, U. Padova  
INFN, U. Roma

### Japan

ICRR Kamioka  
ICRR RCCN  
Kavli IPMU  
KEK  
Kobe U.  
Kyoto U.  
Miyagi U. Edu.  
Okayama U.  
Osaka City U.  
Tokyo Metropolitan U.  
U. Tokyo  
Yokohama National U.

### France

CEA Saclay  
IPN Lyon  
LLR E. Poly.  
LPNHE Paris

### Germany

Aachen

### Switzerland

ETH Zurich  
U. Bern  
U. Geneva

### Poland

IFJ PAN, Cracow  
NCBJ, Warsaw  
U. Silesia, Katowice  
U. Warsaw  
Warsaw U. T.  
Wroclaw U.

### Russia

INR

### United Kingdom

Imperial C. London  
Lancaster U.  
Oxford U.  
Queen Mary U. L.  
Royal Holloway U.L.  
STFC/Daresbury  
STFC/RAL  
U. Liverpool  
U. Sheffield  
U. Warwick

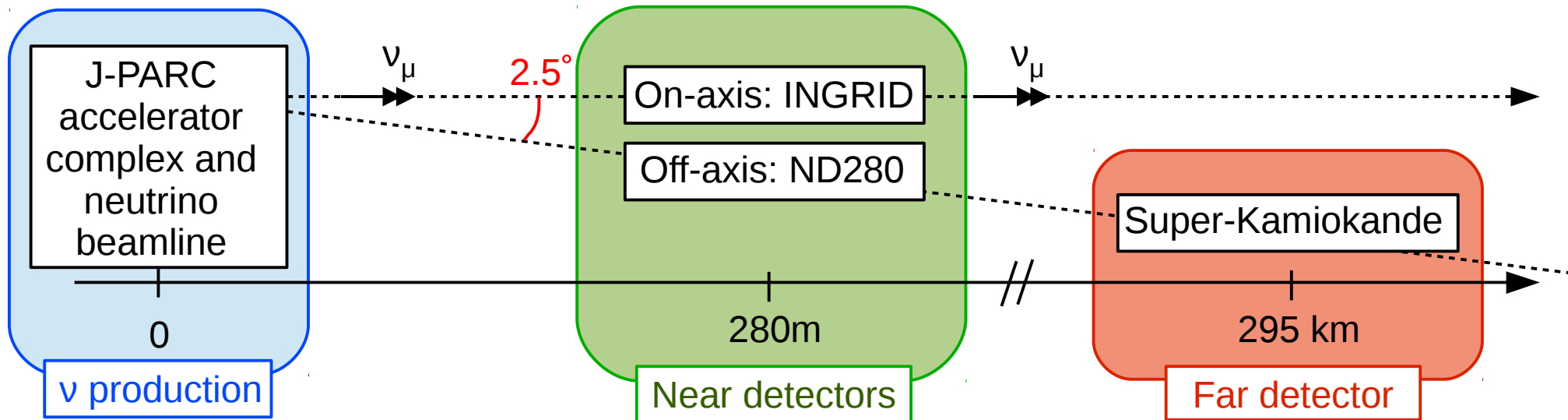
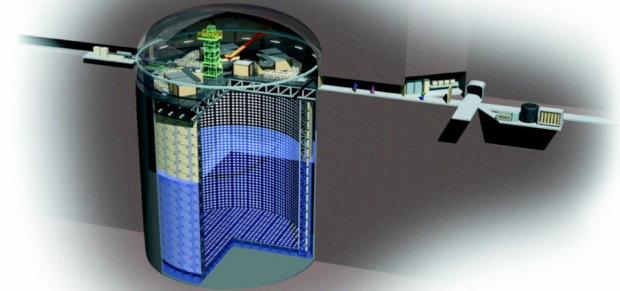
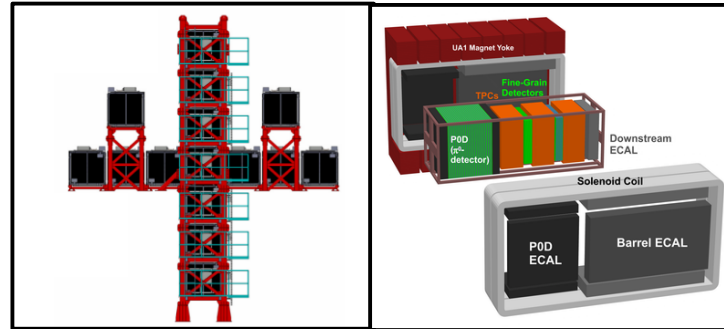
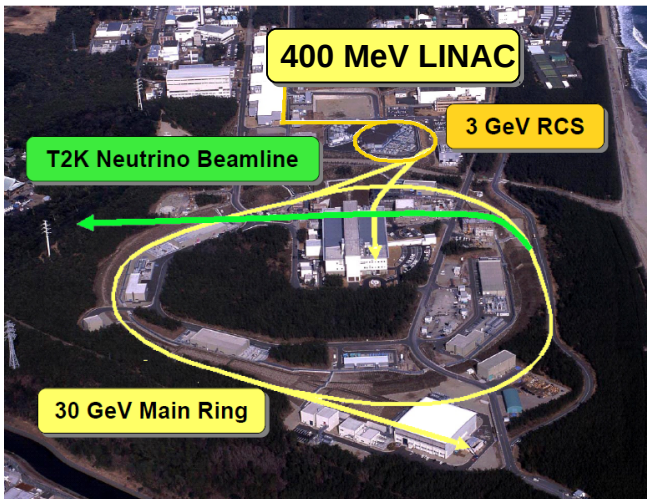
### USA

Boston U.  
Colorado S. U.  
Duke U.  
Louisiana State U.  
Michigan S.U.  
Stony Brook U.  
U. C. Irvine  
U. Colorado  
U. Pittsburgh  
U. Rochester  
U. Washington





# The T2K experiment Overview

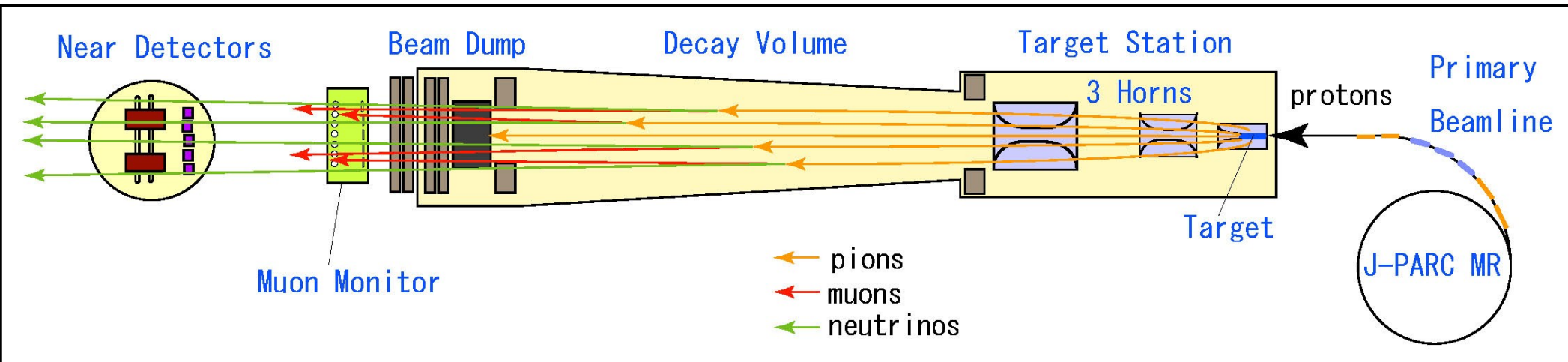


- Baseline: 295 km
- Off-axis beam

# The T2K experiment

## Neutrino production

Conventional neutrino beam produced from 30 GeV protons

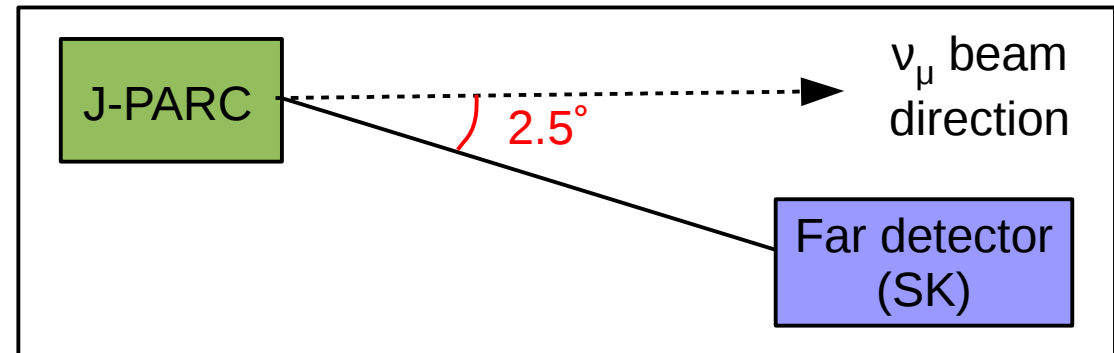
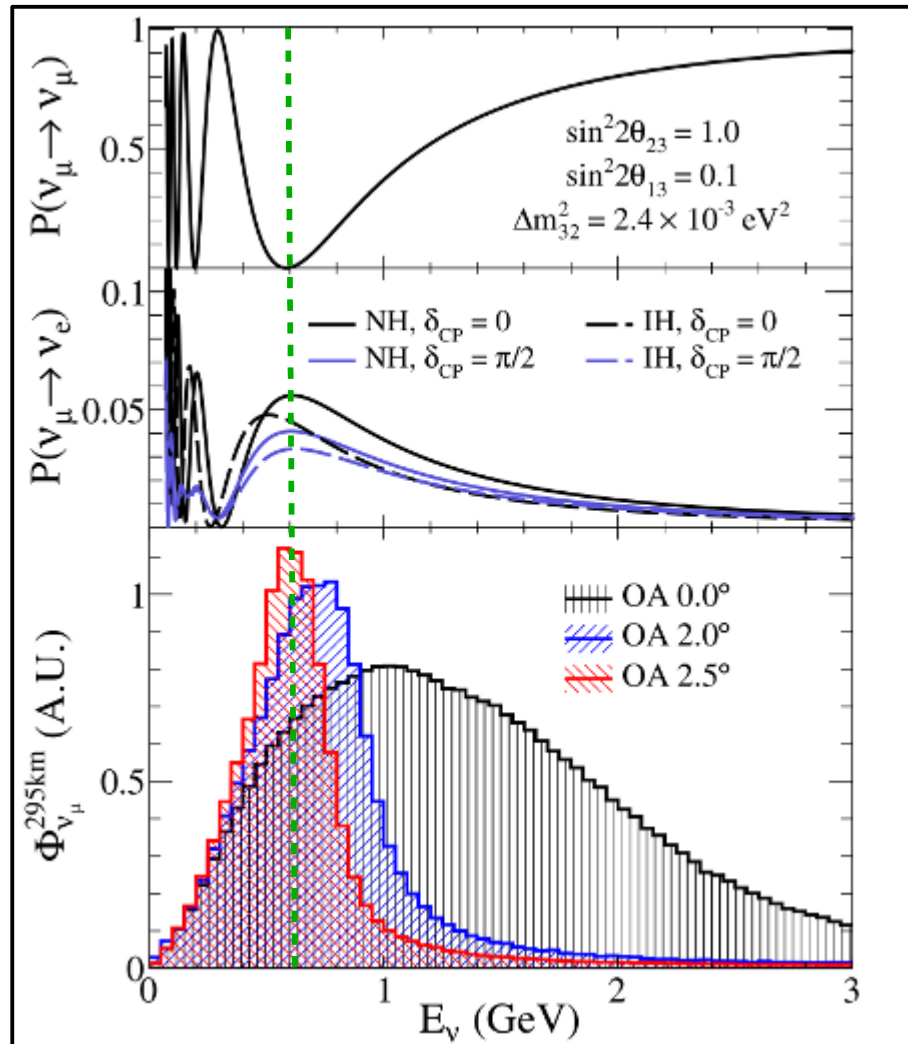


**Almost pure  $\nu_\mu/\bar{\nu}_\mu$  beam,**  
with an intrinsic  $\nu_e/\bar{\nu}_e$   
component (<1% at peak)

Can switch from  $\nu_\mu$  beam to  
 $\bar{\nu}_\mu$  beam by inverting the horn  
polarities

# The T2K experiment

## Off-axis beam



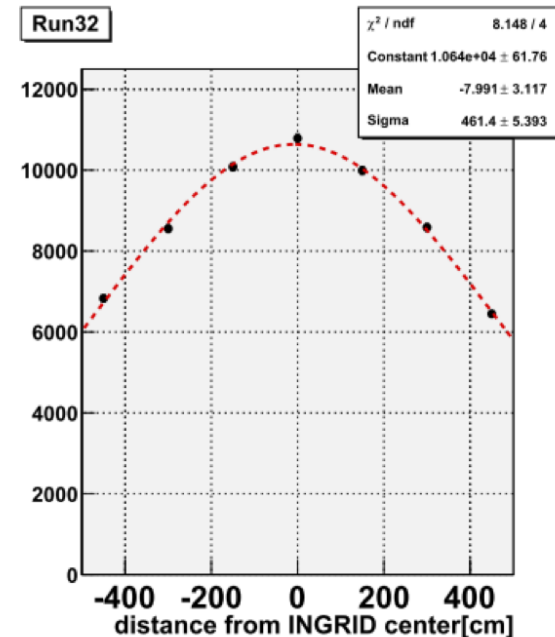
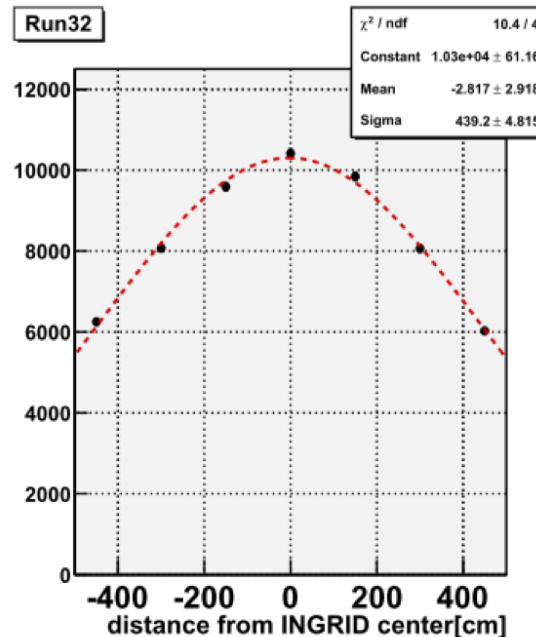
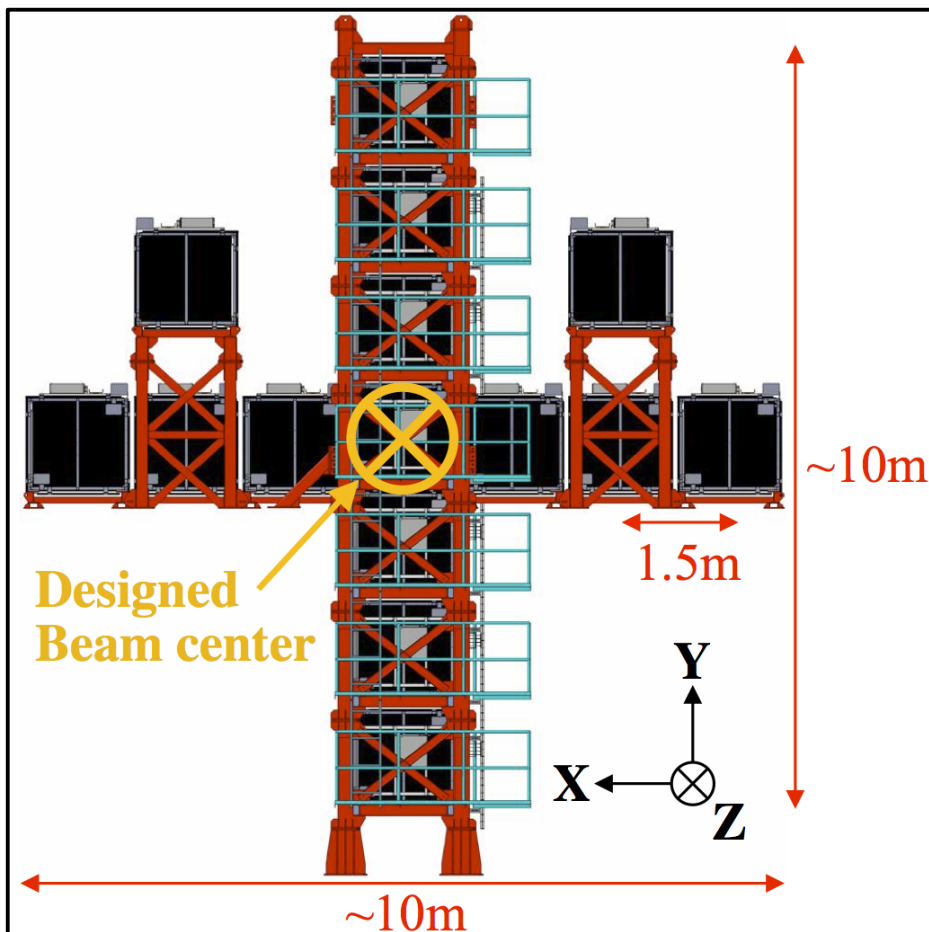
- Narrow band neutrino beam, peaked at oscillation maximum (0.6 GeV)
- Reduces high energy tail
- Reduces intrinsic  $\nu_e$  contamination of the beam at peak energy
- Interactions dominated by CCQE mode

# The T2K experiment

## Near detectors

On-axis detector INGRID (Interactive Neutrino GRID)  
 Located 280m from the target

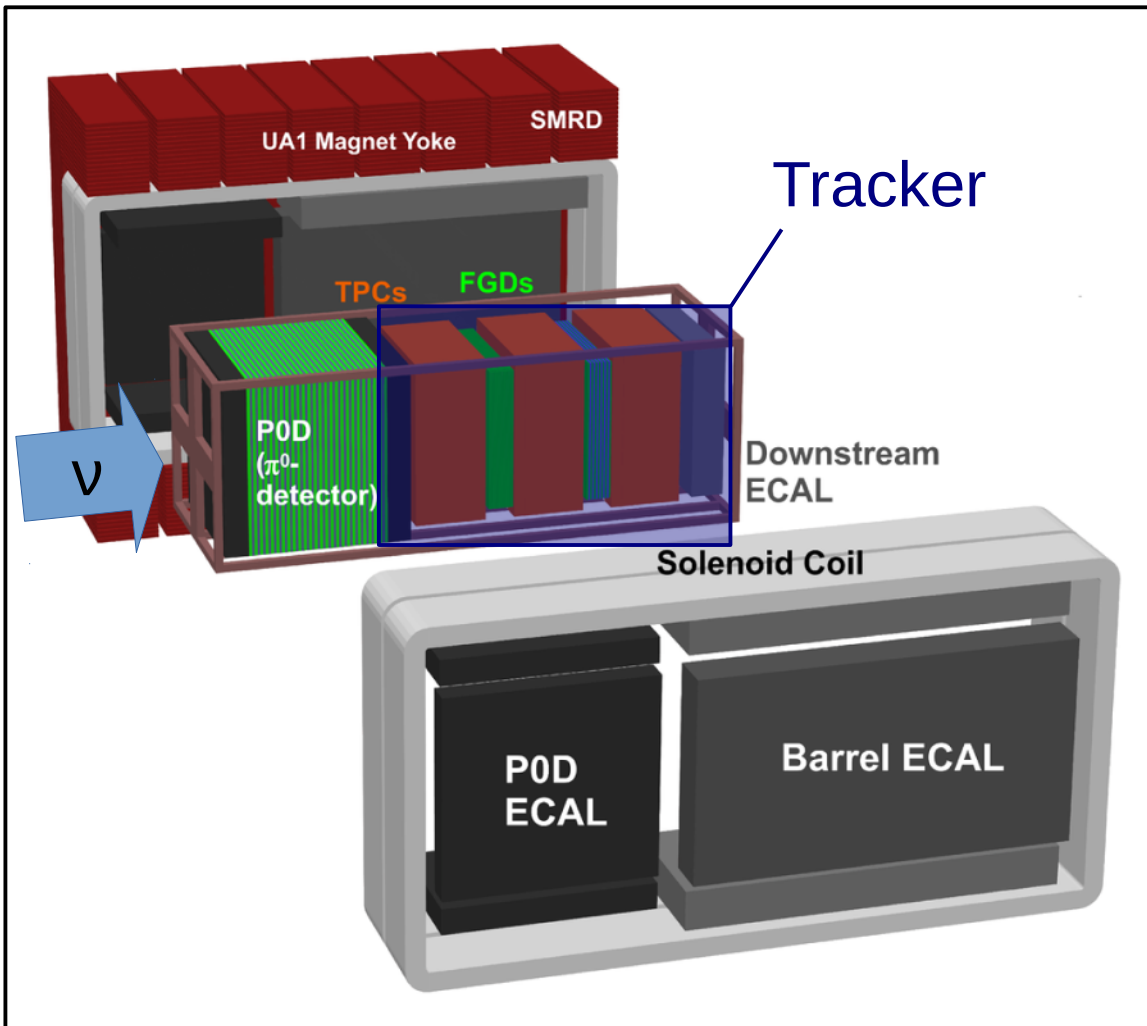
- 16 identical modules made of iron and scintillators
- 'counting neutrinos' by reconstructing muon tracks from  $\nu_\mu$  interactions
- Monitors neutrino beam: rate, direction and stability



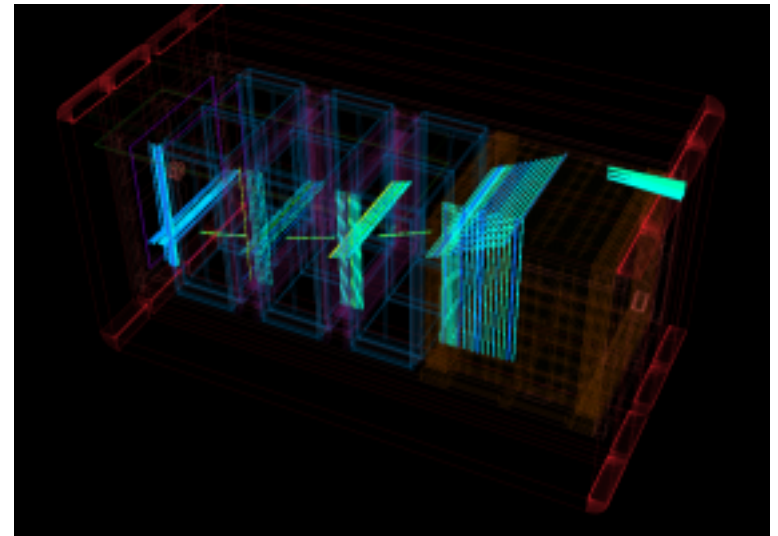
# The T2K experiment

## Off-axis near detectors

Off-axis near detector ND280  
Located 280m from the target



- Several detectors inside a 0.2 T magnetic field
- Good tracking capabilities
- 'Tracker' used to constrain flux and interaction uncertainties for oscillation analysis
- Rich cross-section measurement program



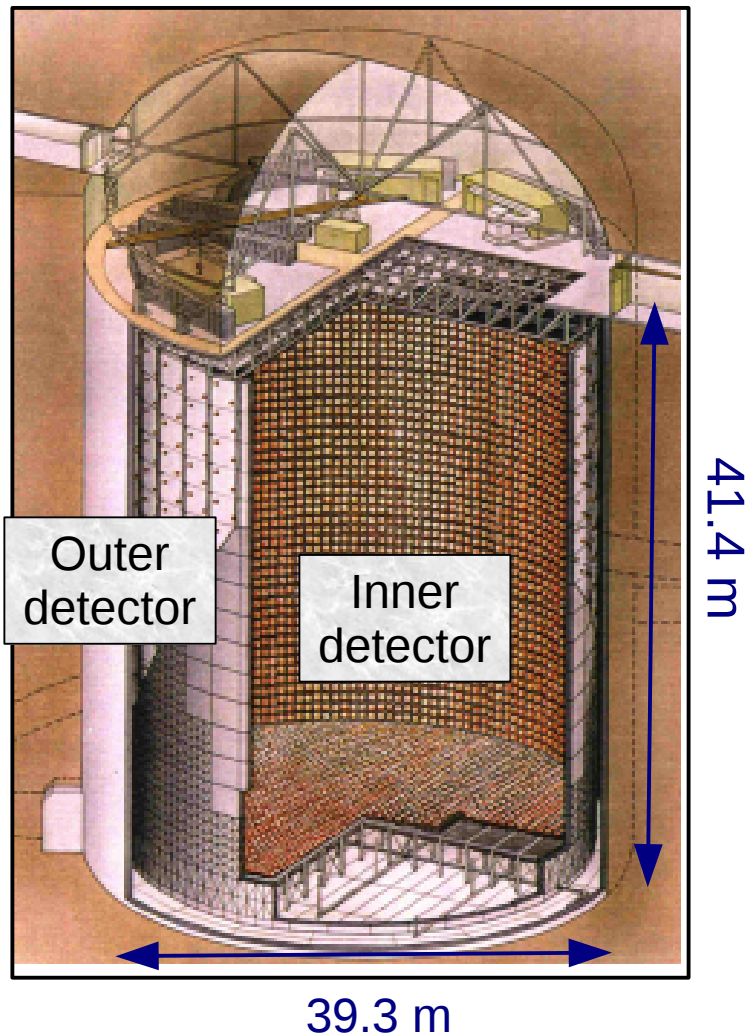


# The T2K experiment

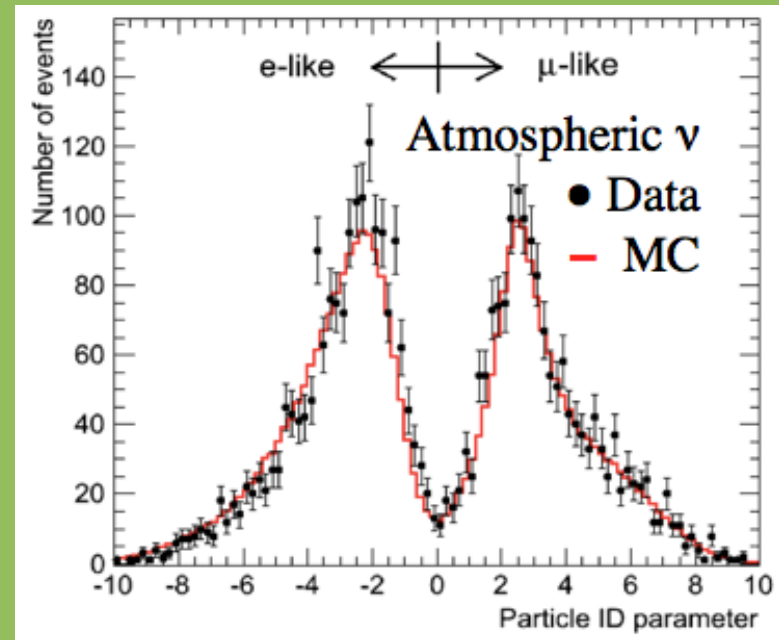
## Far detector: Super-Kamiokande

Located 295 km from the target  
Synchronized with beamline via GPS

- 50 kt (22.5 kt fiducial) water Cherenkov detector
- Operational since 1996



Good separation between  $\mu^\pm$  and  $e^\pm$   
(separate  $\nu_\mu$  and  $\nu_e$  CC interactions)

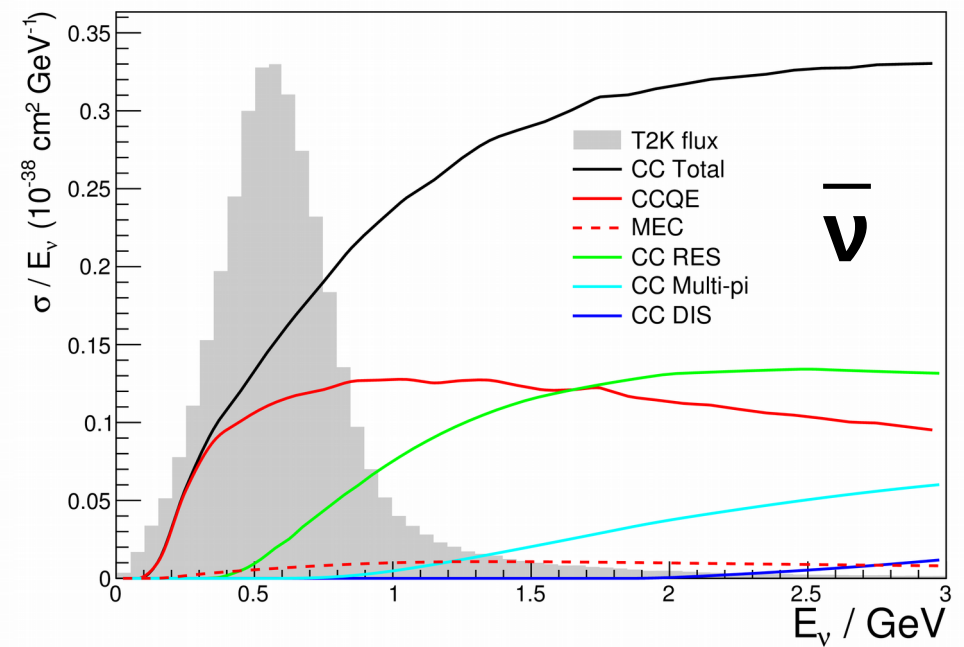
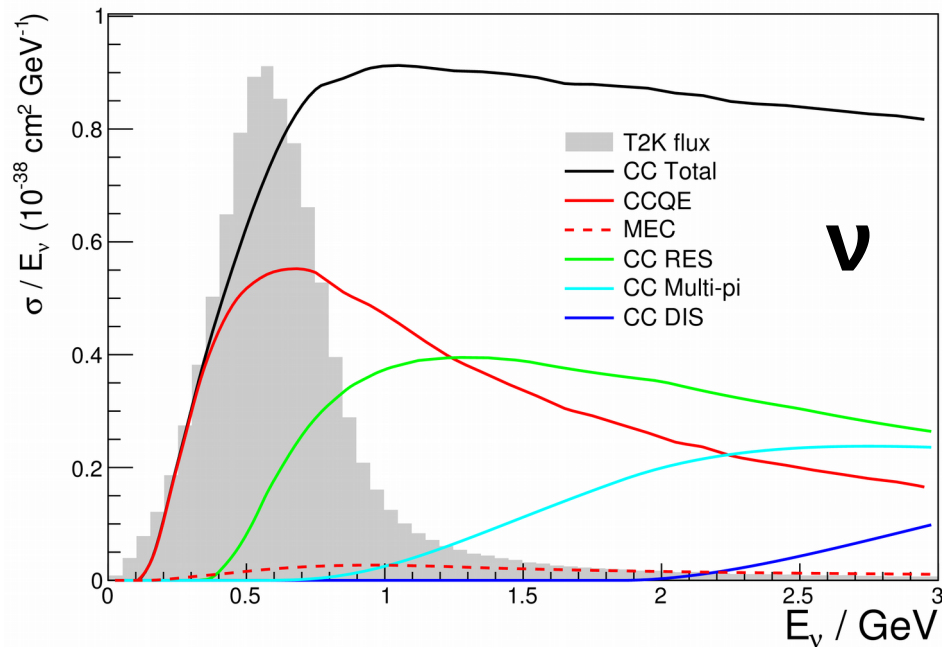


No magnetic field: cannot separate  $\nu$  and  $\bar{\nu}$  on an event by event basis

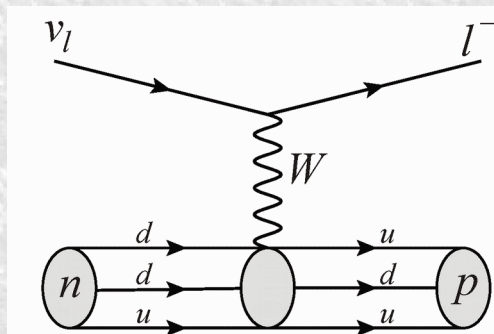
# The T2K experiment

## Neutrino interactions

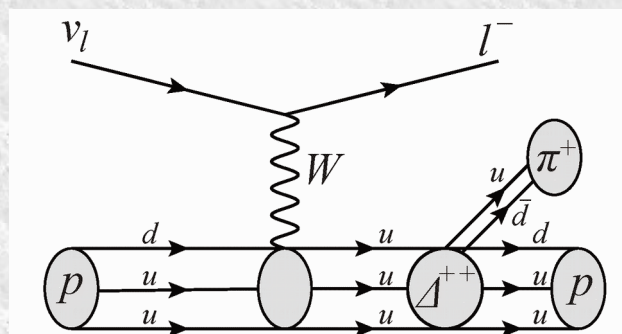
- Need to detect neutrino flavor => charged-current interactions
- At T2K energies, dominant interaction mode is charged-current quasi-elastic



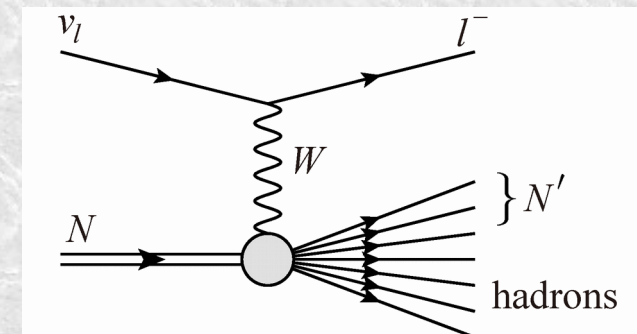
### CCQE



### CC RES



### CC DIS/Multi-pi



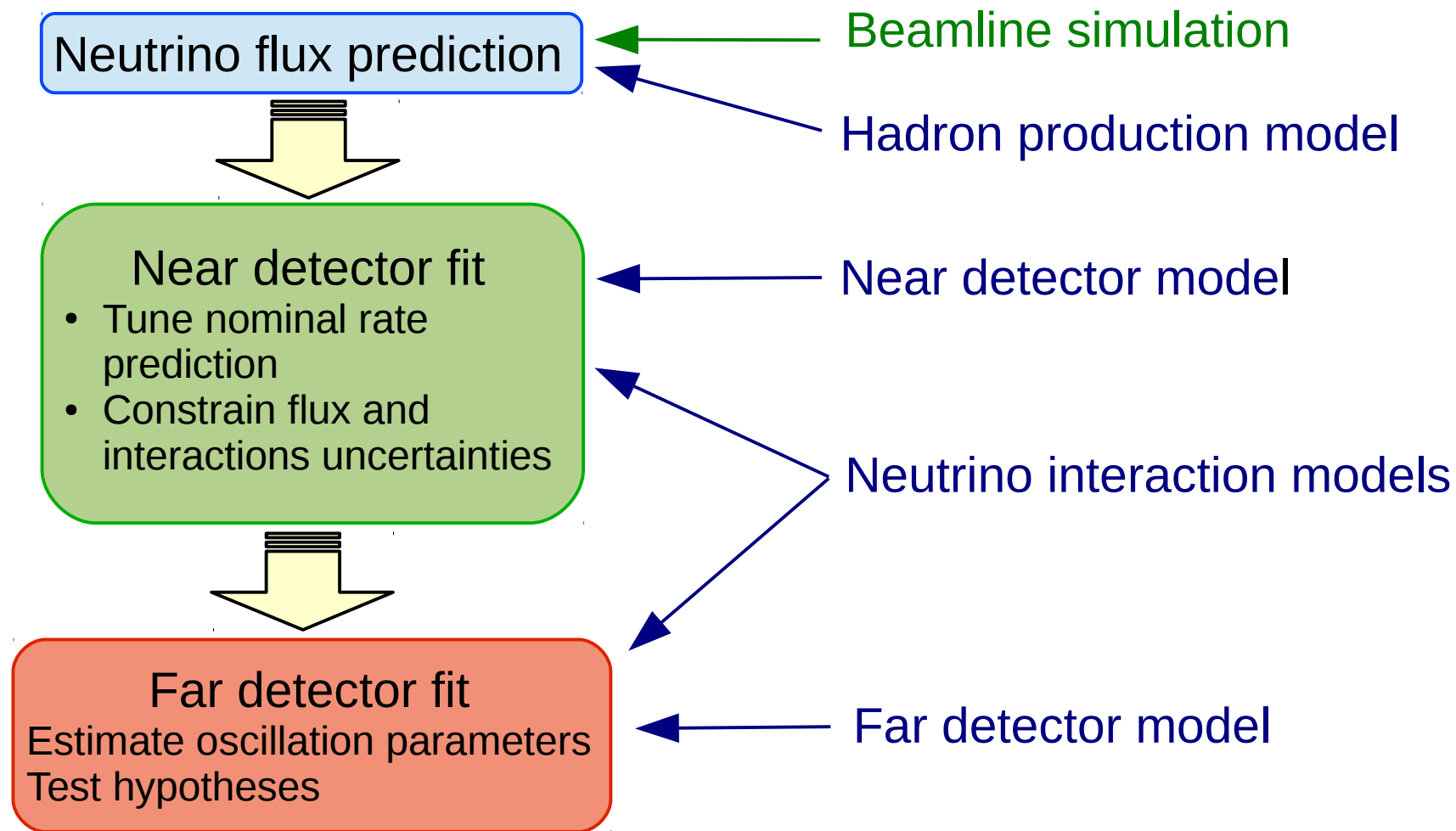
# Oscillation analysis

## Analysis description



# Analysis description overview

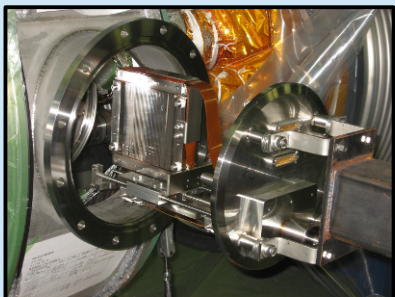
Likelihood analysis: compare observed data at the far detector to predictions based on a model of the experiment to make statistical inferences



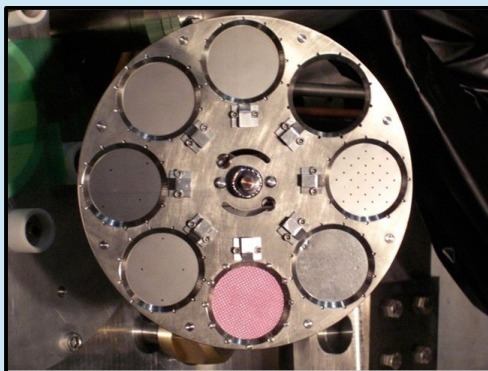
# Analysis description

## Neutrino flux prediction

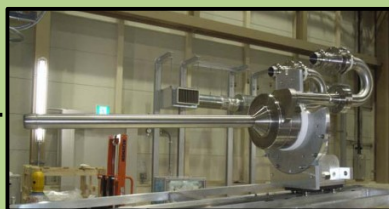
### Proton beam properties



Measured by beam monitors



### Hadron production in target



$\pi^\pm$

$K^\pm$

FLUKA 2011  
Tuned to external data  
(NA61/Shine @ CERN)

### Propagation and decay of hadrons in secondary beamline

$\mu^\pm$

$\nu_\mu/\bar{\nu}_\mu$

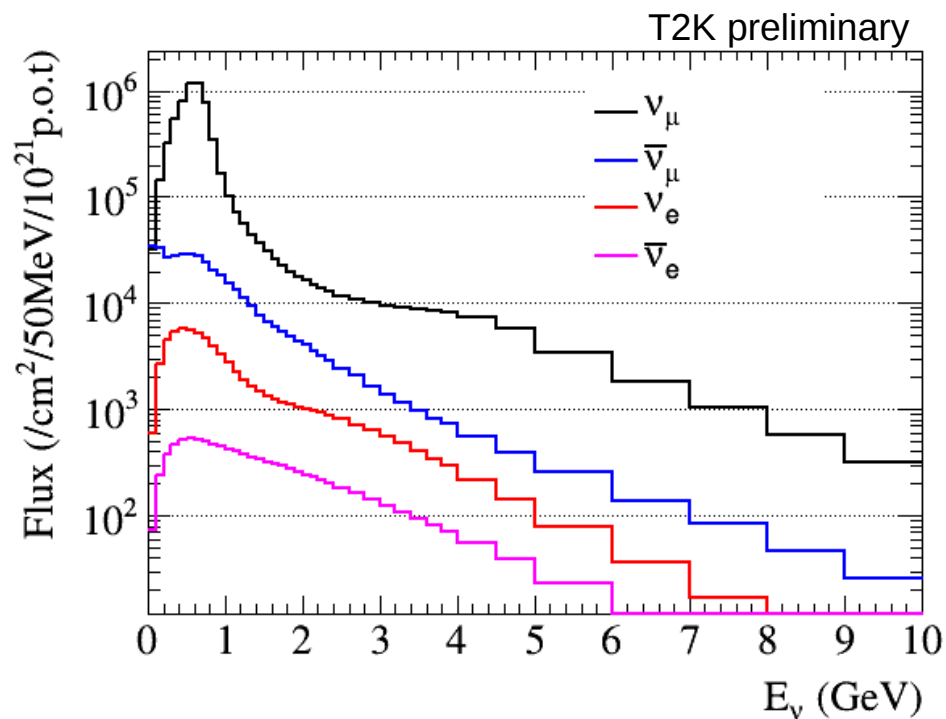
GEANT3 simulation  
GCALOR package

Neutrino flux predicted using a series of simulations

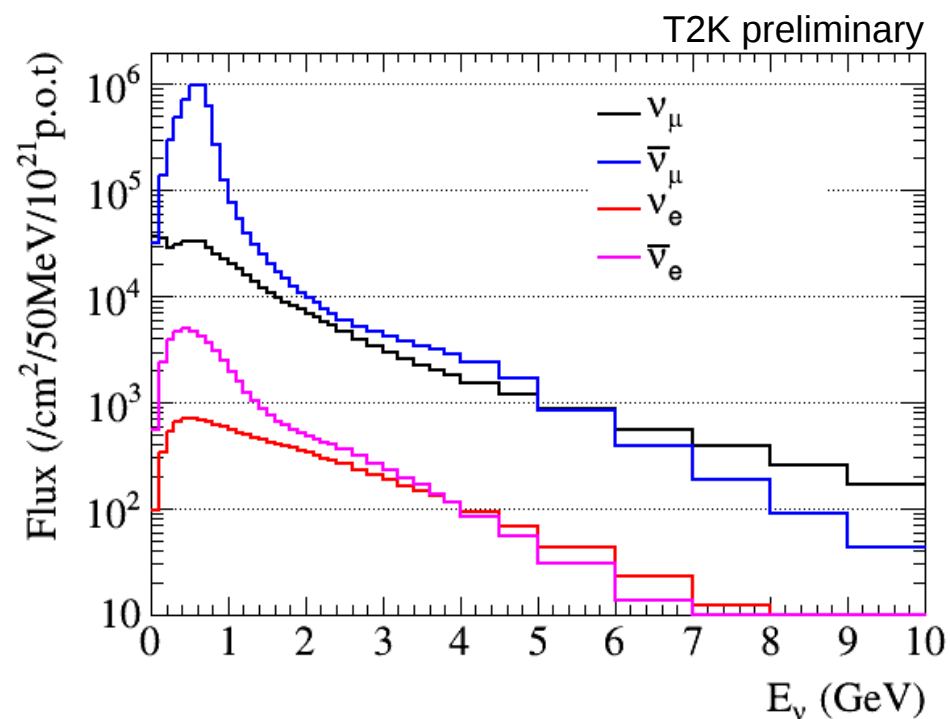
# Analysis description

## Neutrino flux prediction

“Neutrino mode”



“Anti-neutrino mode”



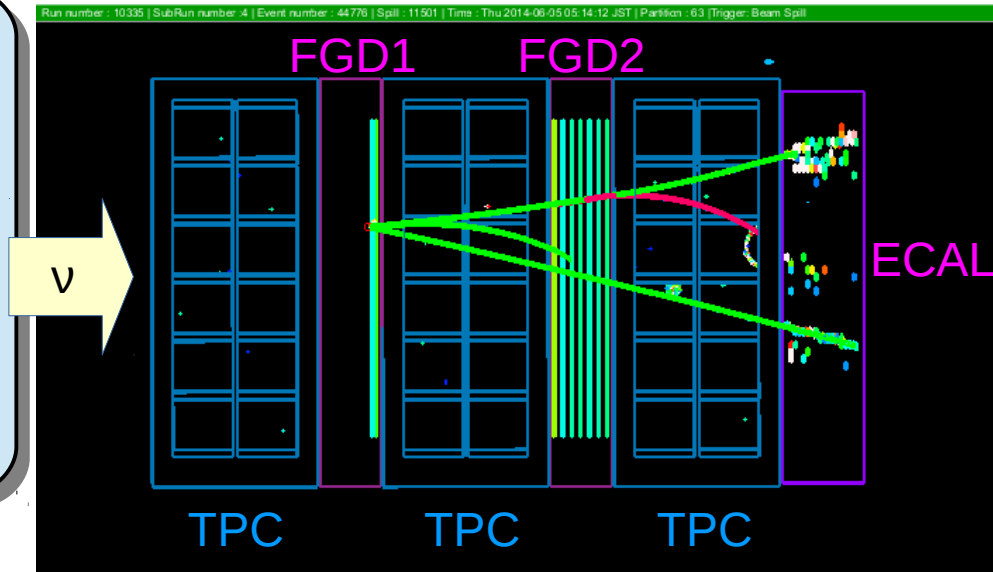
- Intrinsic  $\nu_e/\bar{\nu}_e$  component
- “Wrong sign” component
- Neutrino and anti-neutrino mode fluxes not equivalent (20% less  $\bar{\nu}_\mu$  in  $\bar{\nu}$ -mode than  $\nu_\mu$  in  $\nu$ -mode)

# Near detector analysis

## Event selection – neutrino mode

Select CC  $\nu_\mu$  interactions with vertex in a one of the Fine-Grained Detectors (FGD) Samples separated by FGD:

- FGD1: CH target
- FGD2: 42% water by mass
- Separated by number of tagged pions in each case

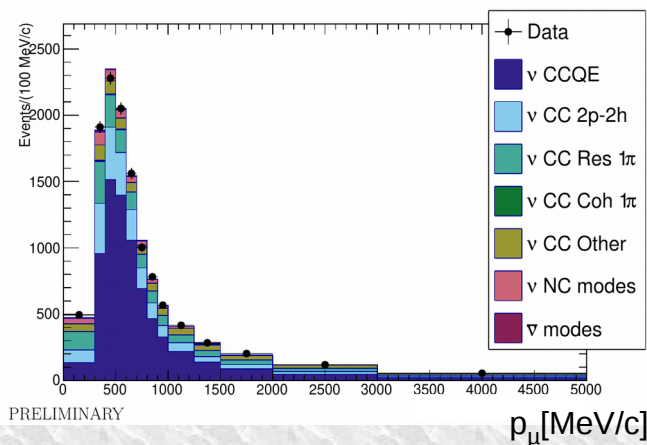


FGD1 samples  
(MC tuned with ND fit)

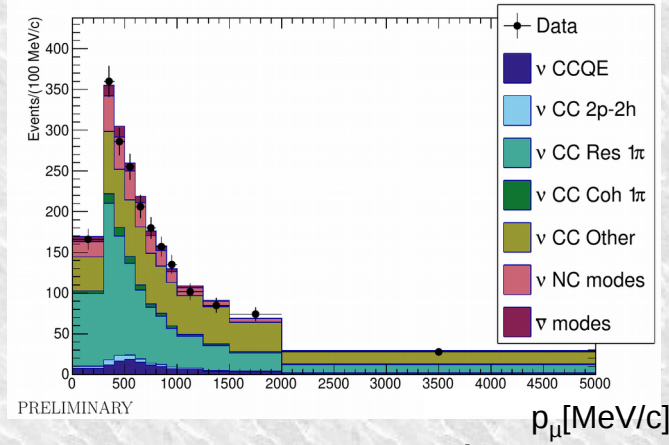
CC0 $\pi$

CC1 $\pi^+$

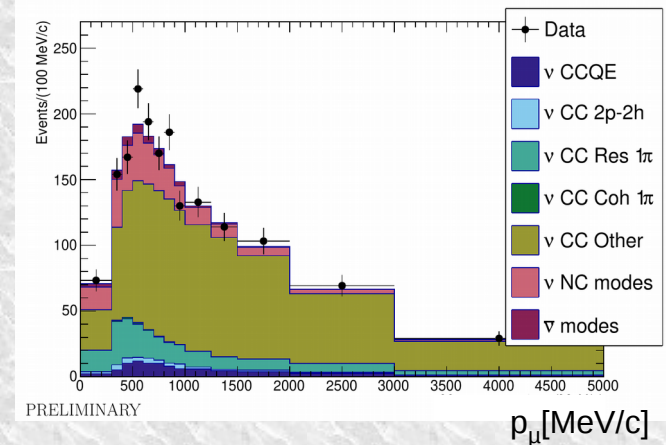
CC other



CCQE enriched



CC resonant pion  
production enriched



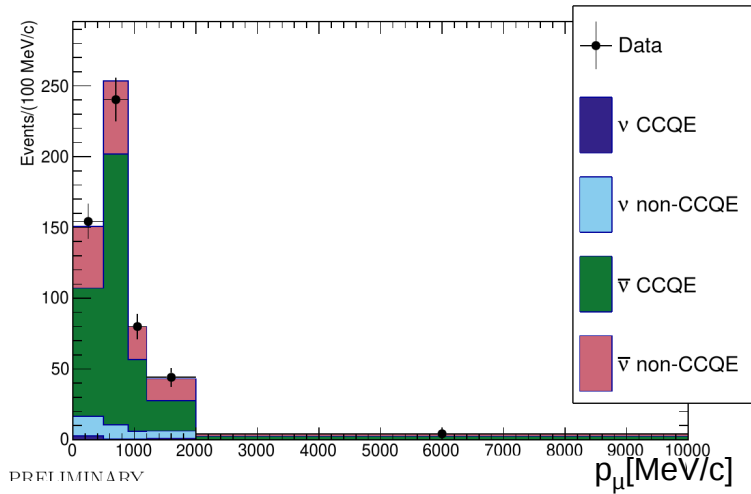
Mainly deep inelastic

# Near detector analysis

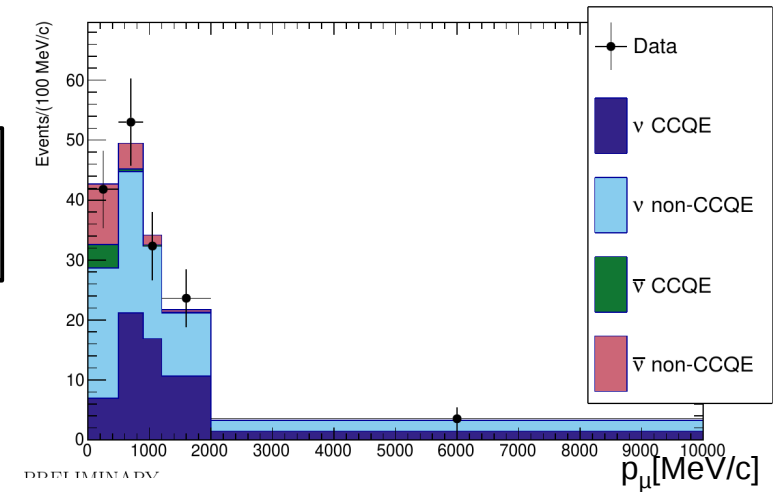
## Event selection – anti-neutrino mode

Large neutrino background in anti-neutrino mode:  
make wrong sign samples to constrain it

### CC $\bar{\nu}_\mu$ samples (FGD1)

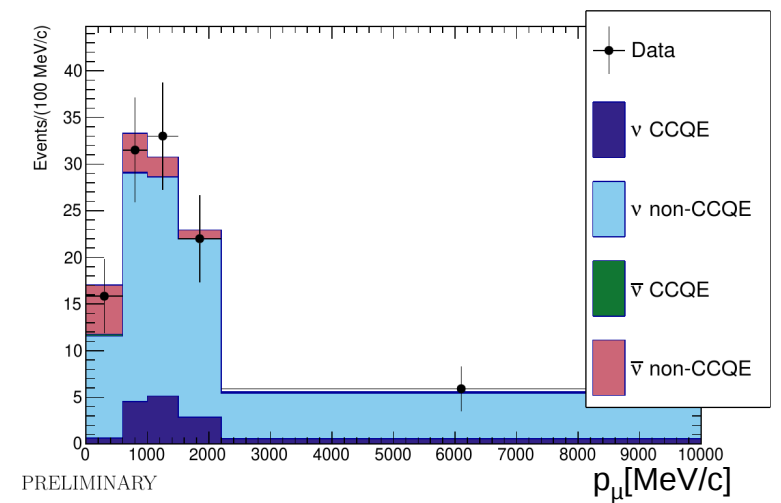
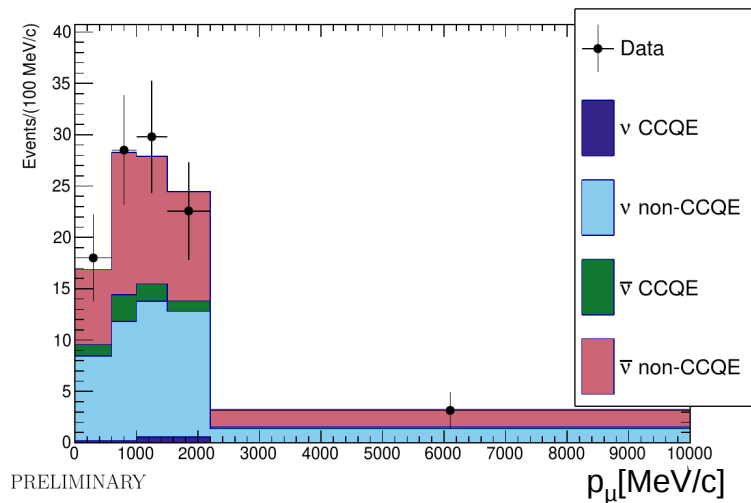


### CC $\nu_\mu$ samples (FGD1)



CC-1track  
(CCQE enriched)

CC-Ntrack  
(CC non-QE  
enriched)



# Far detector

## Energy reconstruction

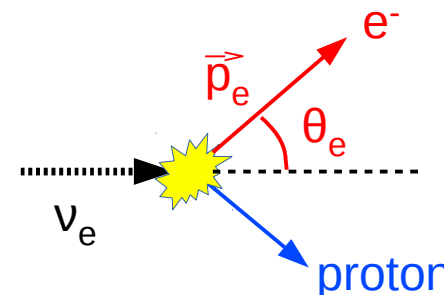
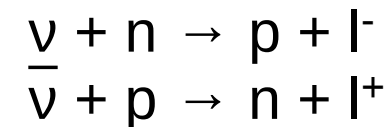
Oscillations depend on  $E_\nu$

$$phase \propto \frac{\Delta m_{ij}^2 L}{E_\nu}$$

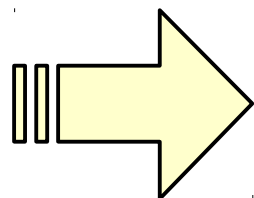
Water Cherenkov detector:

- Only sees charged particles and photons
  - Has a momentum threshold
- See only leptons and pions at T2K energies

CCQE interactions



Knowing  $\nu$  direction, can reconstruct  $E_\nu$  from lepton ( $p, \theta$ )



Build CCQE enriched samples

# Far detector Event selections

## Electron-like samples

- Non  $\nu_\mu \rightarrow \nu_e$  1Re event
- Intrinsic beam  $\nu_e$
  - $\text{NC}\pi^0 \rightarrow 2\gamma$  with missed  $\gamma$

Main  
backgrounds

## Muon-like samples

- Non CCQE 1R $\mu$  events
- MEC/2p2h
  - CCRes/DIS with invisible pions (FSI, below threshold)

Cut	Description
Fully Contained FV	Event on timing in fiducial volume
1 ring only	Only one charged particle for CCQE events
PID	Charged particle should be a $e^-/e^+$
$E_{\text{vis}} > 100$ MeV	Rejects low energy background (NC and invisible muons)
No decay $e^-$	Rejects events with pion/muon below threshold
$E_{\text{rec}} < 1.25$ GeV	Reject intrinsic beam $\nu_e$
“fiTQun” $\pi^0$ cut	Rejects $\text{NC}\pi^0$ events

Cut	Description
Fully Contained Fiducial Volume	Event on timing in fiducial volume
1 ring only	Only one charged particle for CCQE events
PID	Charged particle should be a $\mu^-/\mu^+$
$p_\mu > 200$ MeV/c	
# decay $e^- \leq 1$	Rejects events with pions below threshold



- Maximum likelihood methods to measure the PMNS parameters
- Marginalize (integrate) over the nuisance parameters
- Bayesian and frequentist results

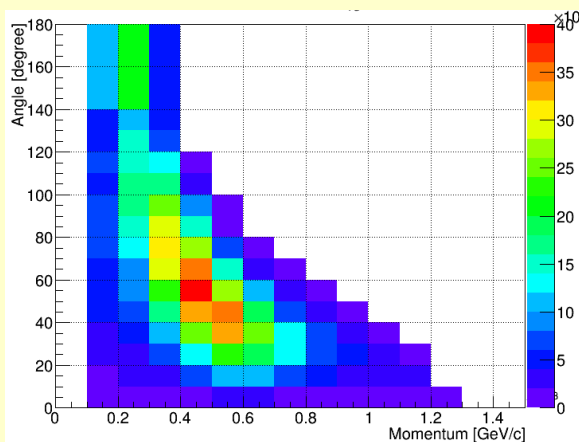
3 different analyses giving consistent results

Different use of near detector data:  
→ 1 joint near/far analysis  
→ 2 use result of ND fit as input

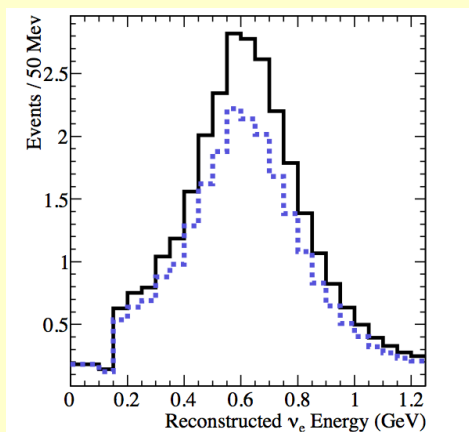
Different fitting methods:  
→ 2 “grid searches”  
→ 1 uses MCMC

## Different ‘shape’ information for e-like samples

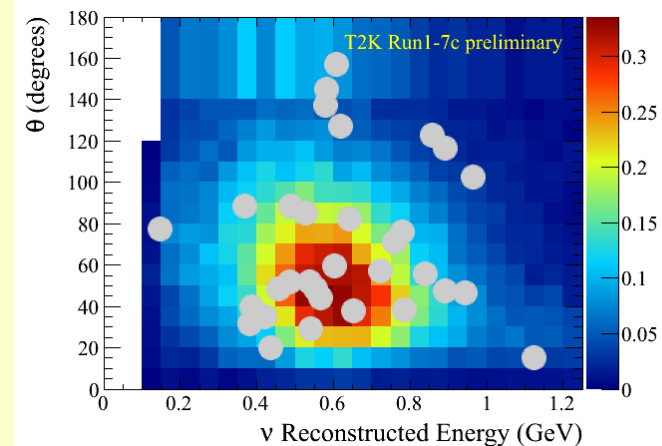
Lepton ( $p, \theta$ )



Neutrino  $E_{\text{rec}}$



$\nu E_{\text{rec}} + \text{lepton } \theta$

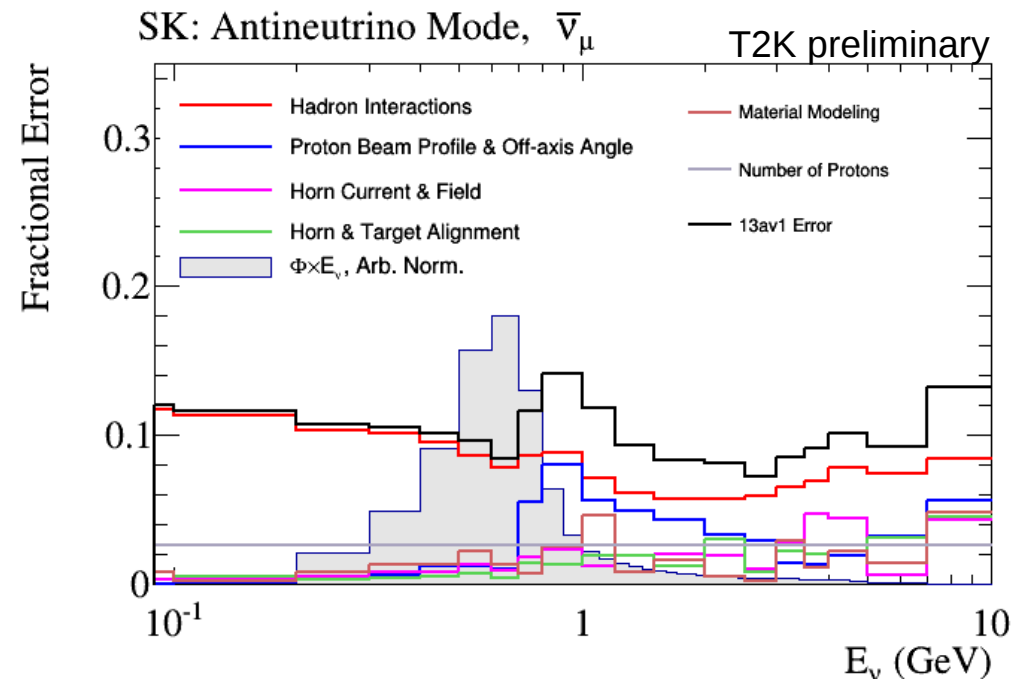
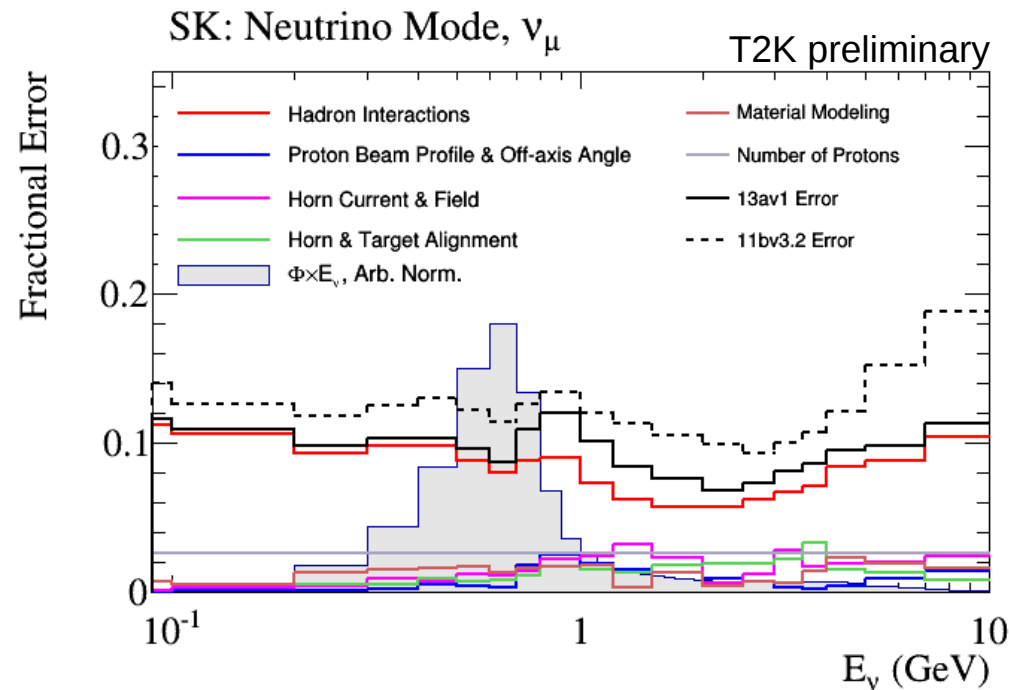




# **Systematic uncertainties**

# Systematic uncertainties Neutrino flux

- Several sources of systematic uncertainties considered : beamline alignment, hadron production, horn current, proton beam parameters...
- Energy-dependent uncertainty for each neutrino flavor
- ~10% uncertainty at peak energy
- Dominant contribution: uncertainty on hadron interactions in target

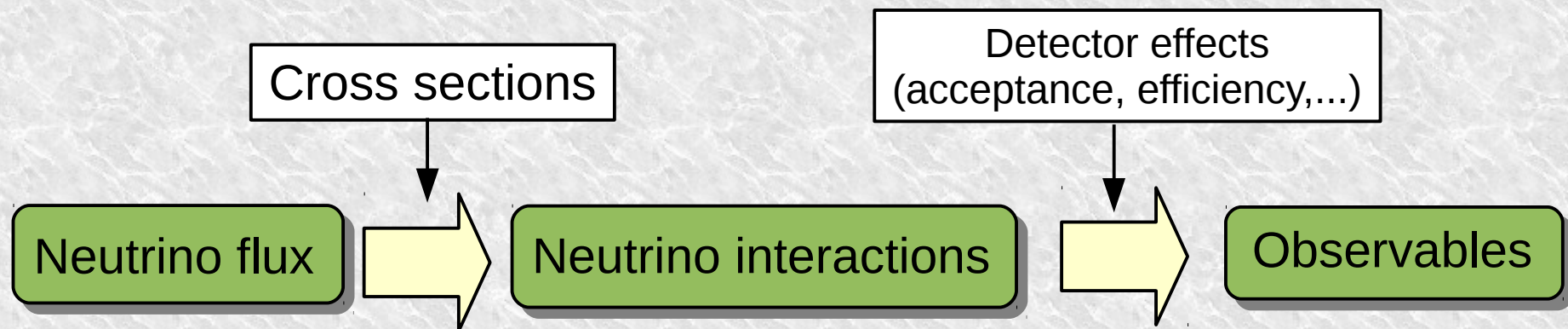


# Systematic uncertainties

## Near to far extrapolation

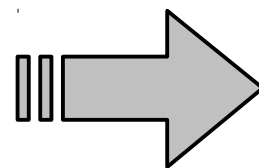
Detectors measure rate as a function of a reconstructed quantity from observables  
 e.g: reconstructed neutrino energy from lepton ( $p, \theta$ )

Want to compare flux between far and near detectors, but have only access to those observables/reconstructed quantities



Differences between ND and FD:

- different fluxes (oscillations)
- different target material
- different acceptance
- different detector technologies

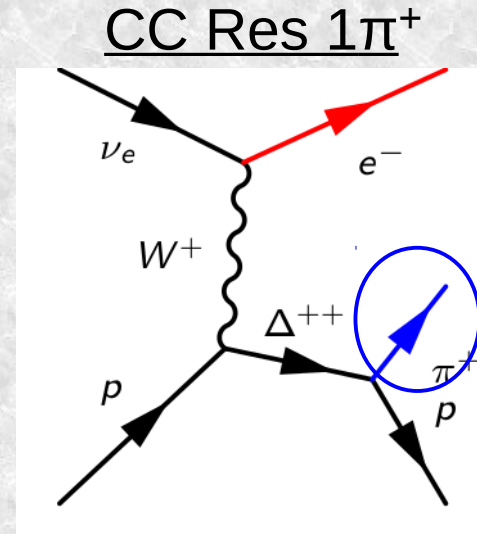
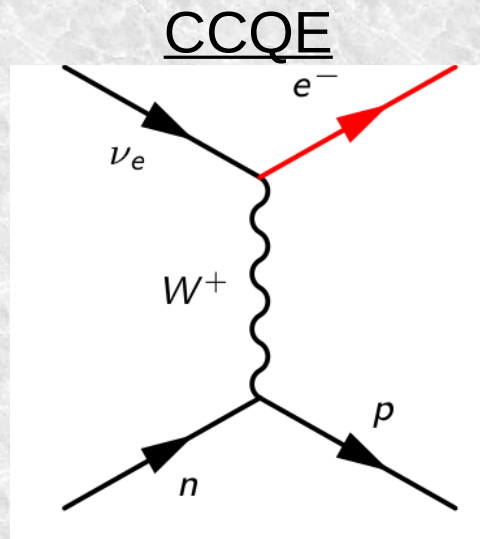


Use models for extrapolation

# Systematic uncertainties

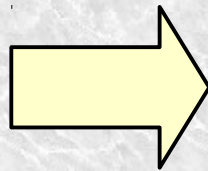
## Neutrino interactions – why it matters

Different relations between neutrino energy and observables in detector for the different types of interactions



can be absorbed  
in nucleus

Different true  
energies

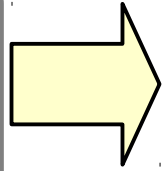


Different effect of oscillations  
at far detector

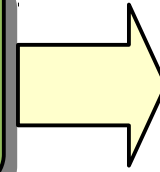
Same observables, but different near to far extrapolation

# Systematic uncertainties Neutrino interactions

Different fluxes at near and far detectors (oscillations)



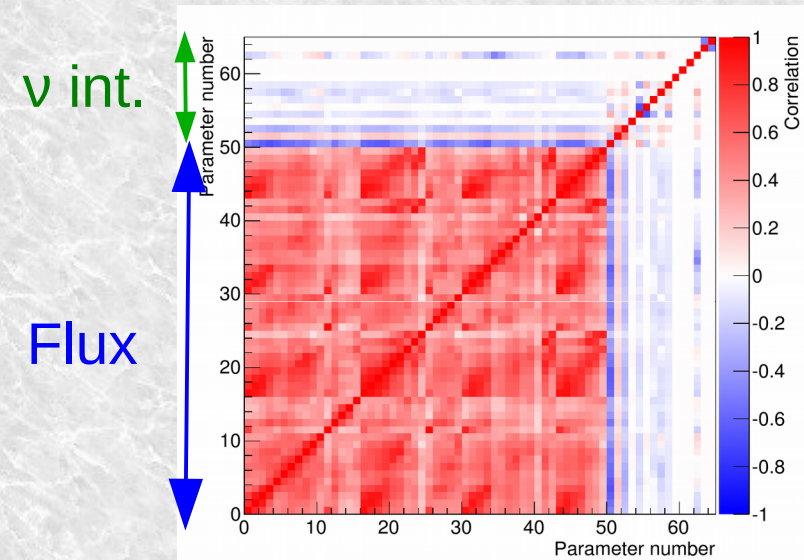
Different fraction of each interaction at ND and FD



Need uncertainties on rate and properties of each interaction type

- Select interaction models using external data
- Nominal predictions from NEUT
- Uncertainties on model parameters ( $M_A$ ,  $pF$ , ...)
- Additional normalization uncertainties for certain modes

Interaction uncertainties fitted in ND with flux uncertainties



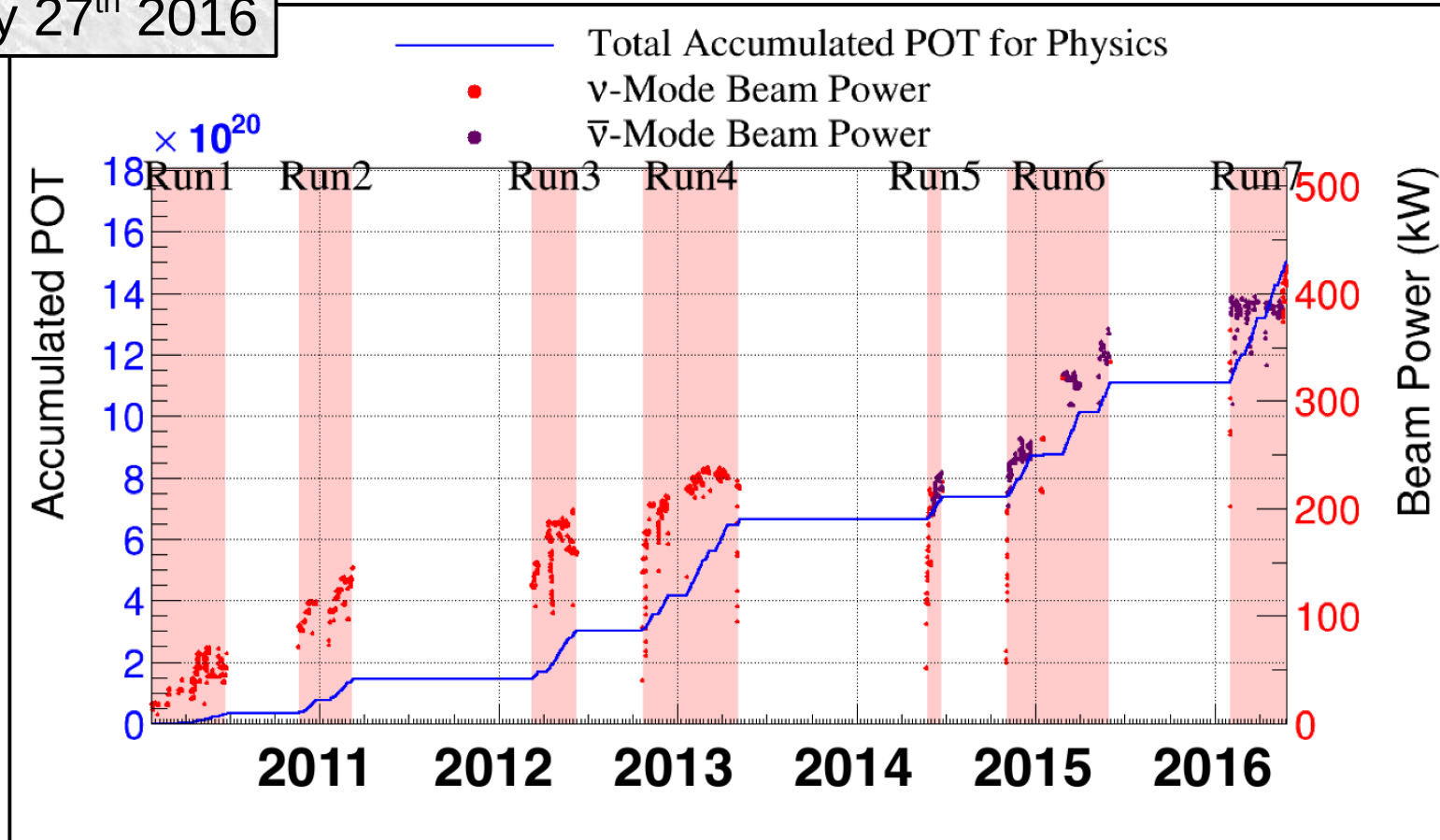
Result applied to FD prediction

# Dataset

# Dataset

## Run 1-7 data

Using data taken  
until May 27<sup>th</sup> 2016



### Near detector analysis

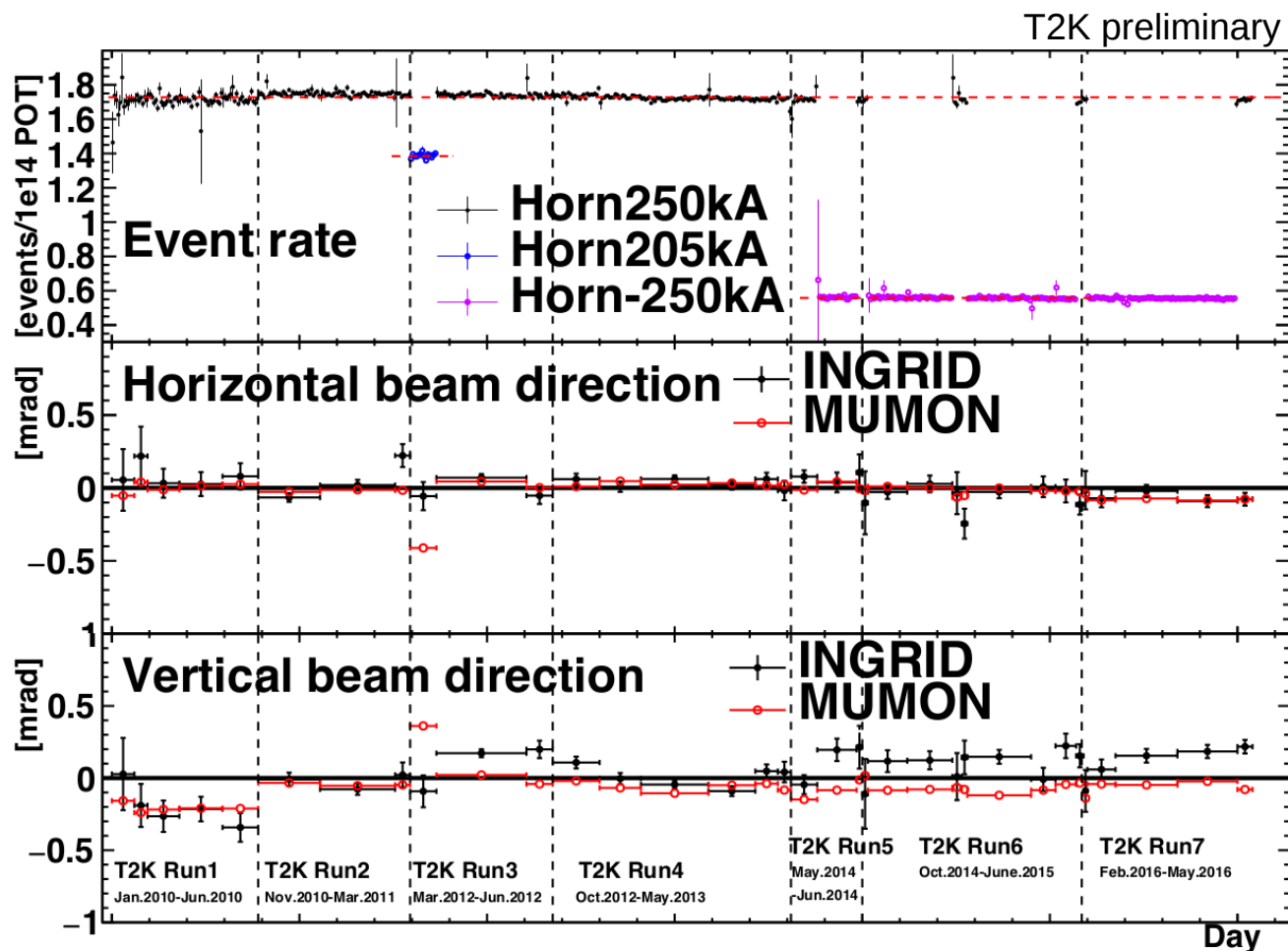
ν-mode:  $5.82 \times 10^{20}$  POT  
 ν̄-mode:  $2.84 \times 10^{20}$  POT

### Far detector analysis

ν-mode:  $7.482 \times 10^{20}$  POT  
 ν̄-mode:  $7.471 \times 10^{20}$  POT

# Beam stability

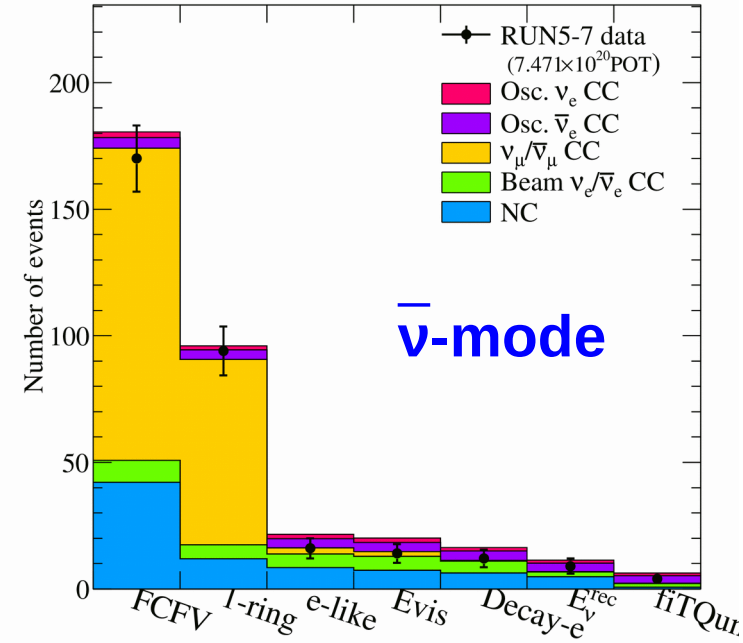
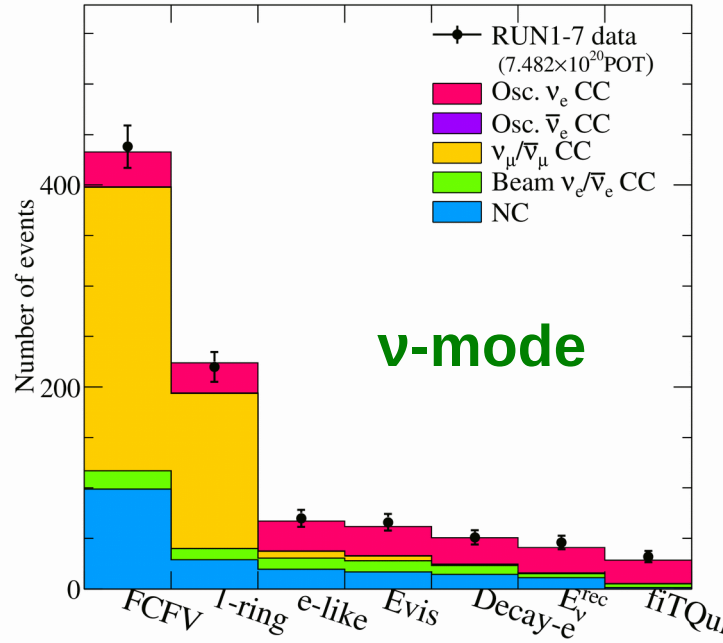
Stable event rate and beam direction from muon monitor and on-axis near detector measurements



Off-axis angle controlled better than 1 mrad target uncertainty  
(= 2% uncertainty on peak energy at SK)



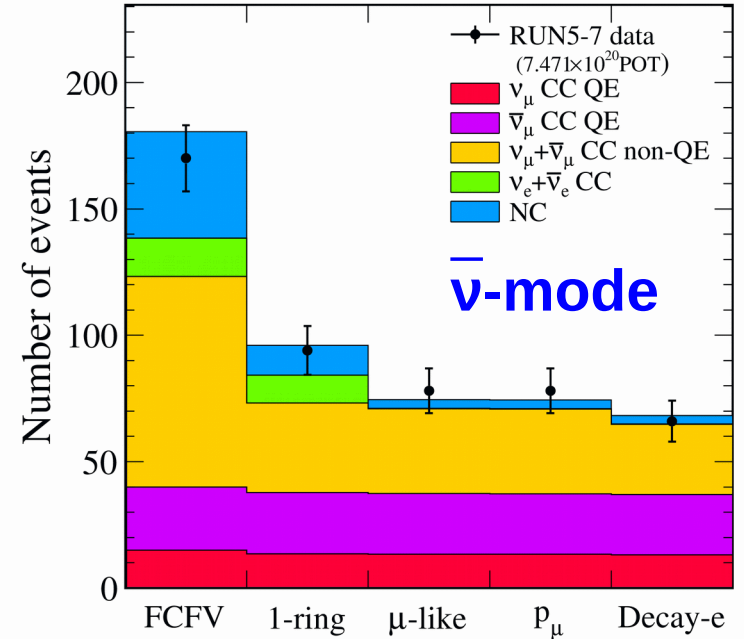
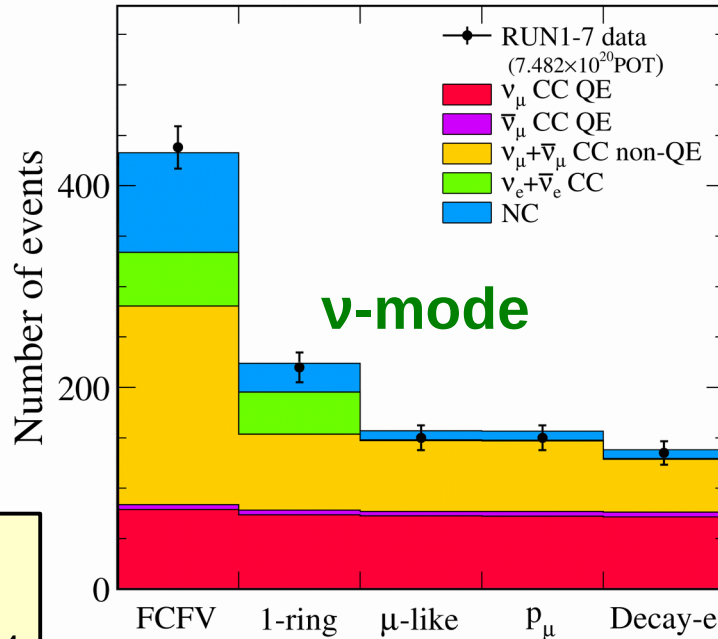
# Far detector data Electron-like samples



$\sin^2(\theta_{23})=0.528$   
 $\Delta m^2_{32}=2.509 \times 10^{-3} \text{ eV}^2 \text{ c}^{-4}$   
 $\sin^2(\theta_{13})=0.0217$

Sample	Mass hierarchy	$\delta=0$ MC	$\delta=\pi$ MC	$\delta=-\pi/2$ MC	$\delta=\pi/2$ MC	Observed
Neutrino mode	Normal	24.2	24.1	28.7	19.6	<b>32</b>
	Inverted	21.3	21.3	25.4	17.1	
Antineutrino mode	Normal	6.9	6.8	6.0	7.7	<b>4</b>
	Inverted	7.4	7.4	6.5	8.4	

# Far detector data Muon-like samples



$\delta = -1.601$   
 $\sin^2(\theta_{23}) = 0.528$   
 $\Delta m^2_{32} = 2.509 \times 10^{-3} \text{ eV}^2 \text{ c}^{-4}$   
 $\sin^2(\theta_{13}) = 0.0217$

Sample	Oscillated MC	No oscillations MC	Observed
Neutrino mode	135.8	521.8	135
Anti-neutrino mode	64.2	184.8	66

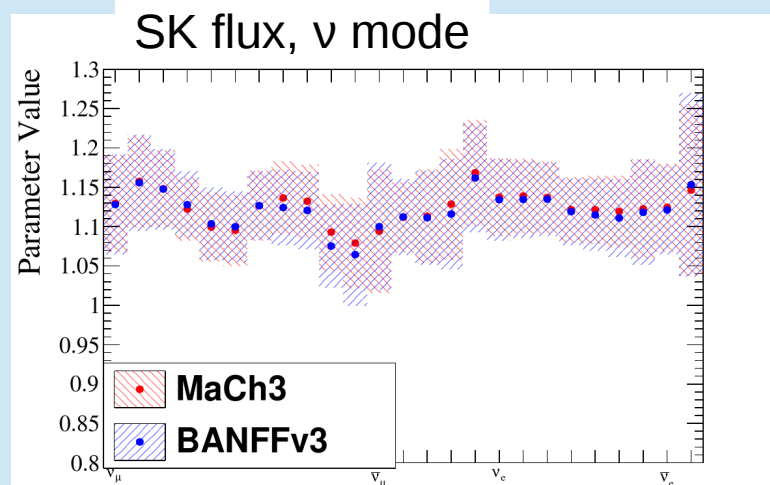
# Results

# Near detector analysis

## Fit results

### 2 different fits:

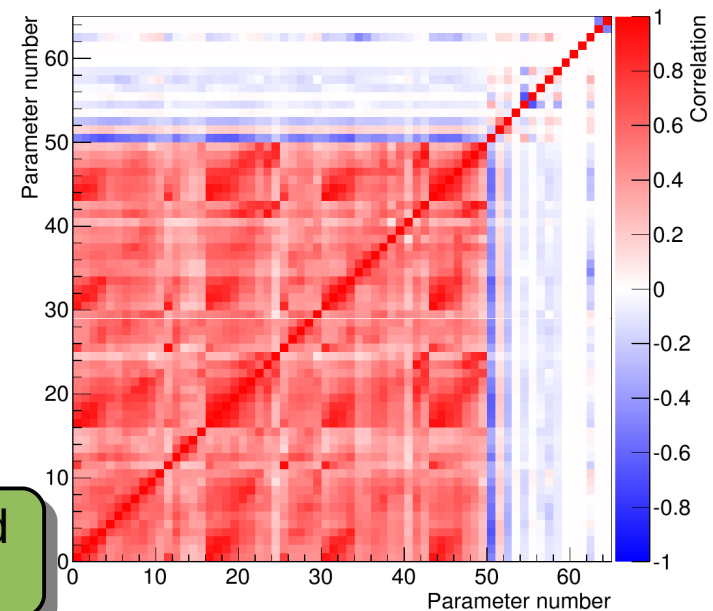
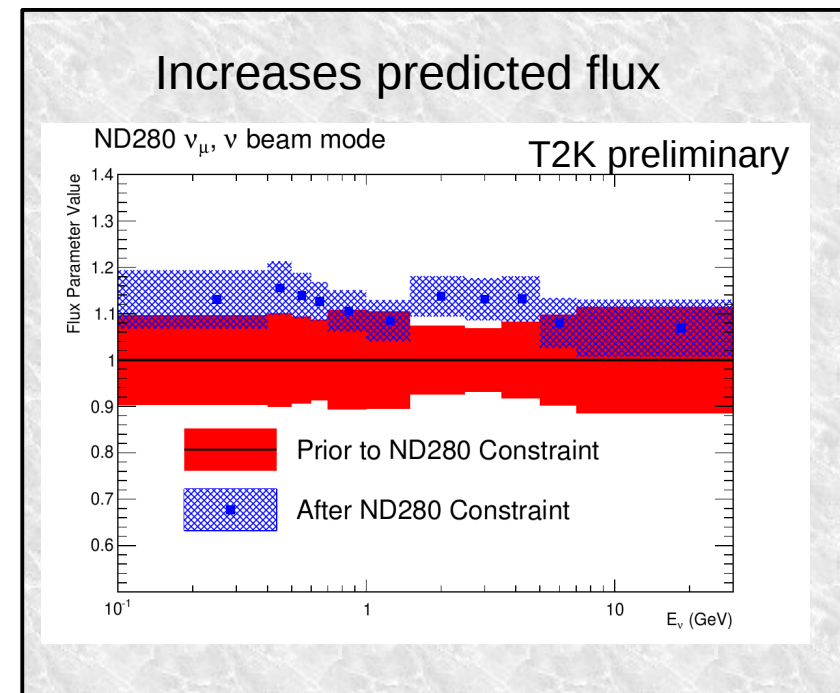
- Minuit minimization “BANFF”
- MCMC “Mach3”



Consistent results

Goodness of fit: **p-value=0.086**

Creates anti-correlations between flux and interaction systematics

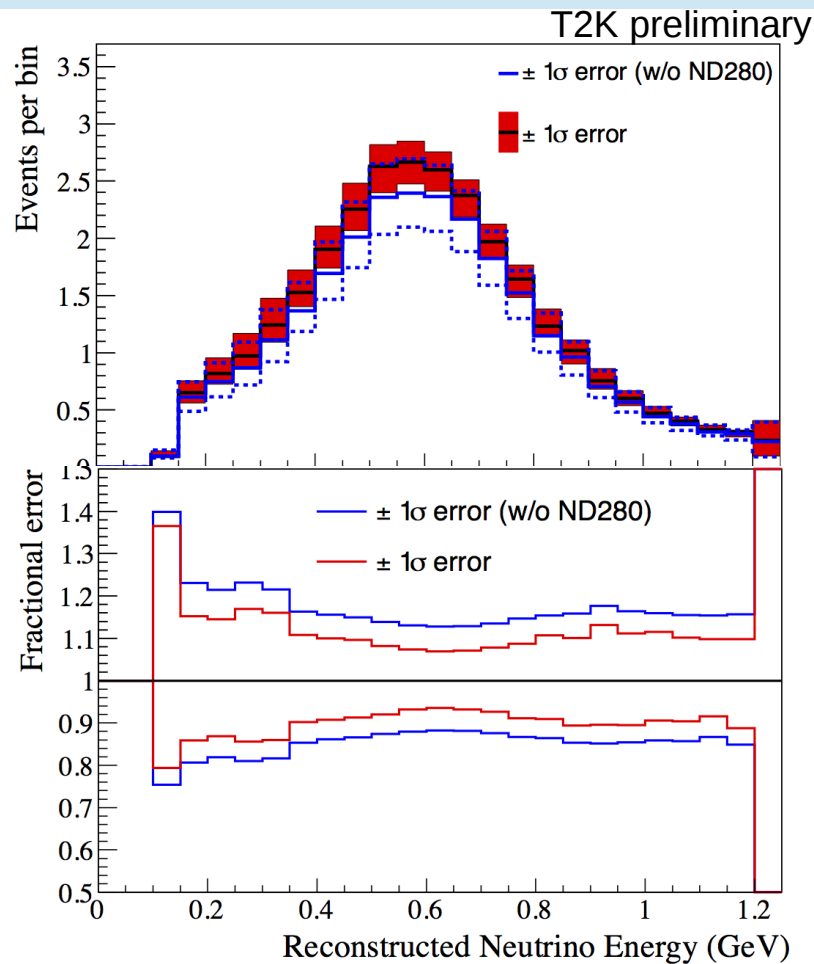


# Near detector analysis

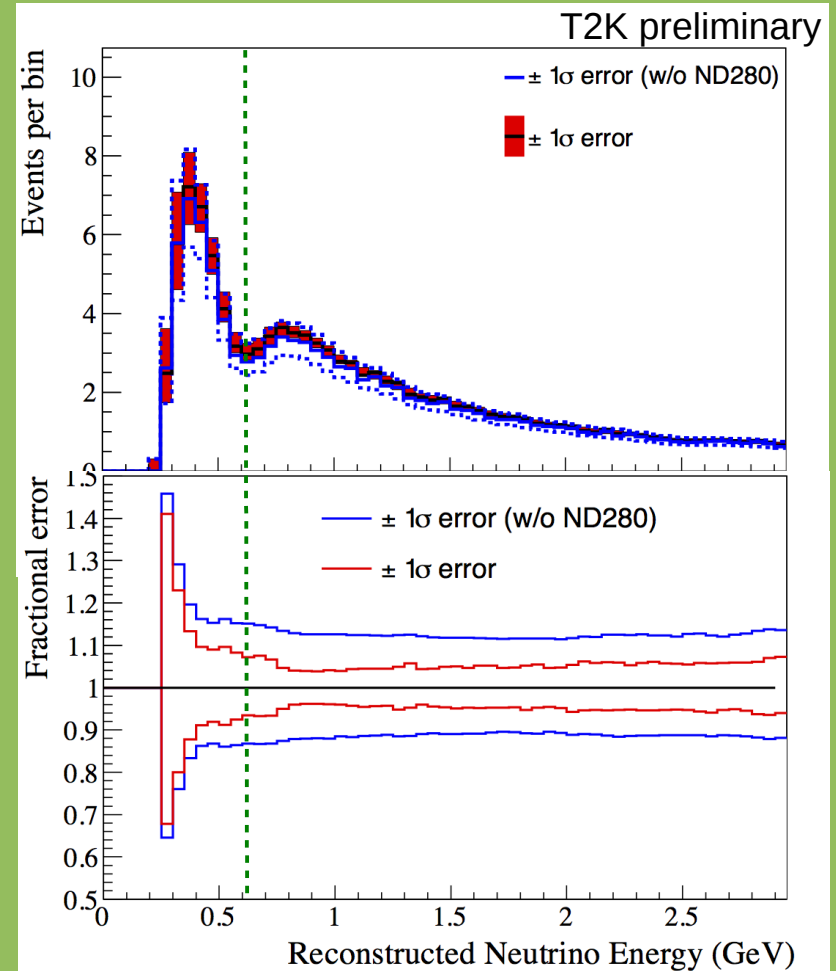
## Systematic uncertainty reduction

Both changes the nominal rate predictions and reduces the uncertainties

### Neutrino mode electron-like



### Neutrino mode muon-like



$$(\delta = -1.601, \sin^2(\theta_{23}) = 0.528, \Delta m_{32}^2 = 2.509 \times 10^{-3} \text{ eV}^2 \text{ c}^{-4}, \sin^2(\theta_{13}) = 0.0217)$$

# Near detector analysis

## Systematic uncertainty reduction

Significantly reduces uncertainty on expected number of events at SK

	$\nu_e$ sample	$\nu_\mu$ sample	$\bar{\nu}_e$ sample	$\bar{\nu}_\mu$ sample
Flux + Xsec (w/o ND fit)	11.4 %	10.9 %	12.8 %	11.6 %
Flux + Xsec (with ND fit)	4.1 %	2.8 %	4.6 %	3.2 %
Far detector (after ND fit)	3.6 %	4.1 %	3.7 %	3.9 %
Total (w/o ND fit)	12.1 %	12.0 %	13.4 %	12.5 %
<b>Total (with ND fit)</b>	<b>5.1 %</b>	<b>5.0 %</b>	<b>6.0 %</b>	<b>5.0 %</b>

Initially:

➤  $\nu_e$  appearance  $\rightarrow \theta_{13}, \delta$

➤  $\nu_\mu$  disappearance  $\rightarrow \theta_{23}, |\Delta m^2|$

But observables depend on all 4 parameters:

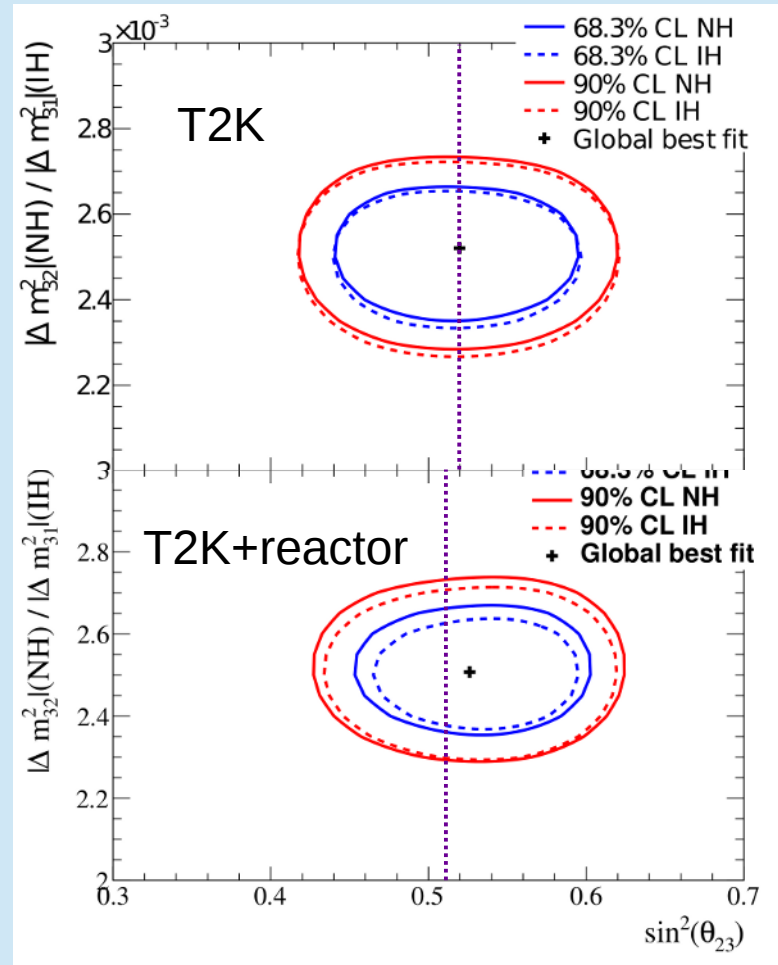
$$P(\nu_\mu \rightarrow \nu_e) \sim 2 \sin^2(2\theta_{13}) \sin^2(\theta_{23}) \sin^2(\Delta m_{31}^2 L/4E)$$

$$P(\nu_\mu \rightarrow \nu_\mu) \sim 1 - \cos^4(\theta_{13}) \sin^2(2\theta_{23}) \sin^2(\Delta m_{31}^2 L/4E) - \sin^2(2\theta_{13}) \sin^2(\theta_{23}) \sin^2(\Delta m_{31}^2 L/4E)$$

Use all samples for most precise measurement of PMNS parameters:

- Electron-like neutrino mode
- Muon-like neutrino mode
- Electron-like anti-neutrino mode
- Muon-like anti-neutrino mode

## 2014 joint 1Re/1R $\mu$ analysis



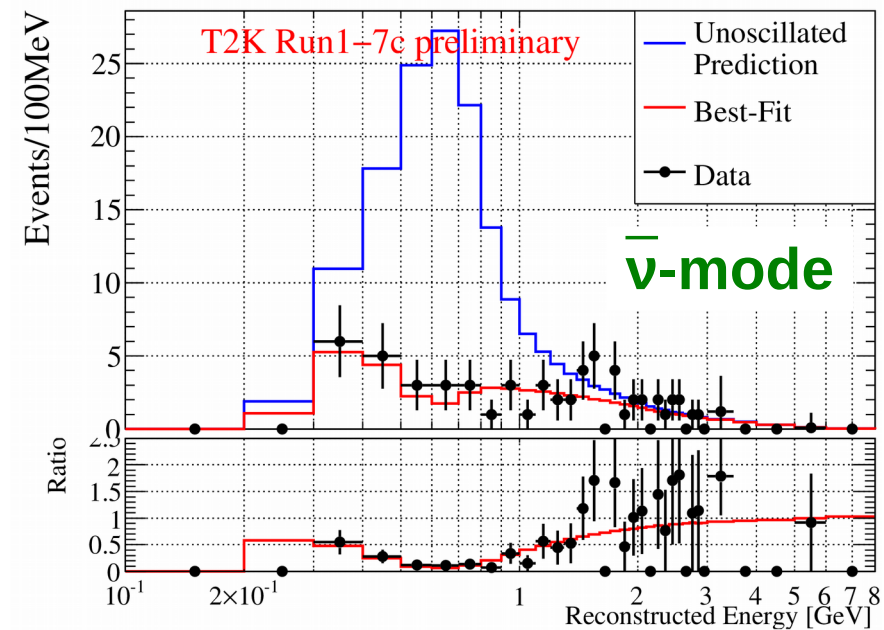
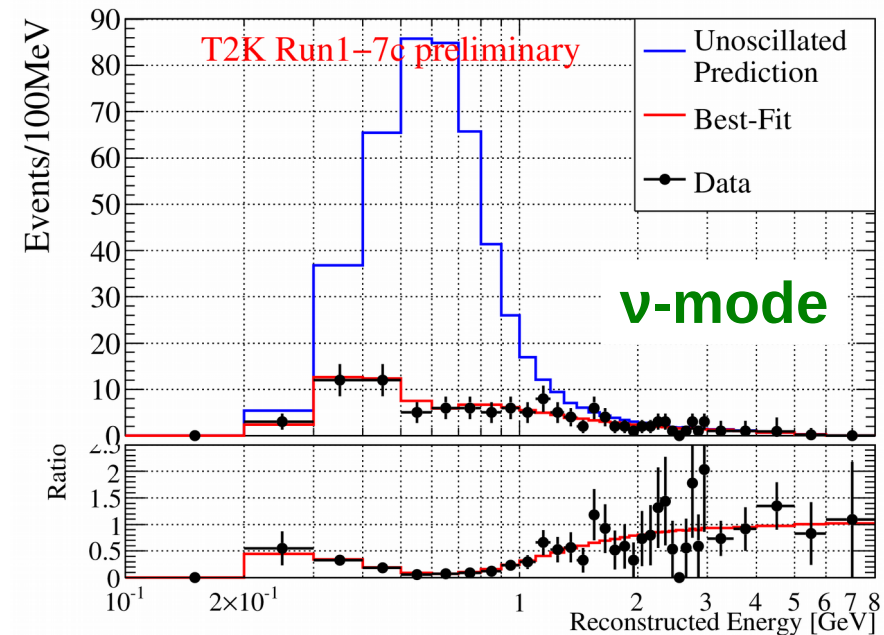
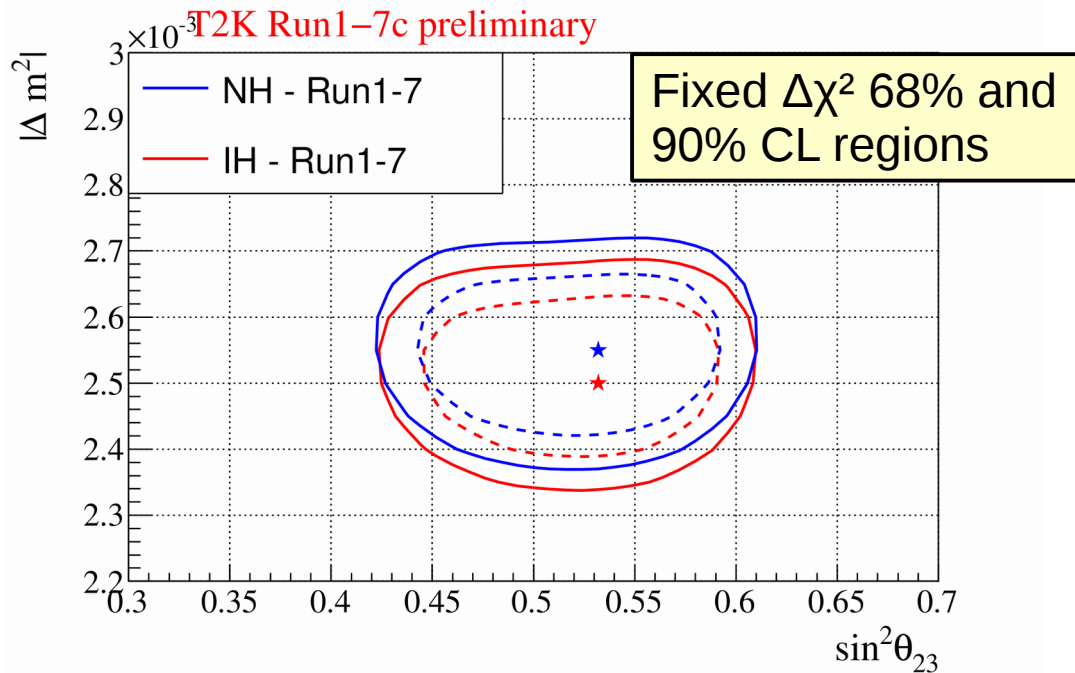
1Re sample also influences measurement of  $\theta_{23}$  and  $\Delta m_{32}^2$



# Combined $\nu\bar{\nu}$ analysis

## Atmospheric parameters

Reactor constraint (PDG2015)  
 $\sin^2(2\theta_{13}) = 0.085 \pm 0.005$

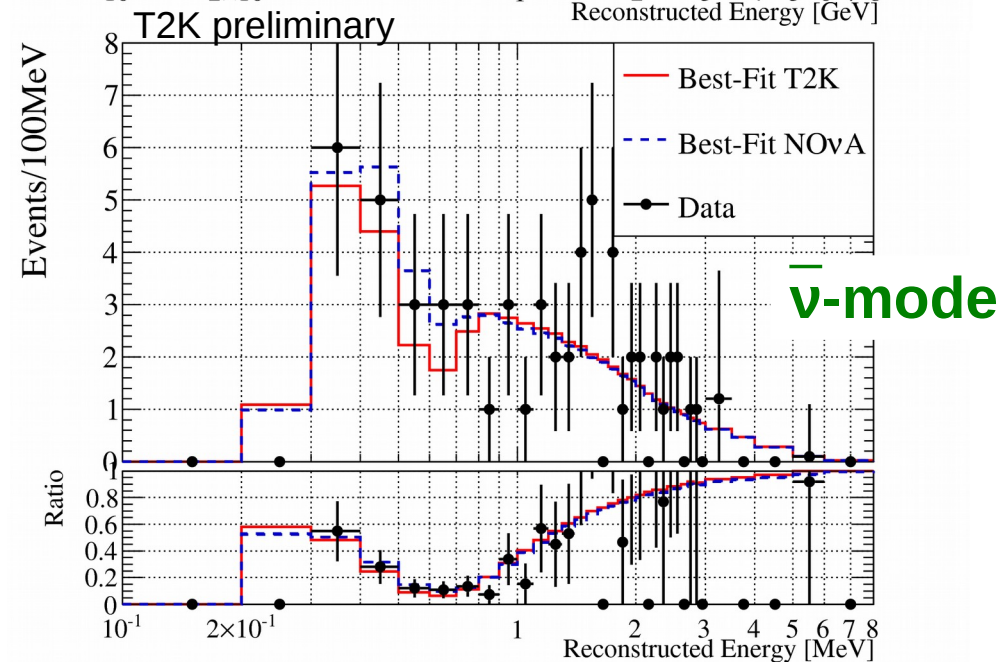
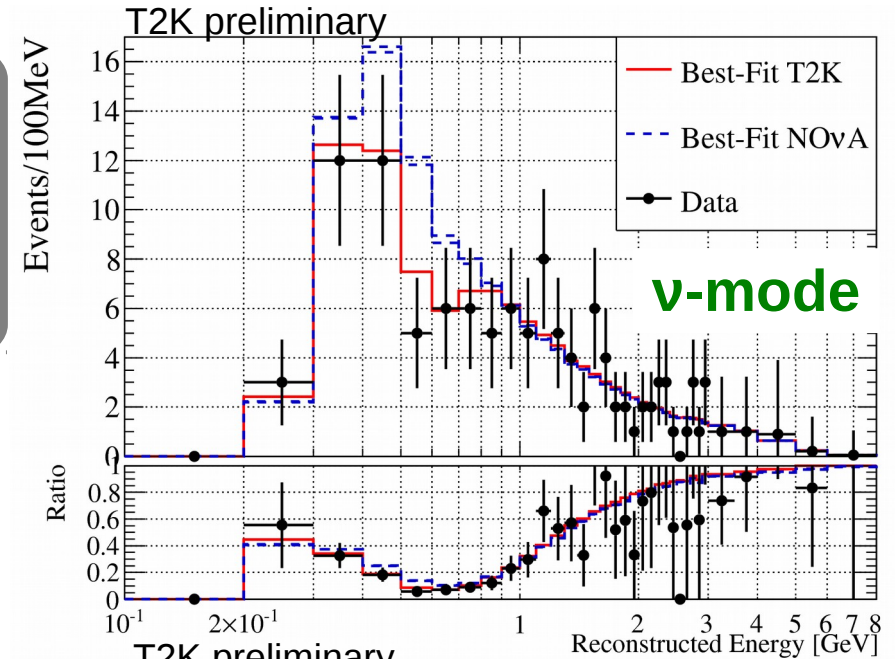
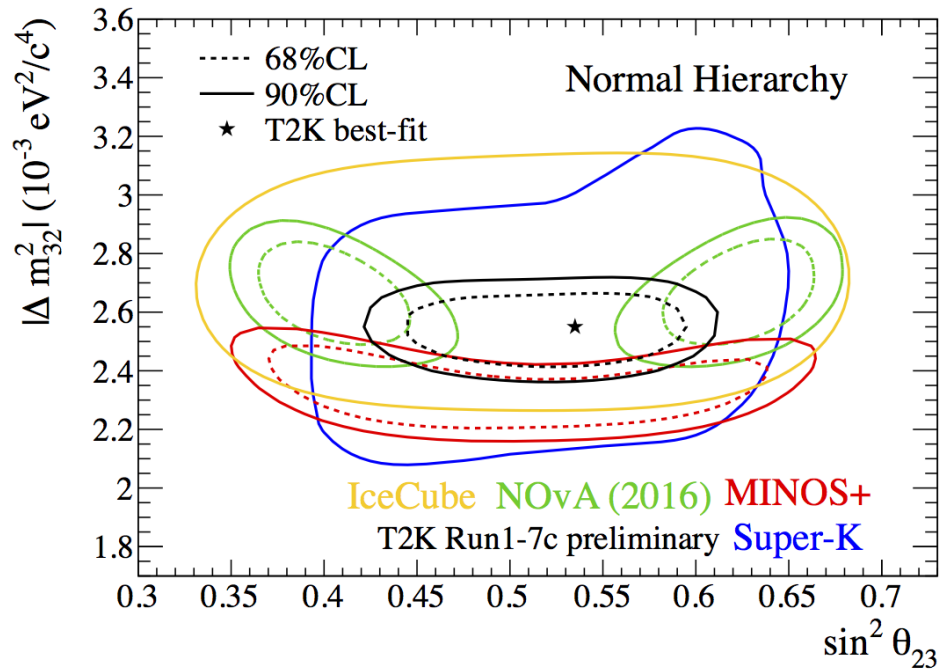


Parameter	Normal hierarchy	Inverted hierarchy
$\sin^2(\theta_{23})$	$0.532^{+0.046}_{-0.068}$	$0.534^{+0.043}_{-0.066}$
$ \Delta m^2_{32} $ ( $10^{-3}\text{eV}^2/\text{c}^4$ )	$2.545^{+0.081}_{-0.084}$	$2.510^{+0.081}_{-0.083}$

# Combined $\nu$ - $\bar{\nu}$ analysis

## Atmospheric parameters

- Measurements compatible with other experiments results
- T2K and NOvA slightly favour different values of the parameters

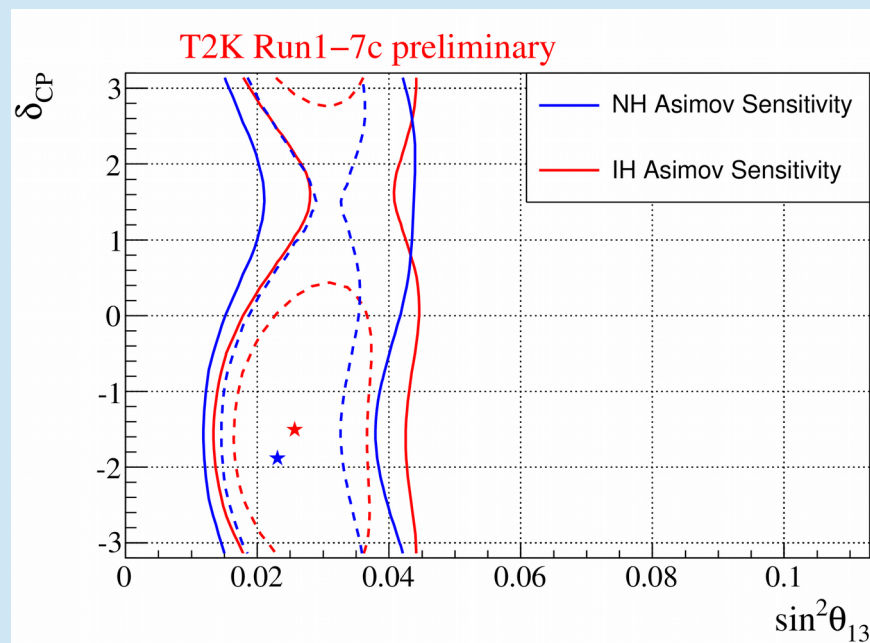


# Combined $\nu$ - $\bar{\nu}$ analysis $\theta_{13}$ and $\delta$ – T2K only

- Compare  $\theta_{13}$  measurement from accelerator and reactor experiments
- Measure  $\delta$  by comparing  $\nu_{\mu} \rightarrow \nu_e$  and  $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e$

- Favor  $\delta \sim -\pi/2$  with T2K data alone
- Compatible with reactor  $\theta_{13}$  measurement

## Sensitivity



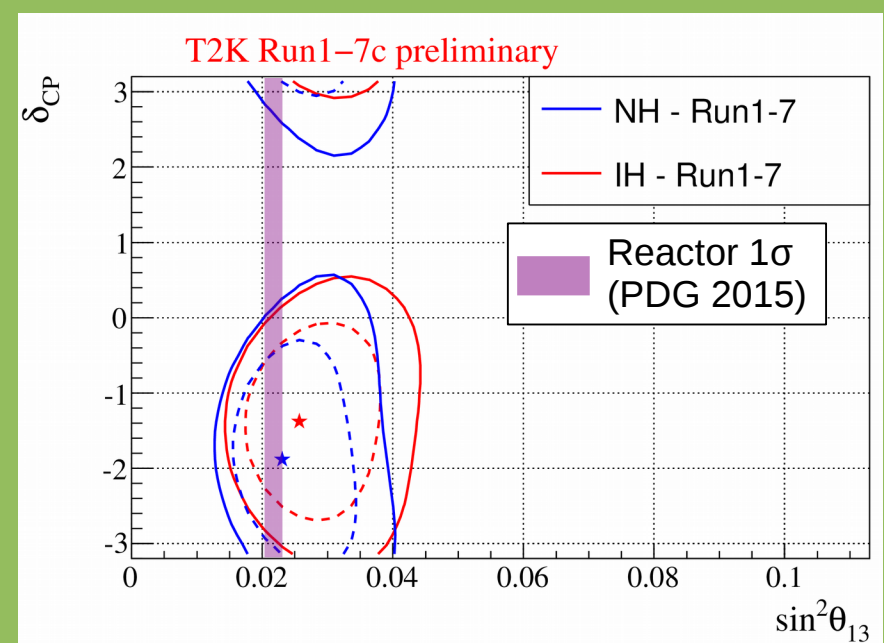
$$\delta = -1.601$$

$$\sin^2(\theta_{23}) = 0.528$$

$$\Delta m_{32}^2 = 2.509 \times 10^{-3} \text{ eV}^2 \text{ c}^{-4}$$

$$\sin^2(\theta_{13}) = 0.0217$$

## Data fit



Fixed  $\Delta\chi^2$  68% and 90% CL regions

# Combined $\nu\bar{\nu}$ analysis

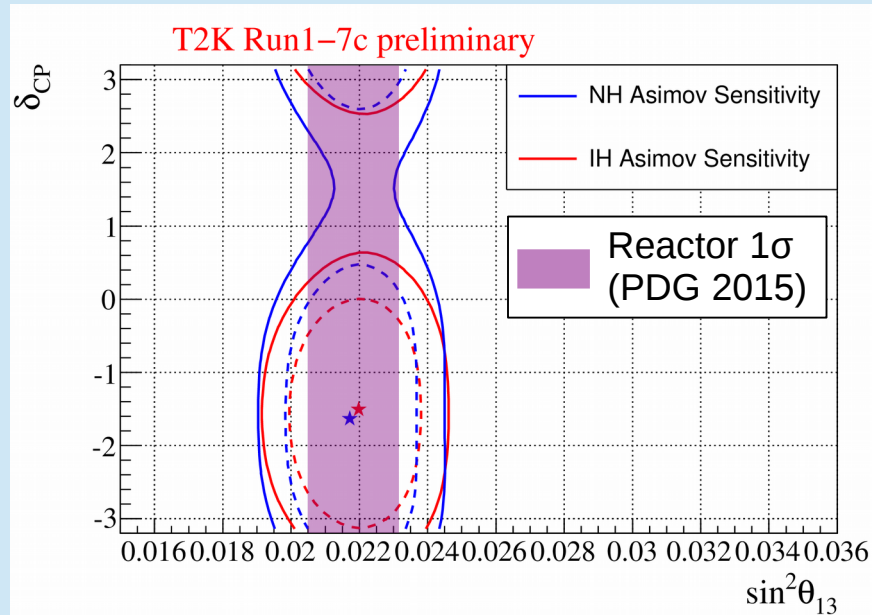
## $\theta_{13}$ and $\delta$ – T2K + reactor

Reactor constraint (PDG2015)  
 $\sin^2(2\theta_{13}) = 0.085 \pm 0.005$

Sensitivity

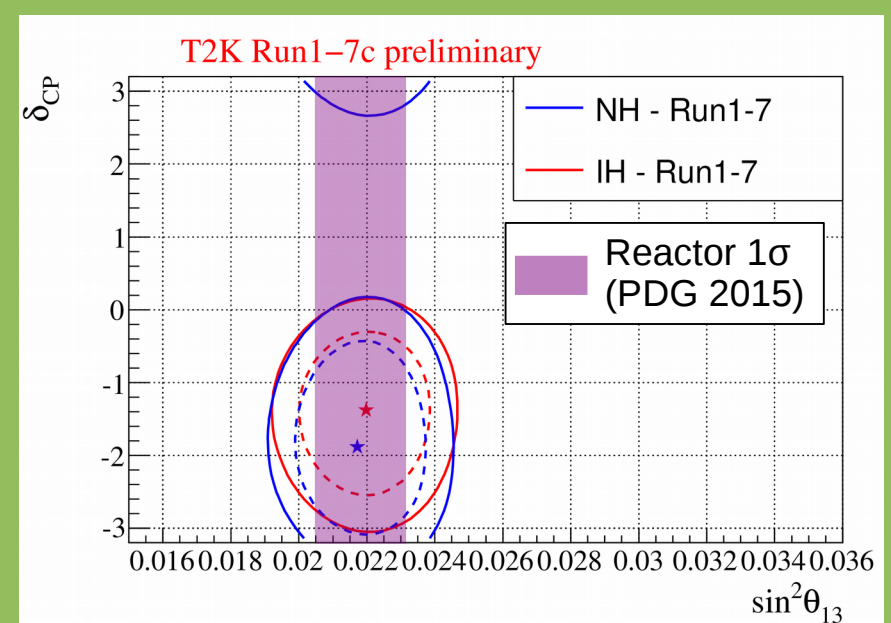
Fixed  $\Delta\chi^2$  68% and  
 90% CL regions

Data fit



$\delta = -1.601$   
 $\sin^2(\theta_{23}) = 0.528$

$\Delta m_{32}^2 = 2.509 \cdot 10^{-3} \text{ eV}^2 \text{ c}^{-4}$   
 $\sin^2(\theta_{13}) = 0.0217$



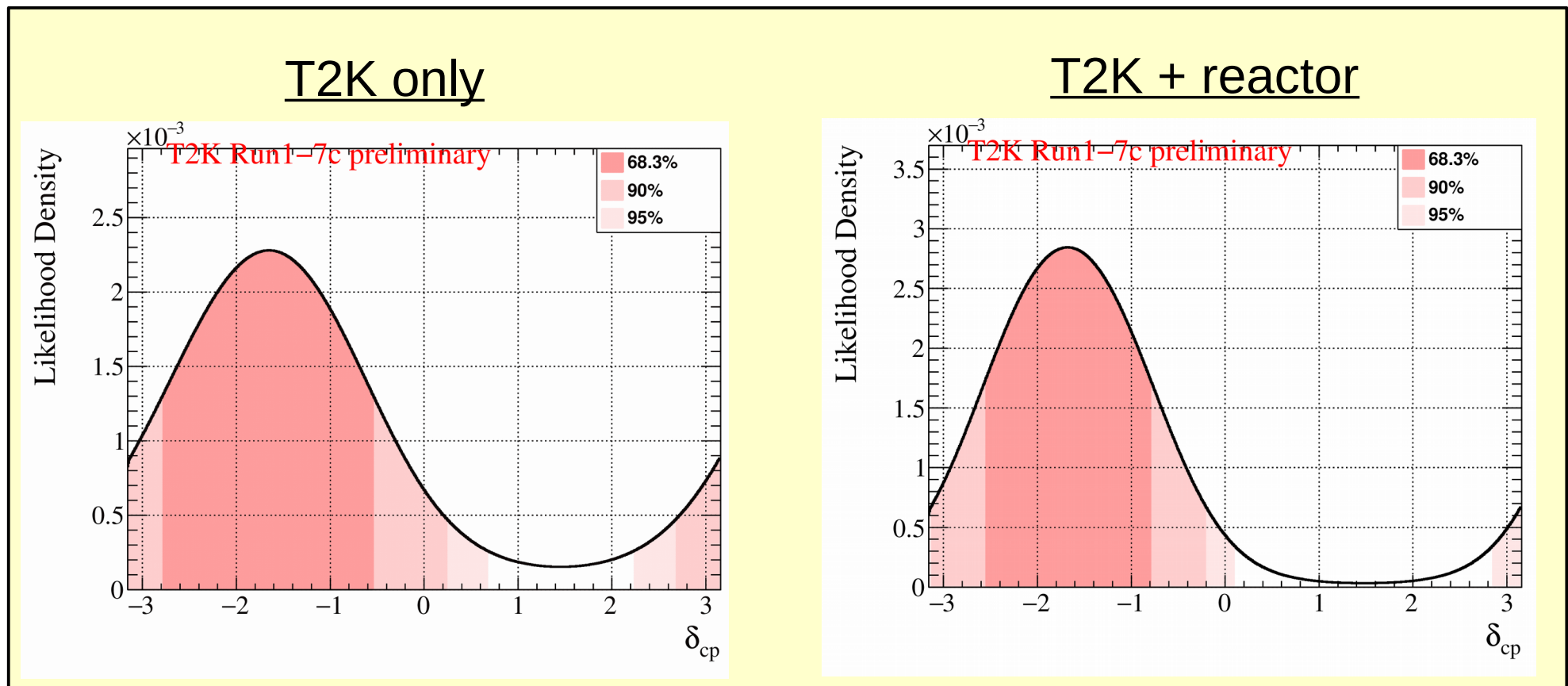


# Combined $\nu$ - $\bar{\nu}$ analysis

## $\delta$ – Bayesian results

Reactor constraint (PDG2015)  
 $\sin^2(2\theta_{13}) = 0.085 \pm 0.005$

Credible intervals marginalizing over everything including mass hierarchy



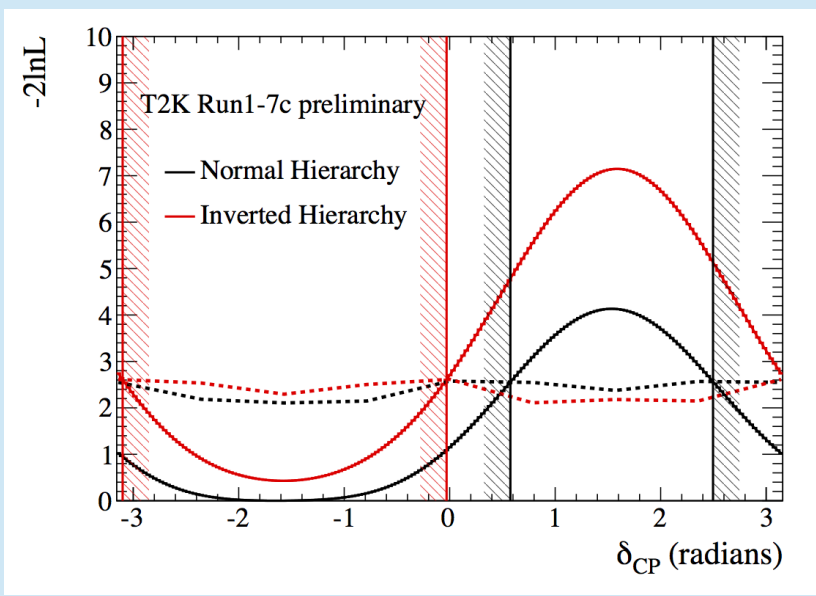
Consistent picture with and without using results of reactor experiments

# Combined $\nu$ - $\bar{\nu}$ analysis $\delta$ – Frequentist results

Reactor constraint (PDG2015)  
 $\sin^2(2\theta_{13})=0.085 \pm 0.005$

Use unified approach by Feldman and Cousins to build CL intervals

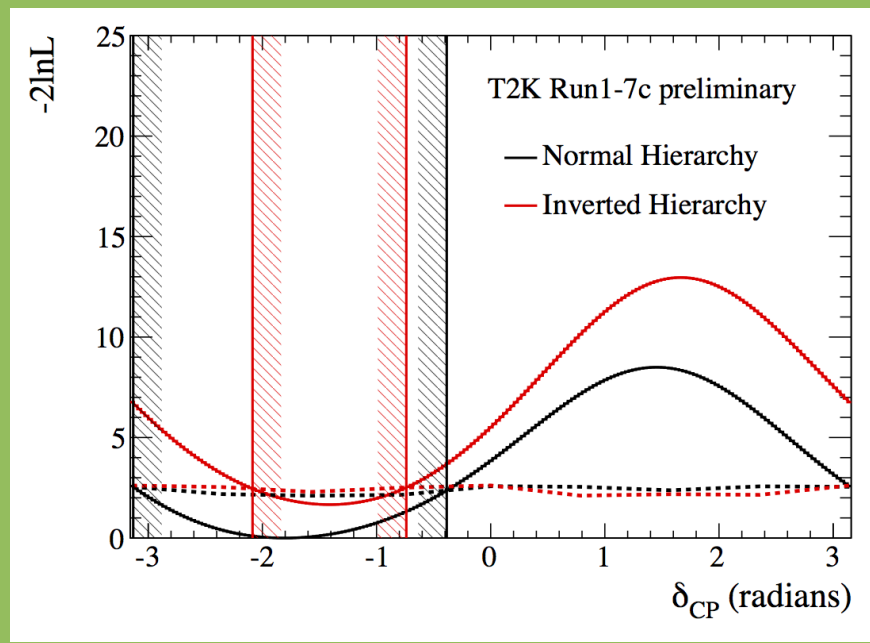
## Sensitivity



$\delta=-1.601$   
 $\sin^2(\theta_{23})=0.528$

$\Delta m^2_{32}=2.509 \cdot 10^{-3} \text{ eV}^2 \text{ c}^{-4}$   
 $\sin^2(\theta_{13})=0.0217$

## Data fit



CP conserving values outside of 90% CL intervals

# Combined $\nu$ - $\bar{\nu}$ analysis

## Model comparisons

Compare posterior probabilities of different models

T2K  
only

T2K Run1-7c preliminary	$\sin^2 \theta_{23} < 0.5$	$\sin^2 \theta_{23} > 0.5$	Line total
Inverted hierarchy	0.19	0.22	0.40
Normal hierarchy	0.27	0.33	0.60
Column total	0.45	0.55	1

T2K  
+  
reactor

T2K Run1-7c preliminary	$\sin^2 \theta_{23} < 0.5$	$\sin^2 \theta_{23} > 0.5$	Line total
Inverted hierarchy	0.10	0.14	0.25
Normal hierarchy	0.29	0.46	0.75
Column total	0.39	0.61	1

Mild preference for normal hierarchy and octant  $\sin^2 \theta_{23} > 0.5$



# $\nu_\mu/\bar{\nu}_\mu$ disappearance comparison

## Motivation

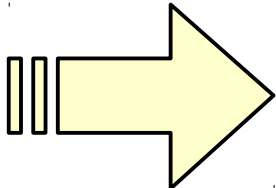
Can test the PMNS framework by comparing oscillations of neutrinos and anti-neutrinos in T2K data

$$\begin{array}{l} \nu \rightarrow \bar{\nu} \\ \delta_{\text{CP}} \rightarrow -\delta_{\text{CP}} \end{array}$$

$\nu_\mu/\bar{\nu}_\mu$  disappearance :

No CP odd order term, limited matter effect

$$\begin{aligned} P(\nu_\mu \rightarrow \nu_\mu) = & 1 - (c_{13}^4 \sin^2 2\theta_{23} + s_{23}^2 \sin^2 2\theta_{13}) \sin^2 \Delta_{\text{atm}} \\ & + \left\{ c_{13}^2 (c_{12}^2 - s_{13}^2 s_{23}^2) \sin^2 2\theta_{23} + s_{12}^2 s_{23}^2 \sin^2 2\theta_{13} - c_{13} \sin 2\theta_{13} \sin 2\theta_{23} \sin 2\theta_{12} \cos \delta \right\} \\ & \times \left\{ \frac{1}{2} \sin 2\Delta_{\text{solar}} \sin 2\Delta_{\text{atm}} + 2 \sin^2 \Delta_{\text{solar}} \sin^2 \Delta_{\text{atm}} \right\} \\ & - \left\{ \sin^2 2\theta_{12} (c_{23}^2 - s_{13}^2 s_{23}^2)^2 + s_{13}^2 \sin^2 2\theta_{23} (1 - c_\delta^2 \sin^2 2\theta_{12}) \right. \\ & \quad + 2s_{13} \sin 2\theta_{12} \cos 2\theta_{12} \sin \theta_{23} \cos 2\theta_{23} c_\delta \\ & \quad \left. - \frac{1}{2} c_{13} \sin 2\theta_{13} \sin 2\theta_{23} \sin 2\theta_{12} \cos \delta s_{23}^2 s_{12}^2 \right. \\ & \quad \left. + \sin^2 2\theta_{23} c_{13}^2 (c_{12}^2 - s_{13}^2 s_{23}^2) + s_{13}^2 s_{23}^2 \sin^2 2\theta_{13} \right\} \times \sin^2 \Delta_{\text{solar}} \end{aligned} \quad (26)$$



Expect similar disappearance pattern for  $\nu_\mu$  and  $\bar{\nu}_\mu$

# $\nu_\mu/\bar{\nu}_\mu$ disappearance comparison Analysis

Compare values of atmospheric parameters measured with neutrinos and anti-neutrinos

$$(\theta_{23}, \Delta m^2_{32})$$

$\nu$

VS

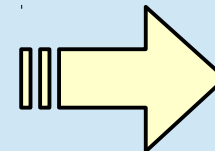
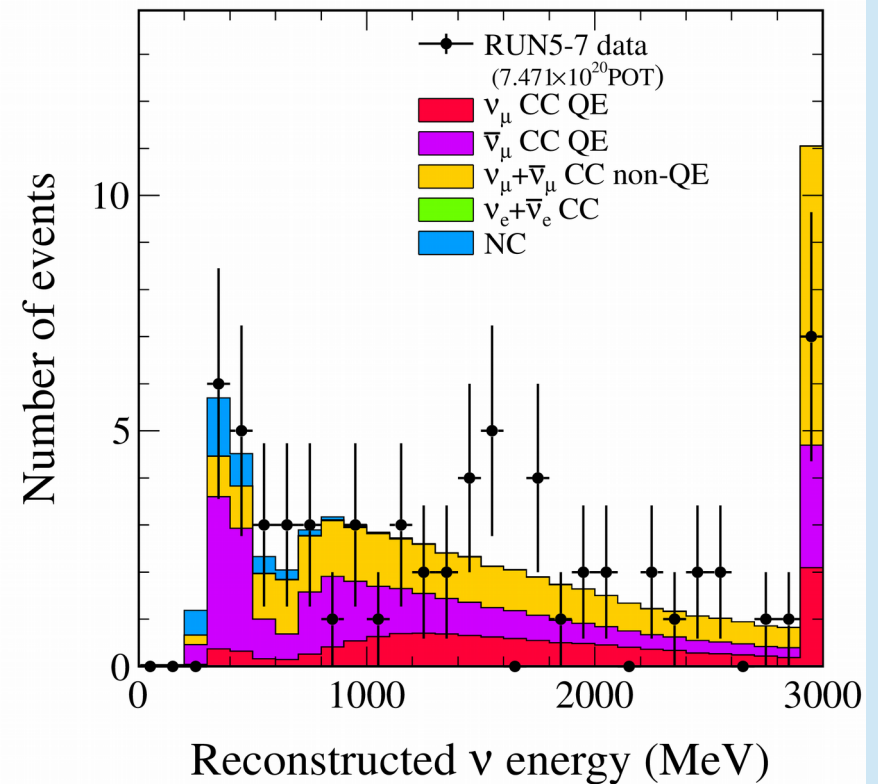
$$(\bar{\theta}_{23}, \bar{\Delta m^2}_{32})$$

$\bar{\nu}$

Other PMNS parameters common

- $\theta_{13}$ ,  $\theta_{12}$  and  $\Delta m^2_{21}$  constrained with PDG2015 values
- $\delta=0$  fixed

Significant  $\nu_\mu$  contamination in anti-neutrino mode

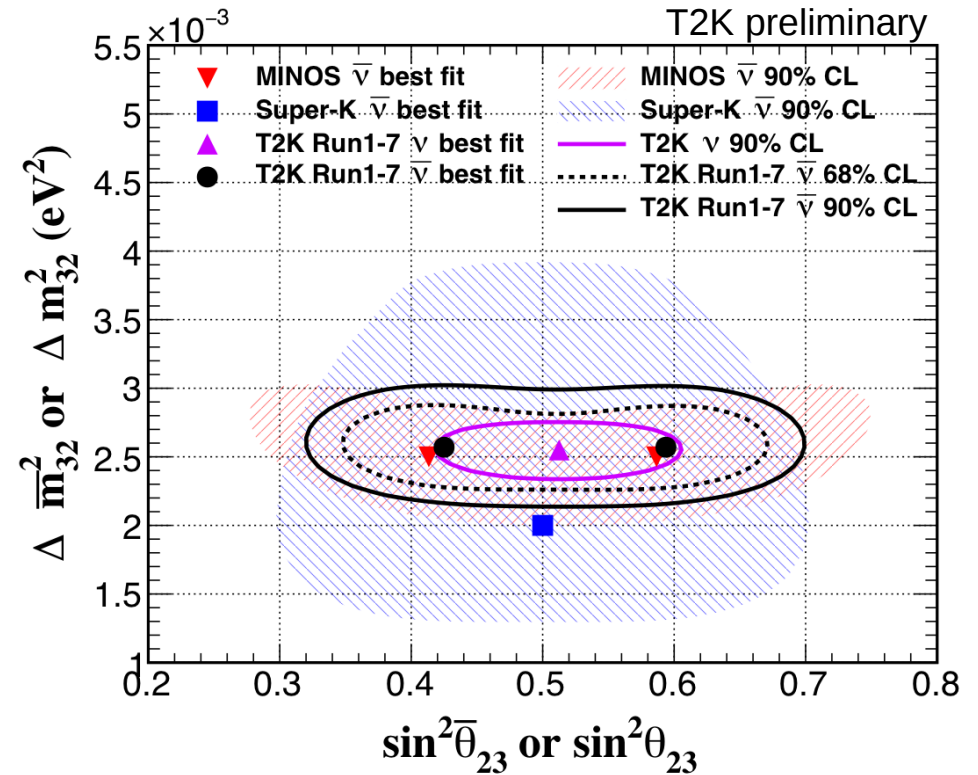
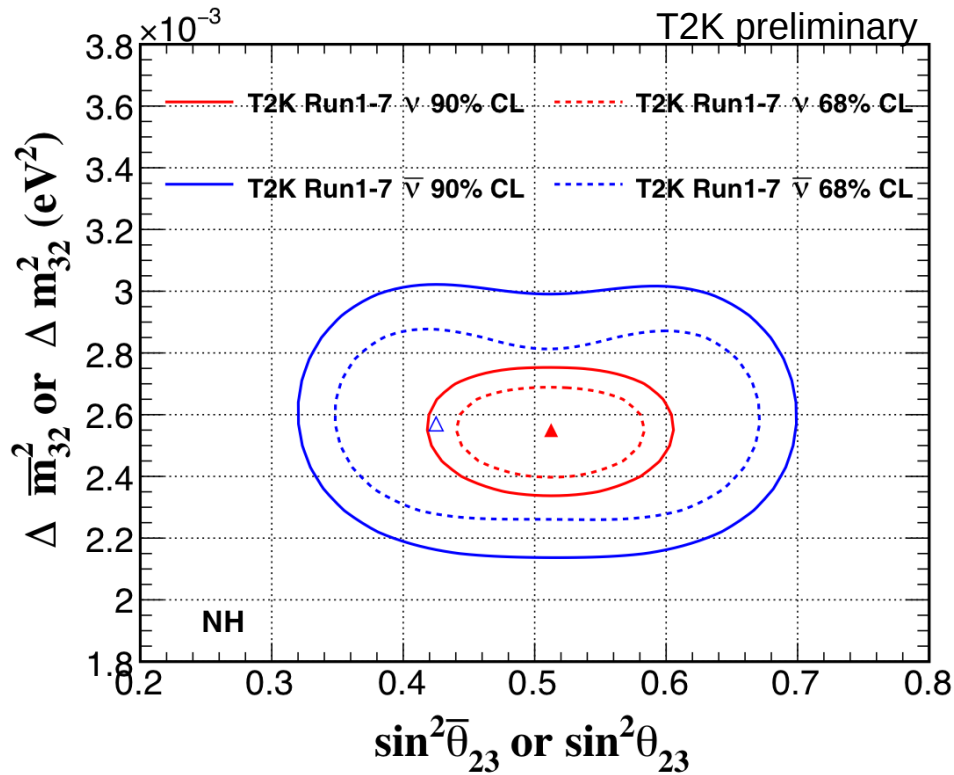


Joint fit of 1R $\mu$  neutrino and anti-neutrino mode samples

# $\nu_\mu/\bar{\nu}_\mu$ disappearance comparison

## Results

- No discrepancies between values measured for neutrinos and anti-neutrinos
- Best measurement of the parameters for anti-neutrinos
- Compatible with measurements by other experiments for anti-neutrinos



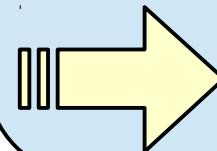
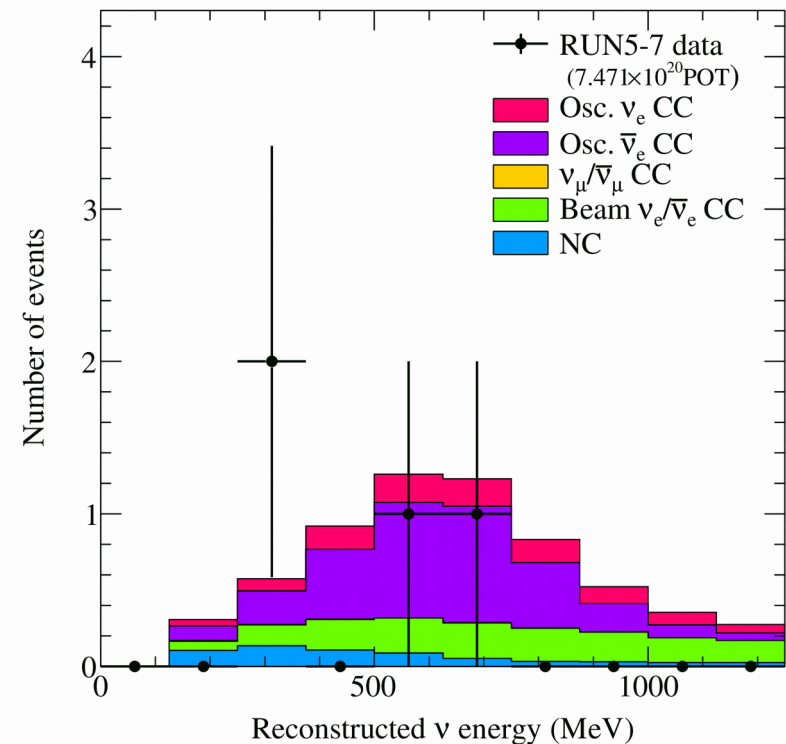
# $\bar{\nu}_e$ appearance Analysis

Look for  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillation using all data samples

Hypothesis test:

- $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) = \beta \times P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e, \text{PMNS})$
- Compare compatibility of data with  $\beta=1$  and  $\beta=0$

Significant  $\nu_e$  contamination in anti-neutrino mode

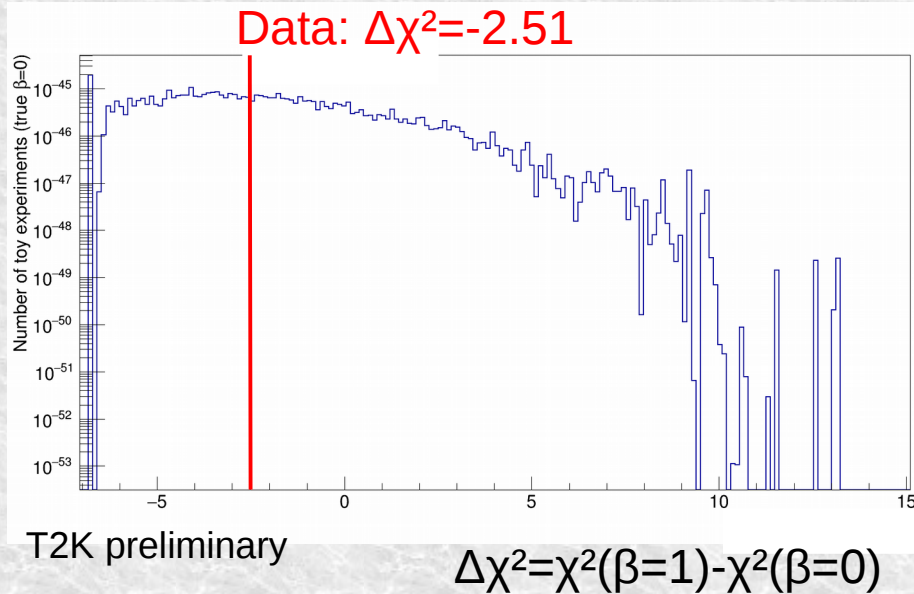


Cannot use simple measurement of  $\theta_{13}$  in 1Re anti-neutrino mode sample

# $\bar{\nu}_e$ appearance

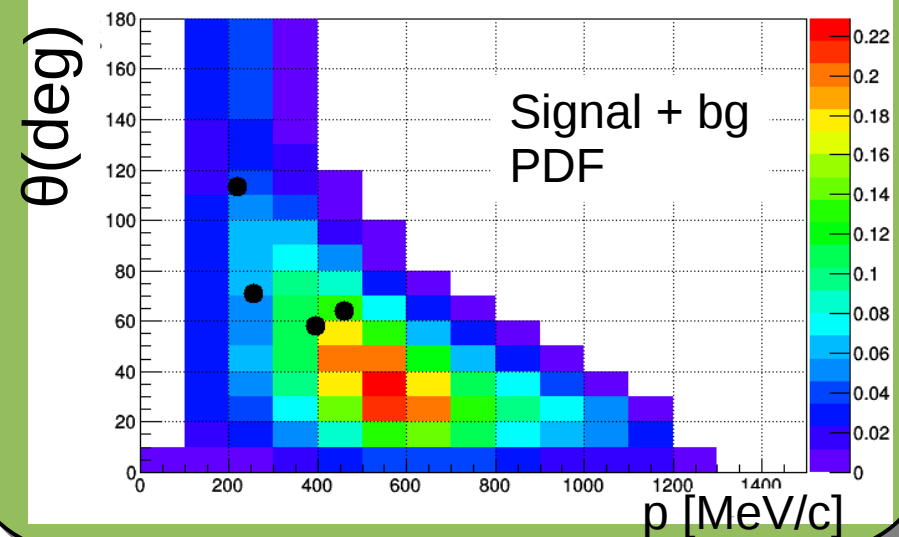
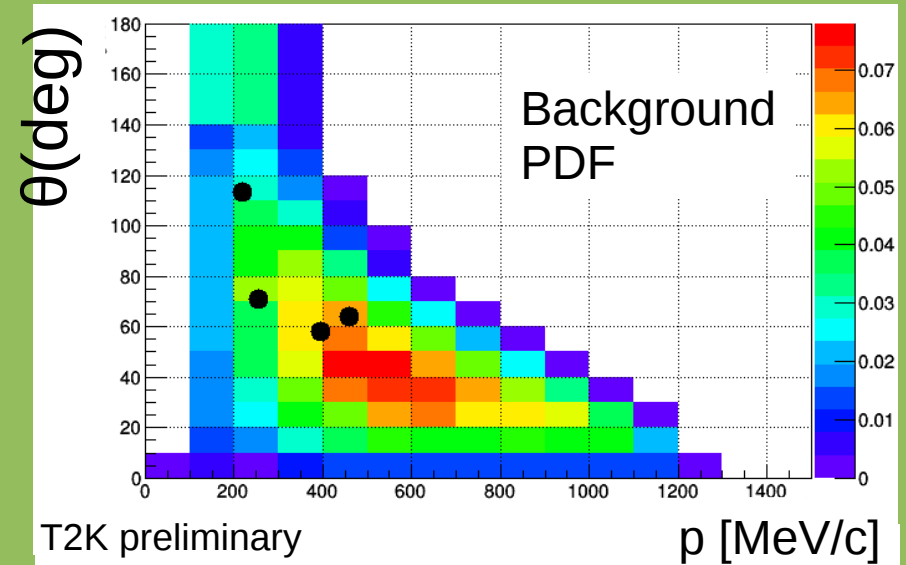
## Results – rate+shape analysis

### Test statistics distribution



P-value for no  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillation:  
 Rate only: 0.41  
 Rate+shape (Erec- $\theta$ ): 0.3742  
 Rate+shape ( $p$ - $\theta$ ): 0.4618

Events look more background like in lepton ( $p, \theta$ )

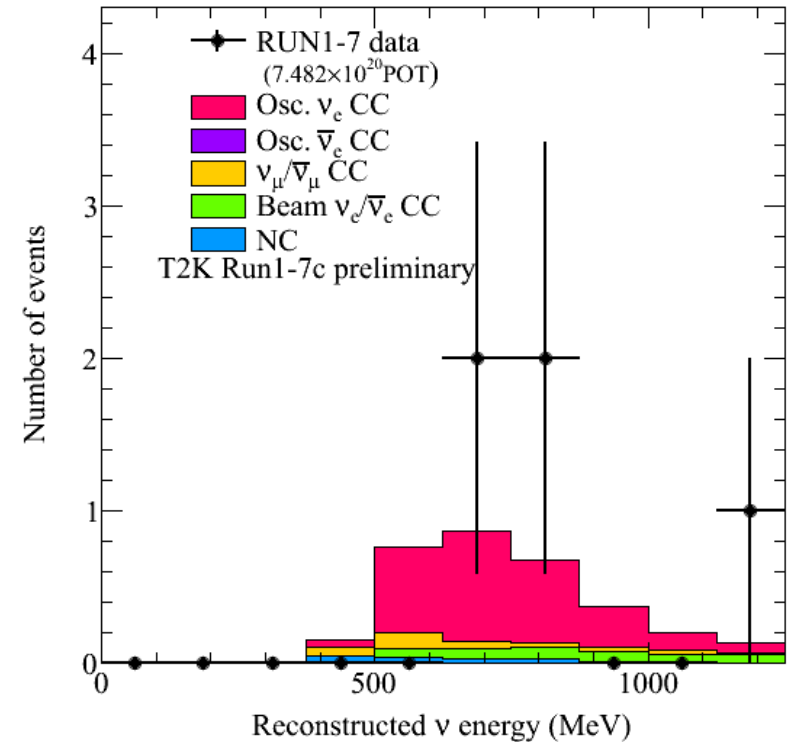
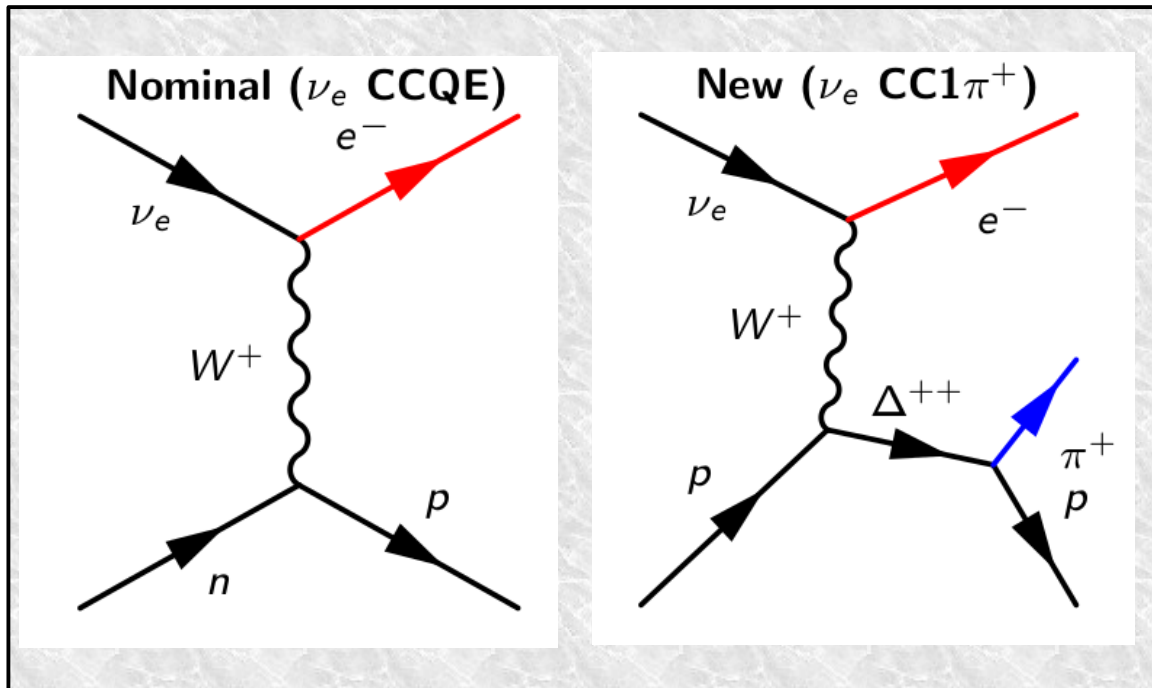


# Perspective for the future

# Near future

## Additional sample – $\nu_e$ CC1 $\pi$

- Selected by normal e-like selection + Michel  $e^-$
- Increase  $\nu$ -mode e-like statistics by  $\sim 11\%$
- 73% purity (defined as CC  $\nu_\mu \rightarrow \nu_e$ )

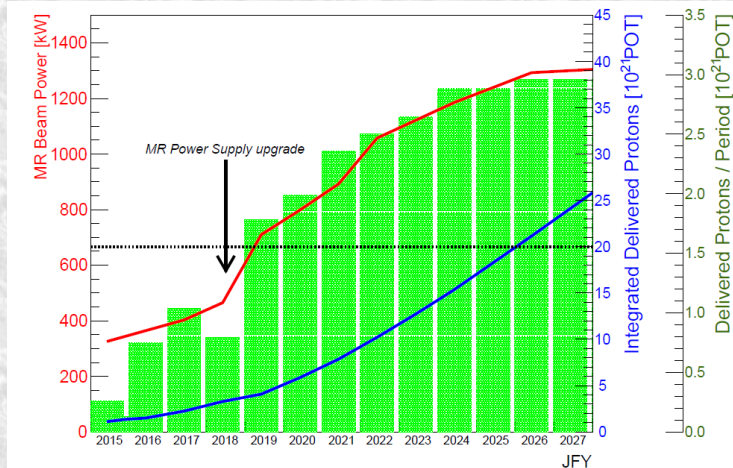


MC expectations (NH)	$\delta=0$	$\delta=\pi$	$\delta=-\pi/2$	$\delta=\pi/2$	Observed
	2.8	2.7	3.1	2.3	5



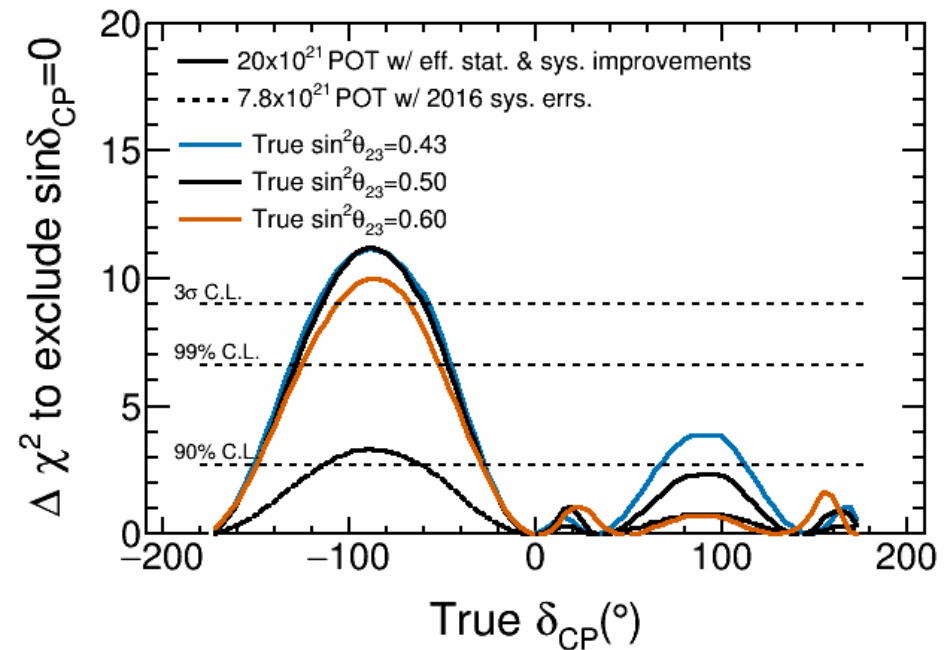
# Medium term Proposal for extended run

- Proposed an extended run until ~2025
- Increased statistics:  $7.8 \times 10^{21}$  POT  $\rightarrow$   $20 \times 10^{21}$  POT + analysis improvements
- Can exclude CP conservation at  $3\sigma$  in favorable case



Beam intensity improvement  
~400kW  $\rightarrow$  1.3MW

T2K phase 2 received stage 1  
status at summer J-PARC PAC



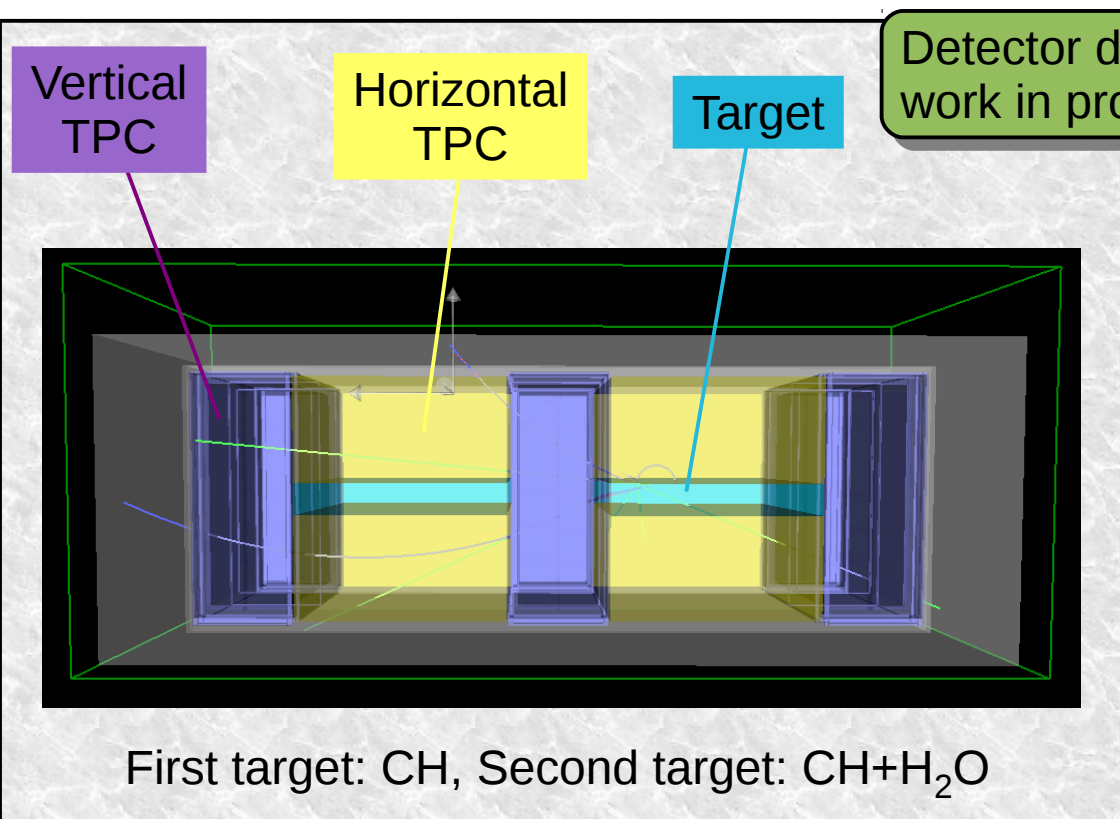
Assumes:

- unknown mass hierarchy
- 50% effective stat improvements
- 1/3 reduction of systematics

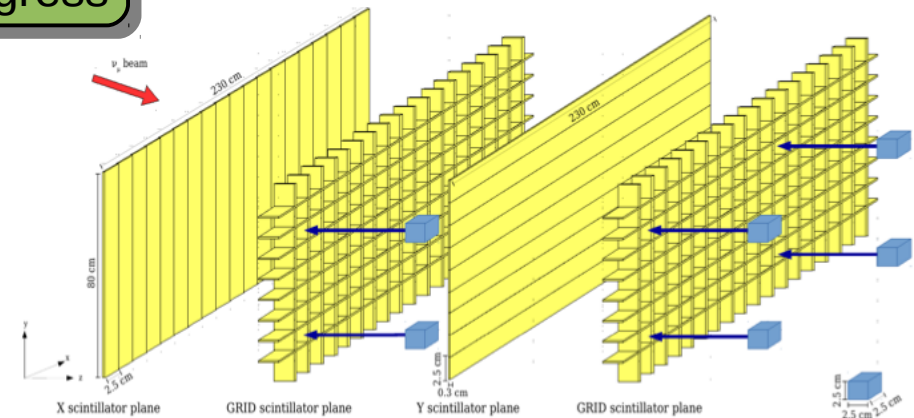
# Medium term (2021-2025)

## Near detector upgrade

- Extended run will require lower systematic uncertainties
- Design a new off-axis detector to avoid limitations of current ND280
- Main requirements:
  - ✓ Water target
  - ✓ Large angular acceptance
  - ✓ Better efficiency for low momentum  $p$  and  $\pi$



### Possible target concept



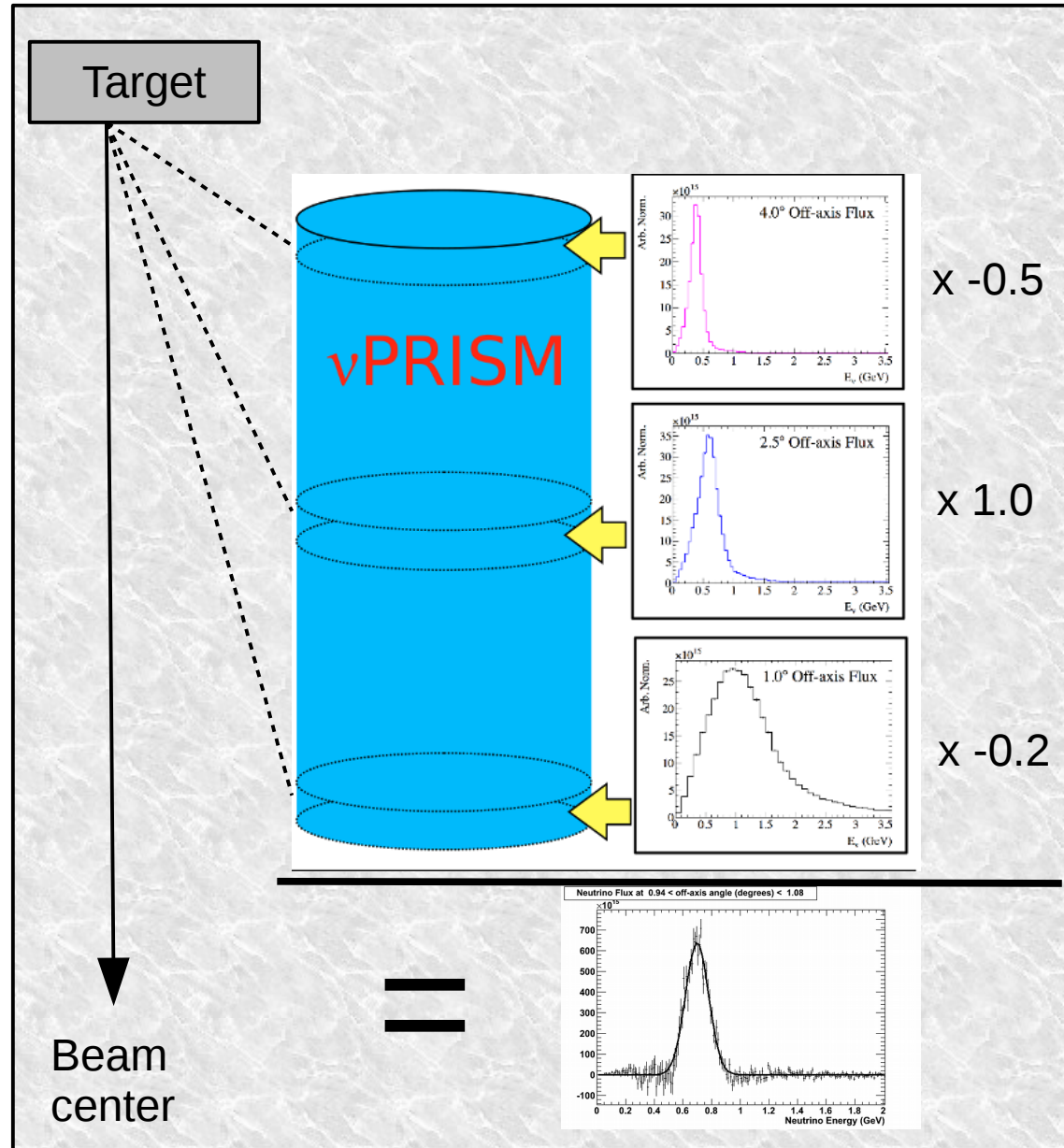
70% water in mass when filled

Workshop@CERN:  
 “Neutrino ND based on gas TPCs”  
<https://indico.cern.ch/event/568177/>

# Medium term Intermediate water Cherenkov detector

- Tall (~50m) water Cherenkov detector, ~1km from target
- Spans 1-4° off-axis angles
- Same target material (H<sub>2</sub>O), angular acceptance and detection technique than far detector
- Ability to look at rates and neutrino interactions as a function of true neutrino energy
- Recreate oscillated flux at SK with little need for interaction models

- Separate collaboration from T2K
- Received stage 1 status at July J-PARC PAC



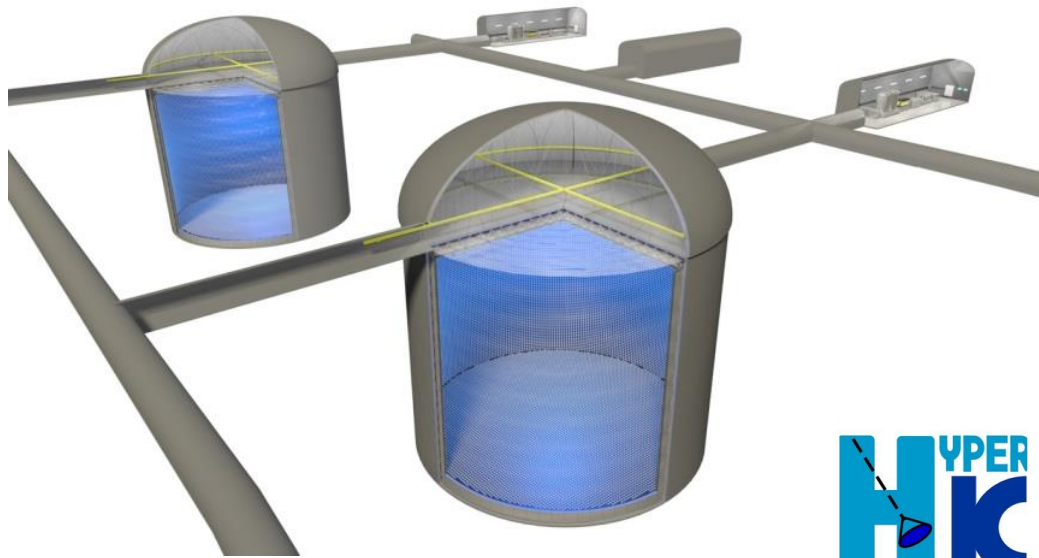
**vPRISM**

# Longer term Hyper-Kamiokande

- 2 tanks 60m height x 74m diameter
- 380 kton fiducial volume (SK: 22.5 kton)
- Improved photo-sensors
- Large statistics to study neutrino oscillations

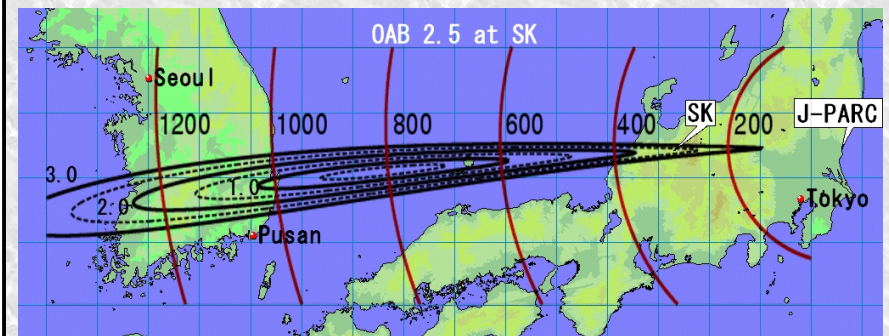
Rich physics program:

- Long baseline neutrinos
- Atmospheric neutrinos
- Proton decay
- Solar / astrophysical / supernova neutrinos



Hyper Kamiokande  
@HyperKamiokande

Also proposal to have  
2<sup>nd</sup> tank in Korea



LOI: 1109.3262 [hep-ex]  
Physics potential: 1309.0184 [hep-ex]



- Presented neutrino oscillation results from combined analysis of T2K  $\nu_\mu/\bar{\nu}_\mu/\nu_e/\bar{\nu}_e$  samples:
  - ✓ Results compatible with maximal  $\nu_\mu$  disappearance
  - ✓  $\theta_{13}$  measurement from T2K alone compatible with measurement by reactor experiments
  - ✓ Favor  $\delta \sim -\pi/2$  with and without combining with reactor experiments  
CP conserving values outside of 90%CL interval when combining
  - ✓ Mild preference for normal hierarchy and octant  $\sin^2\theta_{23} > 0.5$
- Accumulated  $1.5 \times 10^{21}$  POT out of  $7.8 \times 10^{21}$  approved: a lot more results to come
- Proposal to extend run to  $20 \times 10^{21}$  POT with upgraded near detector and additional intermediate detector

**Additional slides**

# Neutrino oscillations

## Looking for second order effects

Oscillation probabilities for a muon neutrino beam

$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_e) = & \underbrace{4c_{13}^2 s_{13}^2 s_{23}^2 \sin^2 \Delta_{31}}_{\text{Leading term}} \quad \theta_{13} \\
 & + 8c_{13}^2 s_{12} s_{13} s_{23} (c_{12} c_{23} \cos \delta - s_{12} s_{13} s_{23}) \cos \Delta_{32} \sin \Delta_{31} \sin \Delta_{21} \quad \text{CPC} \\
 & - 8c_{13}^2 c_{12} c_{23} s_{12} s_{13} s_{23} \sin \delta \sin \Delta_{32} \sin \Delta_{31} \sin \Delta_{21} \quad \text{CPV} \\
 & + 4s_{12}^2 c_{13}^2 (c_{12}^2 c_{23}^2 + s_{12}^2 s_{23}^2 s_{13}^2 - 2c_{12} c_{23} s_{12} s_{23} s_{13} \cos \delta) \sin^2 \Delta_{21} \quad \text{Solar}
 \end{aligned}$$

$$P(\nu_\mu \rightarrow \nu_\mu) \sim 1 - \left( \underbrace{\cos^4 \theta_{13}}_{\text{Leading-term}} \times \underbrace{\sin^2 2\theta_{23}}_{\text{Next-to-leading}} + \sin^2 2\theta_{13} \times \underbrace{\sin^2 \theta_{23}}_{\text{Next-to-leading}} \right) \times \sin^2 \frac{\Delta m_{31}^2 \times L}{4E}$$

$$c_{ij} = \cos \theta_{ij}, s_{ij} = \sin \theta_{ij}$$

$$\Delta_{ij} = \Delta m_{ij}^2 \frac{L}{4E_\nu}$$



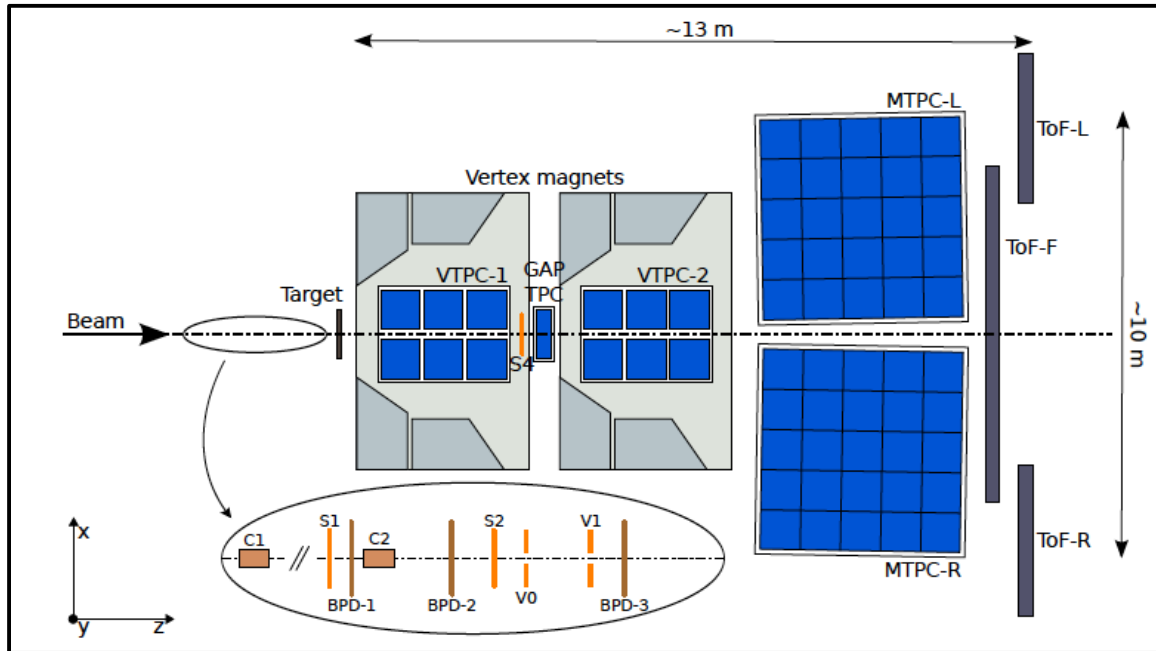
# Analysis description

## Hadron production measurements

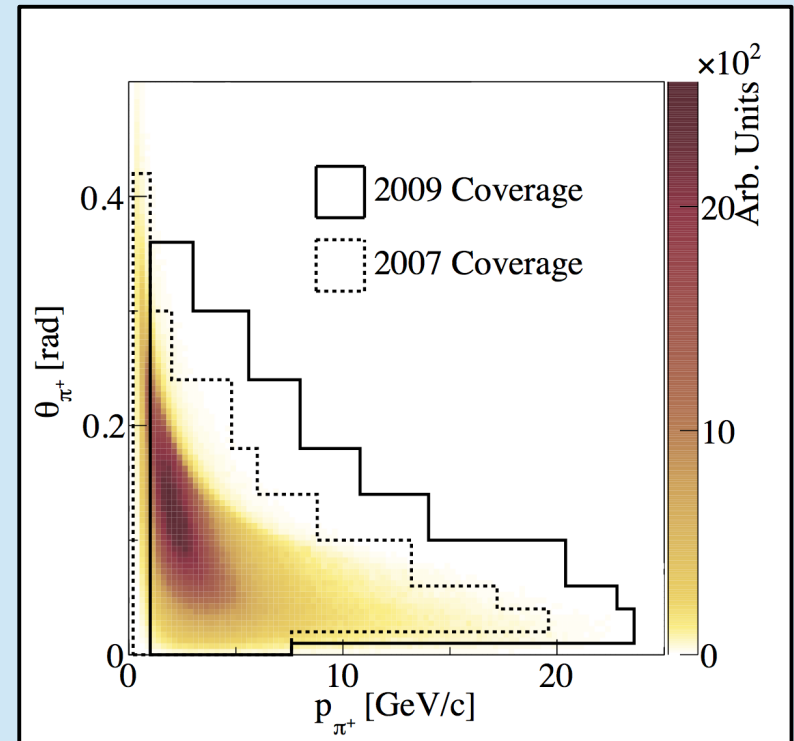
The NA61/Shine experiment measures hadron production from 30 GeV protons on carbon

2 targets:

- › 'thin'  $\sim 0.04\lambda$
- › Replica T2K target



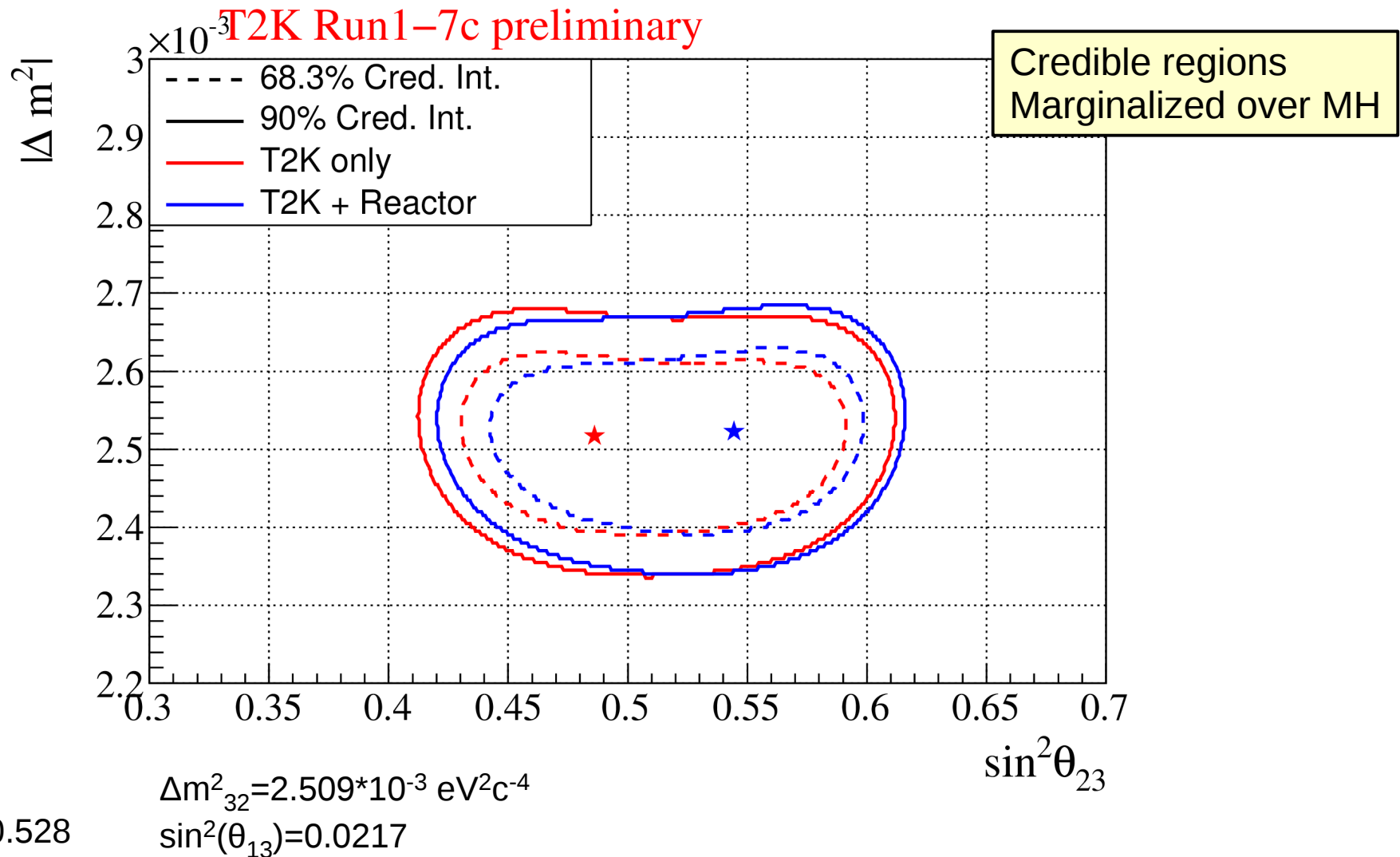
Covers most of the phase space for T2K neutrino production



# Combined $\nu$ - $\bar{\nu}$ analysis

## Effect of reactor constraint for atmospheric

Sensitivity for the atmospheric parameters changes depending on whether the reactor constraint is used or not



# Combined $\nu$ - $\bar{\nu}$ analysis

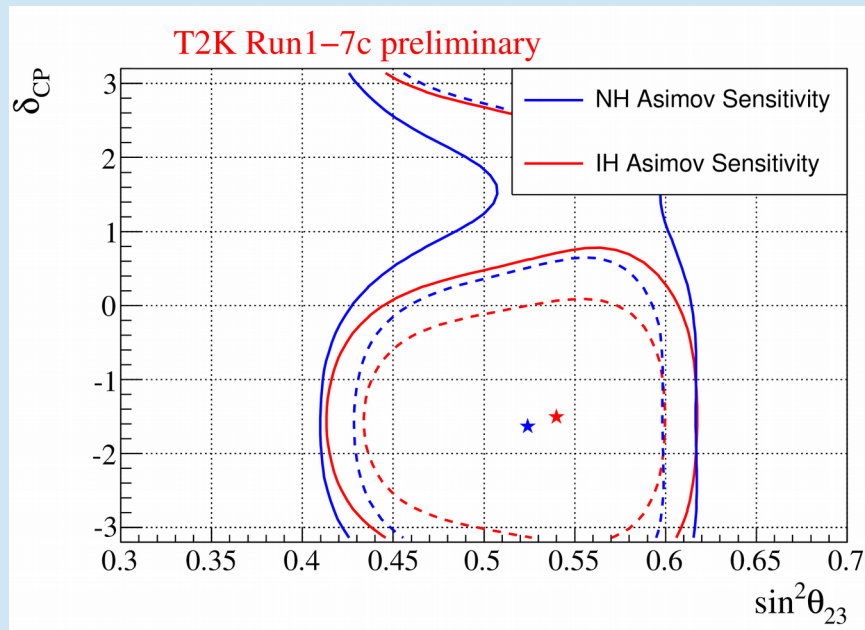
## $\theta_{23}$ and $\delta$ – T2K + reactor

Reactor constraint (PDG2015)  
 $\sin^2(2\theta_{13}) = 0.085 \pm 0.005$

Sensitivity

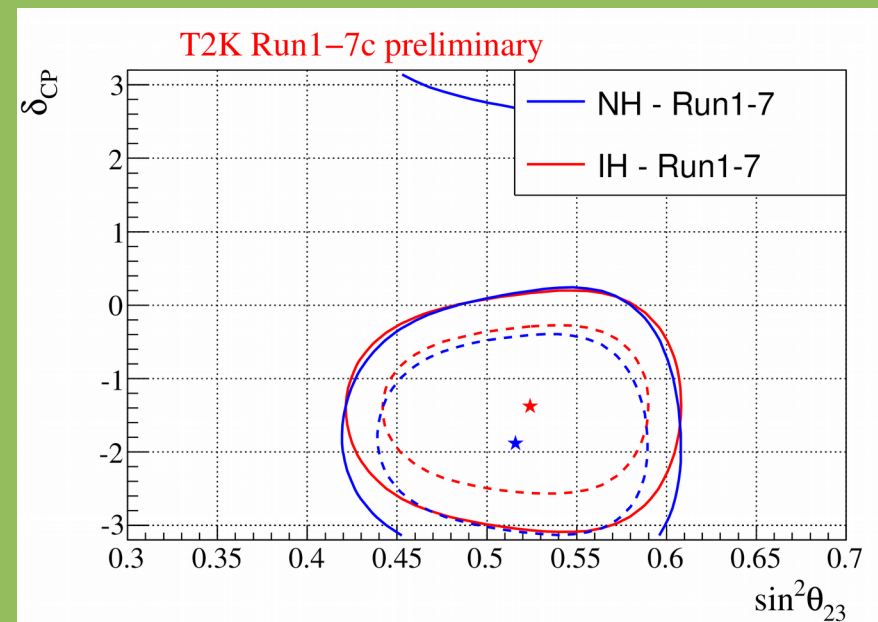
Fixed  $\Delta\chi^2$  68% and  
 90% CL regions

Data fit



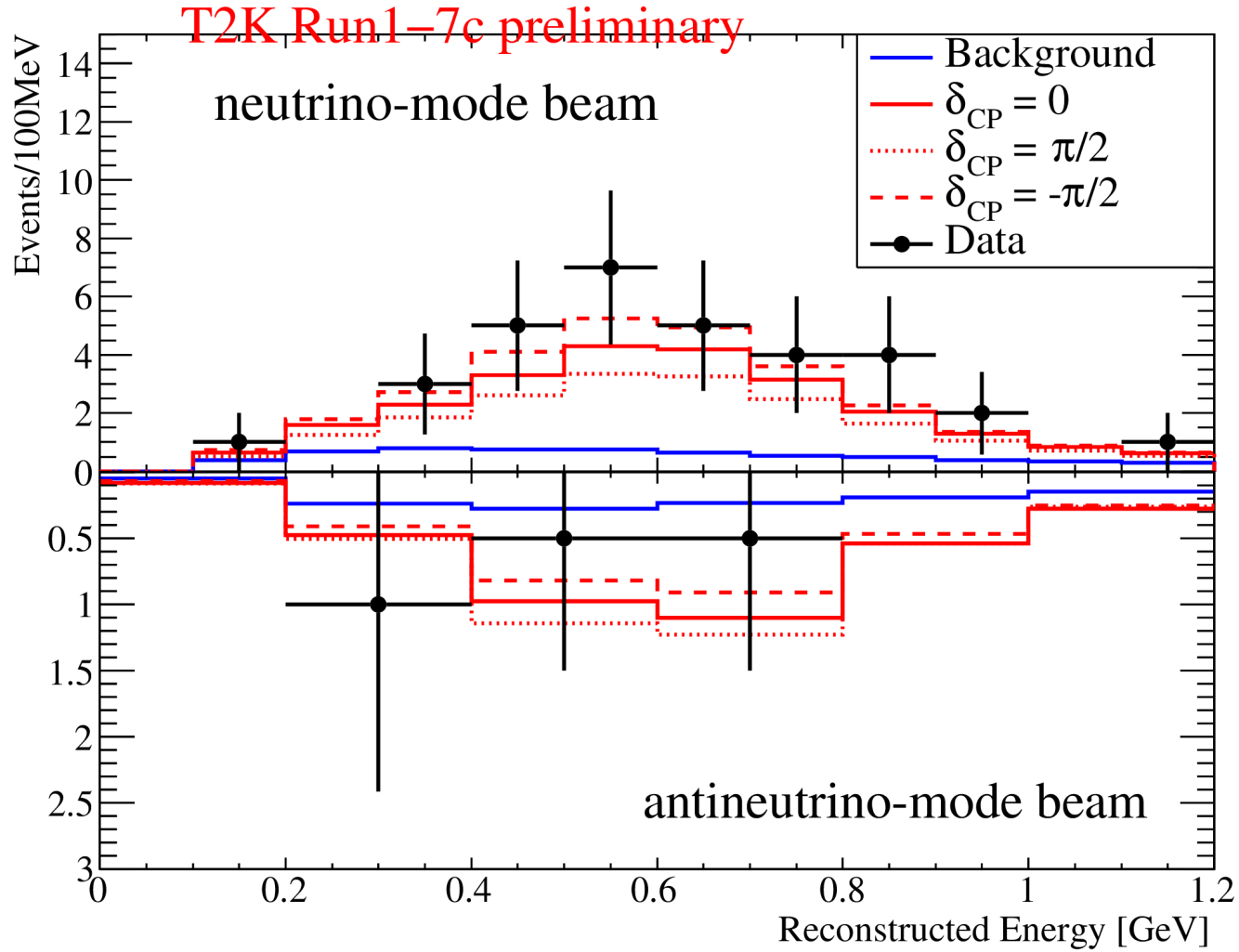
$\delta = -1.601$   
 $\sin^2(\theta_{23}) = 0.528$

$\Delta m_{32}^2 = 2.509 \times 10^{-3} \text{ eV}^2 \text{ c}^{-4}$   
 $\sin^2(\theta_{13}) = 0.0217$



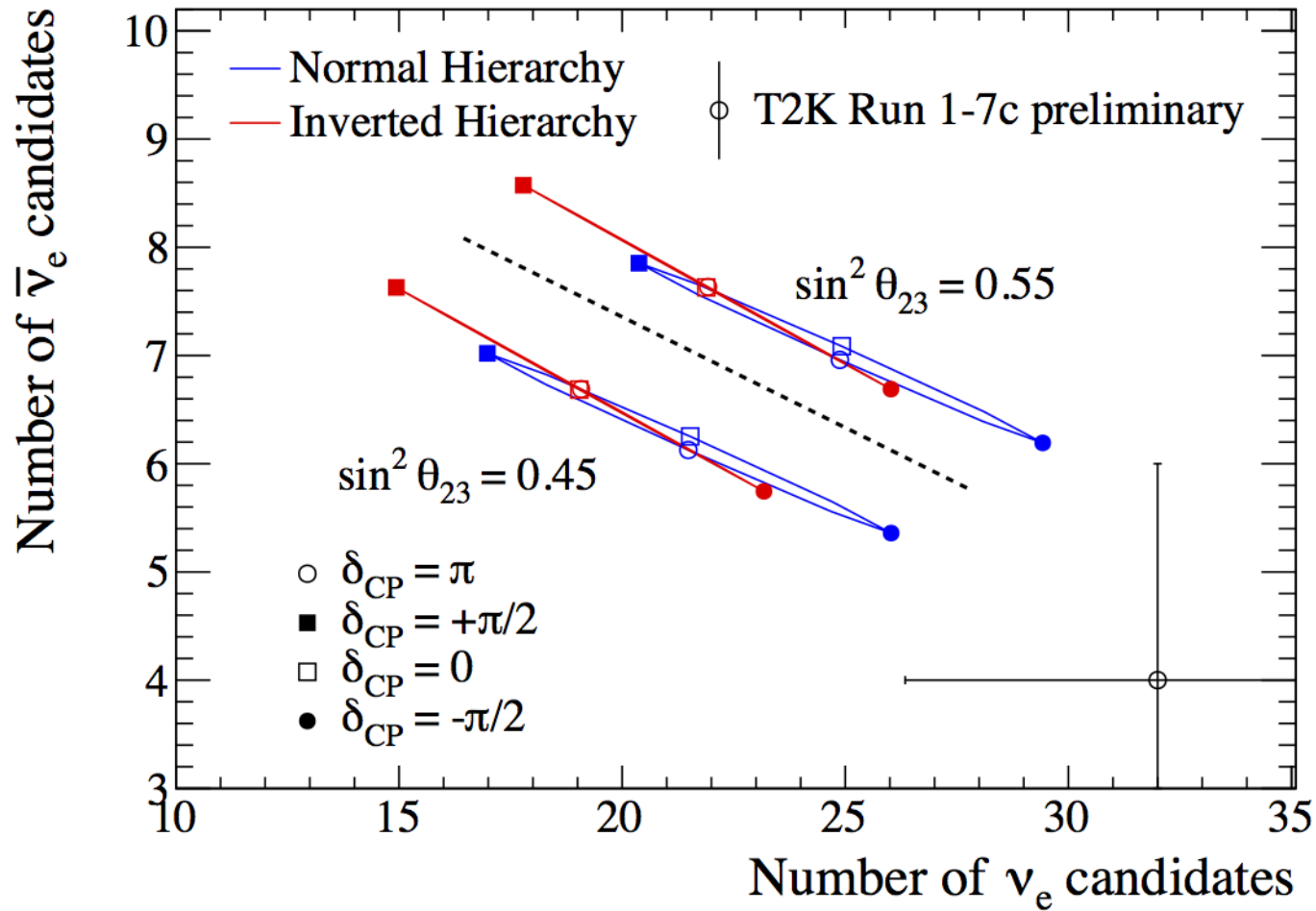
# Combined $\nu$ - $\bar{\nu}$ analysis

## Electron-like spectra



# Combined $\nu\bar{\nu}$ analysis

## Electron-like number of events



# Systematic uncertainties

## T2K systematics uncertainties (joint oscillation analysis)

Fractional error on the number of expected events at SK with and without ND280

	$\nu_\mu$ sample 1R $_\mu$ FHC	$\nu_e$ sample 1R $_e$ FHC	$\bar{\nu}_\mu$ sample 1R $_\mu$ RHC	$\bar{\nu}_e$ sample 1R $_e$ RHC
$\nu$ flux w/o ND280	7,6%	8,9%	7,1%	8,0%
$\nu$ flux with ND280	3,6%	3,6%	3,8%	3,8%
$\nu$ cross-section w/o ND280	7,7%	7,2%	9,3%	10,1%
$\nu$ cross-section with ND280	4,1%	5,1%	4,2%	5,5%
$\nu$ flux+cross-section	2,9%	4,2%	3,4%	4,6%
Final or secondary hadron int.	1,5%	2,5%	2,1%	2,5%
Super-K detector	3,9%	2,4%	3,3%	3,1%
Total w/o ND280	12,0%	11,9%	12,5%	13,7%
<b>Total with ND280</b>	<b>5,0%</b>	<b>5,4%</b>	<b>5,2%</b>	<b>6,2%</b>

# Systematic uncertainties

## T2K systematics uncertainties (joint oscillation analysis)

Fractional error on the number of expected events at SK

	$\nu_\mu$ sample 1R $_\mu$ FHC	$\nu_e$ sample 1R $_e$ FHC	$\bar{\nu}_\mu$ sample 1R $_\mu$ RHC	$\bar{\nu}_e$ sample 1R $_e$ RHC	1R $_e$ FHC/RHC
$\nu$ flux+cross-section constrained by ND280	2,8%	2,9%	3,3%	3,2%	2,2%
$\nu_e/\nu_\mu$ and $\bar{\nu}_e/\bar{\nu}_\mu$ cross-sections	0,0%	2,7%	0,0%	1,5%	3,1%
NC $\gamma$	0,0%	1,4%	0,0%	3,0%	1,5%
NC other	0,8%	0,2%	0,8%	0,3%	0,2%
Final or secondary hadron int.	1,5%	2,5%	2,1%	2,5%	3,6%
Super-K detector	3,9%	2,4%	3,3%	3,1%	1,6%
<b>Total</b>	<b>5,0%</b>	<b>5,4%</b>	<b>5,2%</b>	<b>6,2%</b>	<b>5,8%</b>