

Dark Energy

Daniel Gruen, NASA Einstein Fellow at SLAC/KIPAC

APS DPF Meeting, Fermilab, Aug 3 2017



THE DARK ENERGY SURVEY

Dark Energy Survey Year 1 Results: Cosmological Constraints

Daniel Gruen, NASA Einstein Fellow at SLAC/KIPAC
on behalf of the DES Collaboration

APS DPF Meeting, Fermilab, Aug 3 2017

Structure of this talk

- Dark Energy
 - Introduction
 - How to observe Dark Energy
 - How to explain Dark Energy
 - Recent results
 - geometry of expanding universe vs. growth of structure
 - early universe vs. late-time universe
- Cosmological Constraints from DES

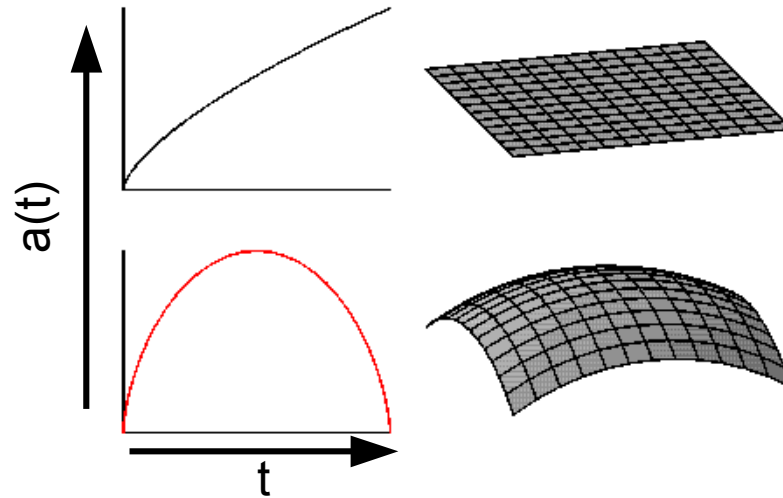
What goes up must come down?

- on large scales, Universe described as homogenous fluid in expanding space

$$\left(\frac{\dot{a}}{a} \right)^2 = \Omega_{m,0} a^{-3} + \text{radiation} + \text{relativistic species} + \text{curvature}$$

scale factor of Universe

mean matter density

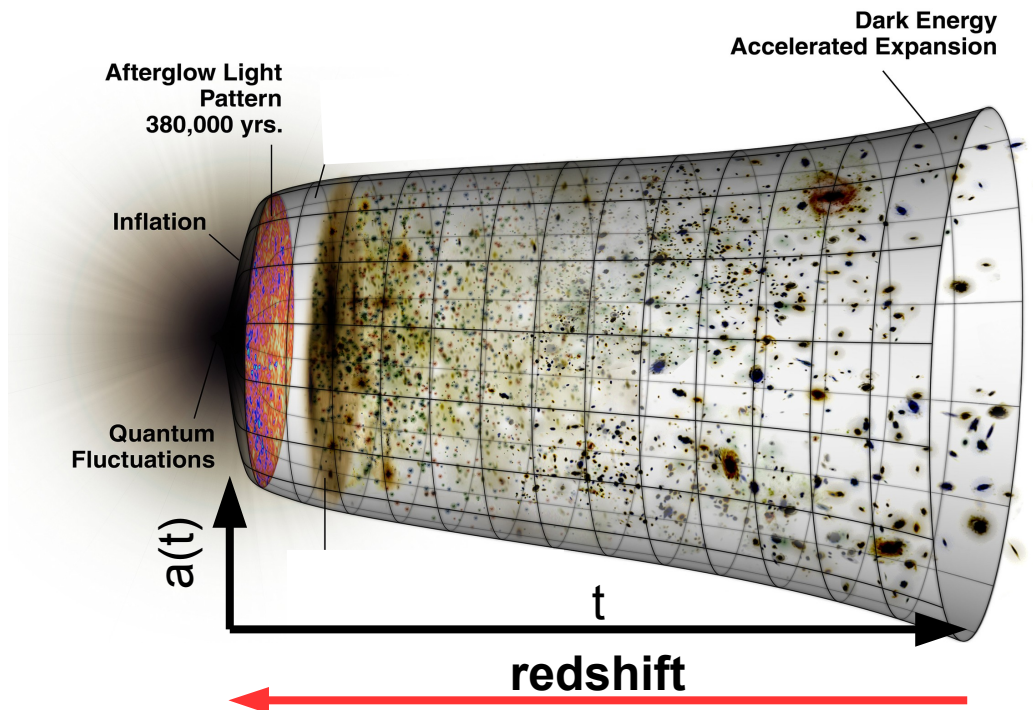


What goes up keeps getting faster!

- on large scales, Universe described as homogenous fluid in expanding space

$$\left(\frac{\dot{a}}{a}\right)^2 = \Omega_{m,0}a^{-3} + \Omega_{\Lambda}$$

constant
vacuum
energy
density



What goes up keeps getting faster!

- on large scales, Universe described as homogenous fluid in expanding space

$$\left(\frac{\dot{a}}{a}\right)^2 = \Omega_{m,0} a^{-3} + \Omega_{\Lambda}$$

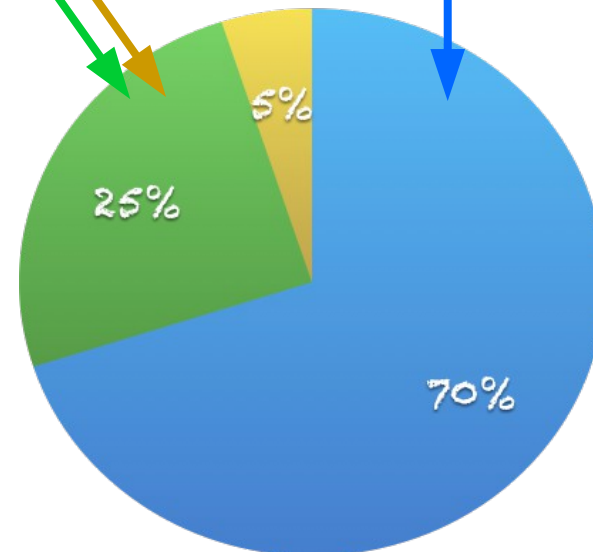
4 additional parameters:

σ_8 / S_8 : amplitude of density fluctuations

m_ν / Ω_ν : mass/density of neutrinos

$h / H / \left(\frac{\dot{a}}{a}\right)$: rate of expansion today

n_s : scale dependence of early density fluctuations



$\Omega_m=0.3, \sigma_8=0.8$
“fiducial Λ CDM”

This is a remarkably odd model

- 70% of energy content of Universe is an unknown substance that appears like vacuum energy, but 120 orders of magnitude smaller than QFT prediction
- 80% of matter is an unknown matter-like substance that does only interacts via gravitation
- We have a wide range of independent observations that cannot be explained without these assumptions

This is a remarkably odd model, but alternatives are even odder

Lovelock (1969) theorem:

GR + Λ are the only **local, second-order** gravitational field equations that can be derived from a **four-dimensional action** that is constructed **solely from the metric tensor**, and admitting Bianchi identities.

Theory zoo of:

- non-local field equations
- higher order field equations $f(R)$
- higher dimensions:
e.g. strings & branes
- new degrees of freedom
= substances

This is a remarkably odd model, but alternatives are even odder

Lovelock (1969) theorem:

GR + Λ are the only **local, second-order** gravitational field equations that can be derived from a **four-dimensional action** that is constructed **solely from the metric tensor**, and admitting Bianchi identities.

Theory zoo of:

- non-local field equations
- higher order field equations $f(R)$
- higher dimensions:
e.g. strings & branes
- new degrees of freedom
= substances

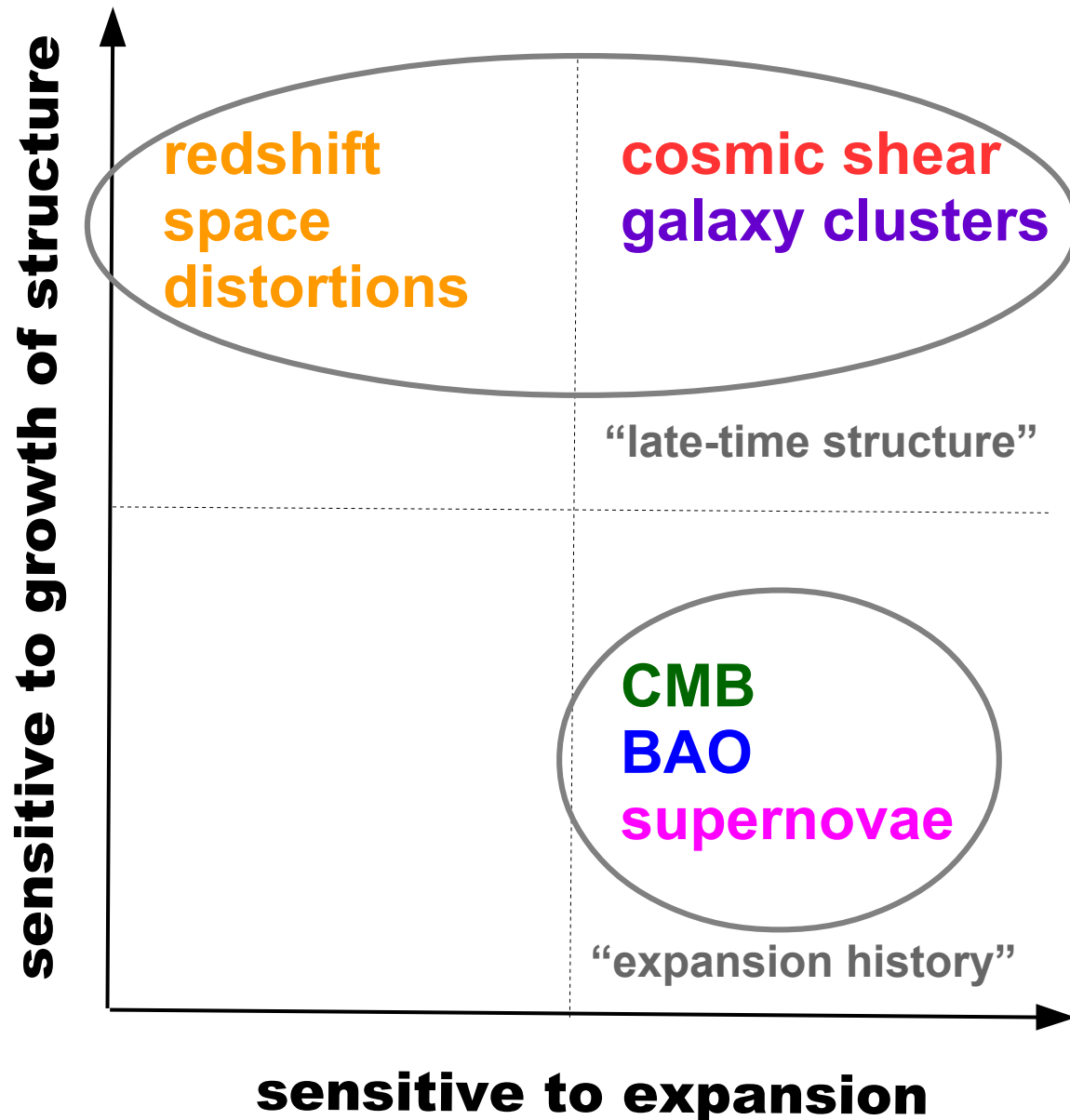
Need phenomenological tests of the most simple model:

Are data from
early Universe
and **late Universe**
fit by the same parameters?

Do measurements of
cosmic distances and
growth of **structure**
agree?

Does the dark energy density
change as space expands?
“Equation of state” parameter
 $w = \text{pressure/density}$

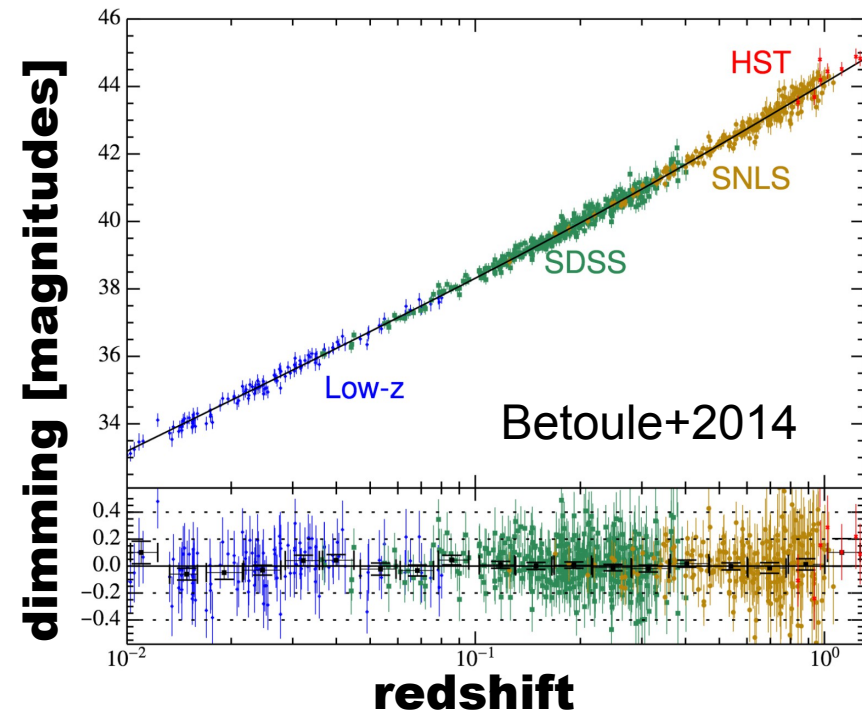
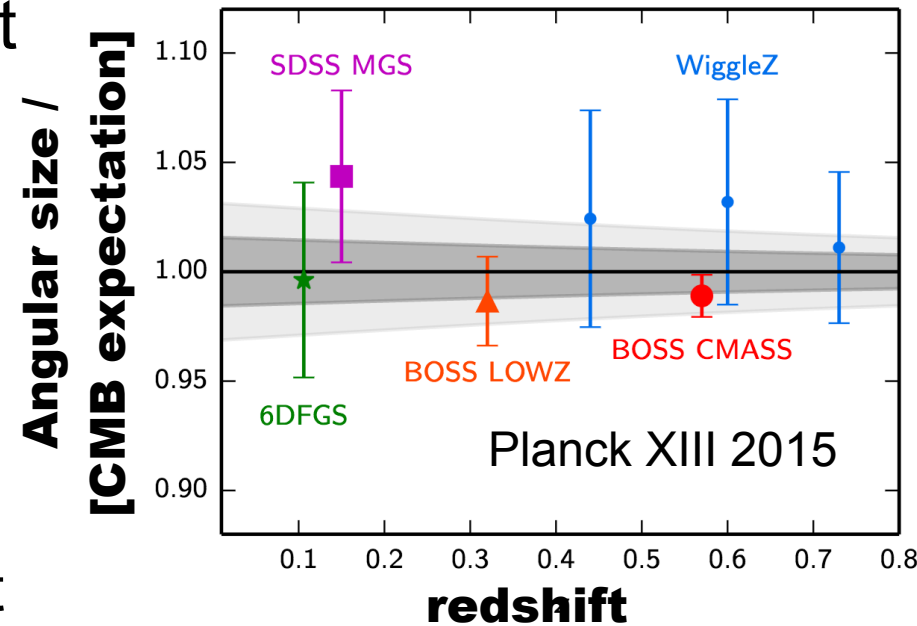
How to survey Dark Energy



Q: Do all these measurements agree with predictions in the same, fiducial Λ CDM model?

Measurements of expansion history

- Comparison of distance and redshift
- **Standard ruler:**
angle subtended by known scale
 - **CMB**: sound horizon in early Universe (380,000 years)
 - **BAO**: same scale, but expanded at later times (billions of years)
- **Standard candle:** brightness of source with known luminosity
 - **SNe**: luminosity can be determined from duration/color
- These are consistent and very tightly constrain $w=-1$, Ω_m , Ω_{DE} , flatness

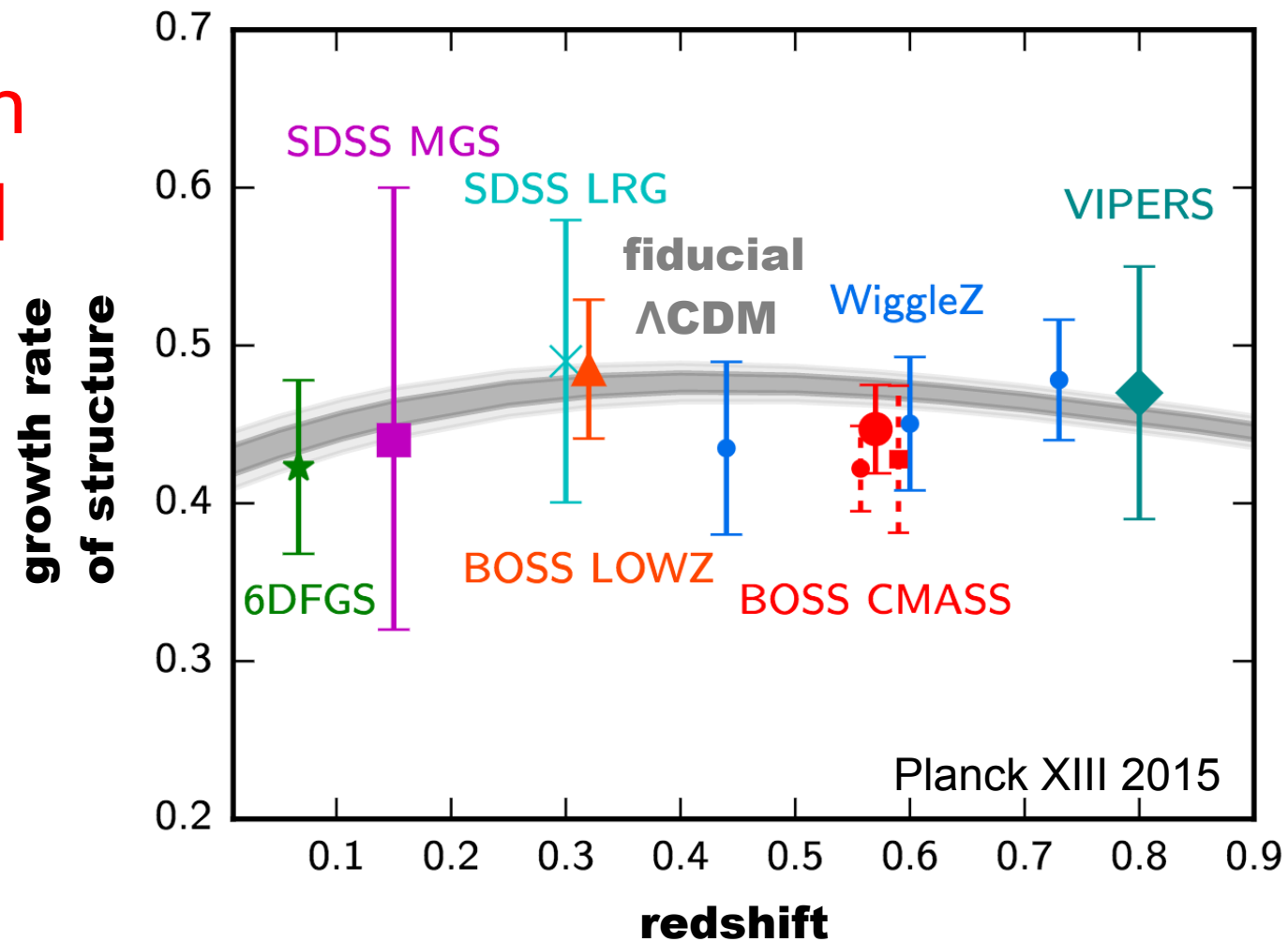


Measurement of late-time structure

- redshift space distortions (RSD):

growth rate

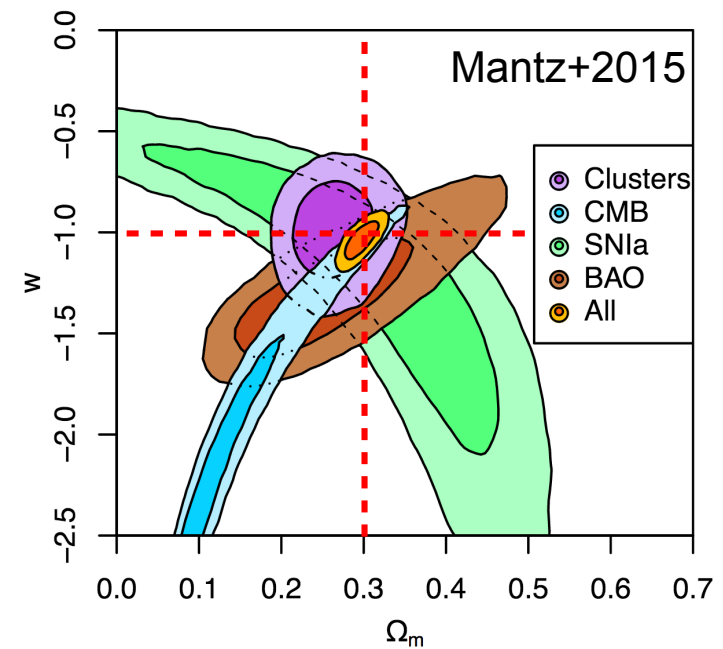
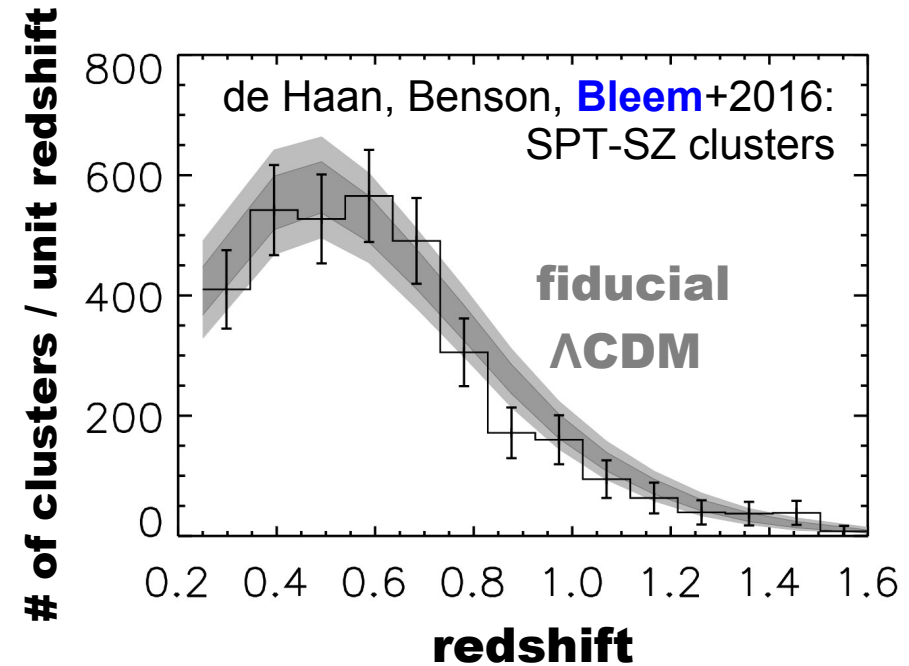
consistent with
fiducial Λ CDM



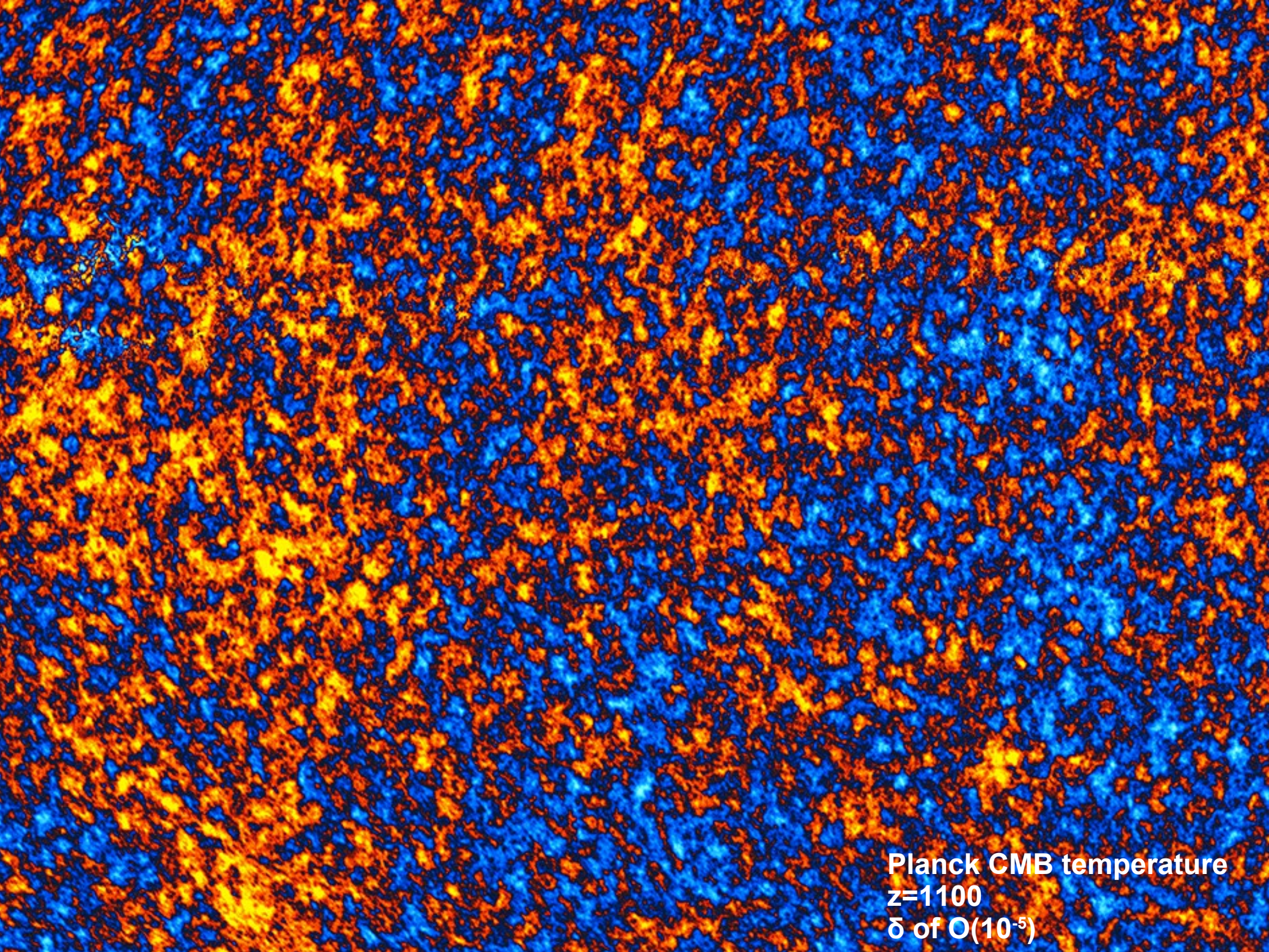
Measurement of late-time structure

✓ RSD

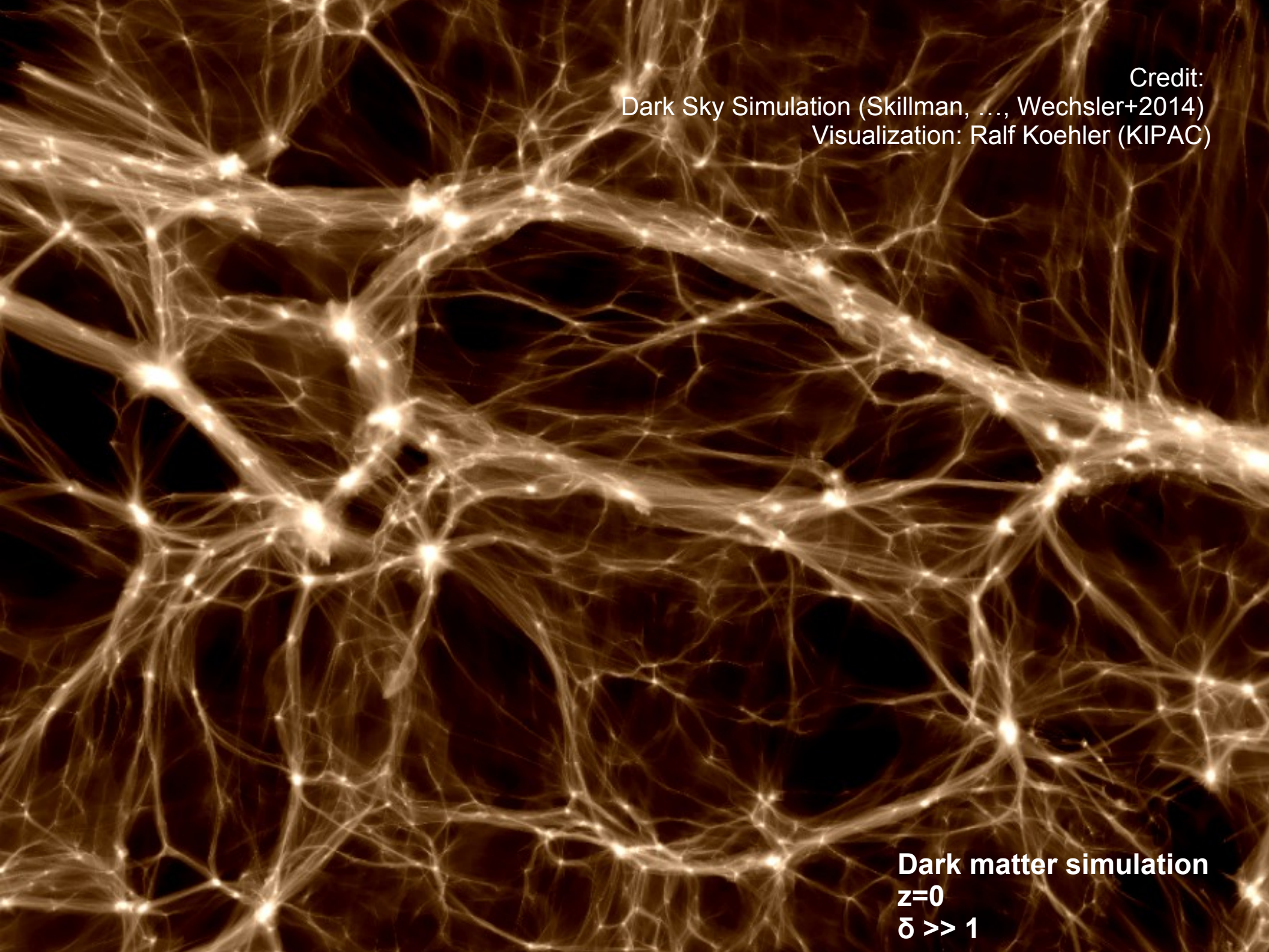
- Galaxy clusters:
count of clusters as a
function of mass and redshift
consistent with fiducial Λ CDM



see also: talks by [Antonella Palmese & Huan Lin](#),
[Yuanyuan Zhang](#), [Mathew Madhavacheril](#)



Planck CMB temperature
z=1100
 δ of $O(10^{-5})$



Credit:
Dark Sky Simulation (Skillman, ..., Wechsler+2014)
Visualization: Ralf Koehler (KIPAC)

Dark matter simulation
 $z=0$
 $\delta \gg 1$

Measurement of late-time structure

✓ RSD

✓ Galaxy clusters

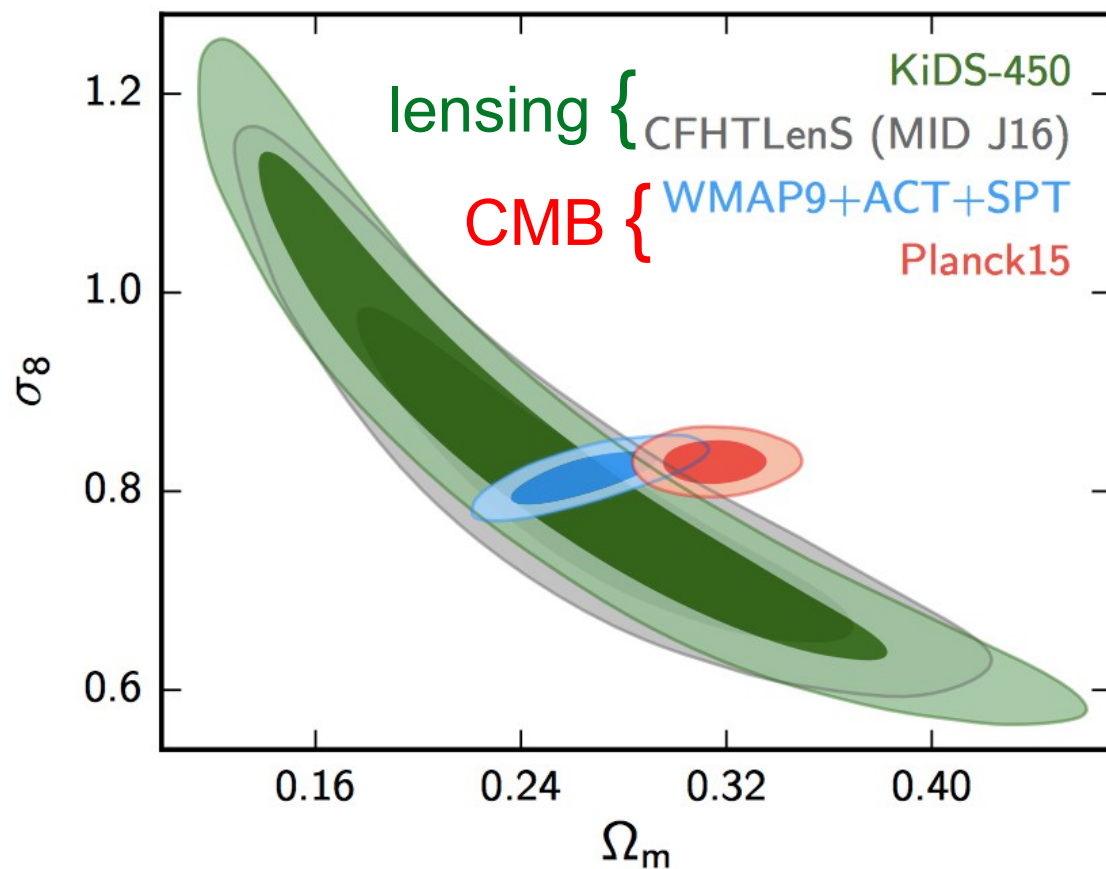
• cosmic shear:

recent studies have
claimed 2-3 σ offset
from Planck CMB in
 Ω_m - σ_8

A non-issue?

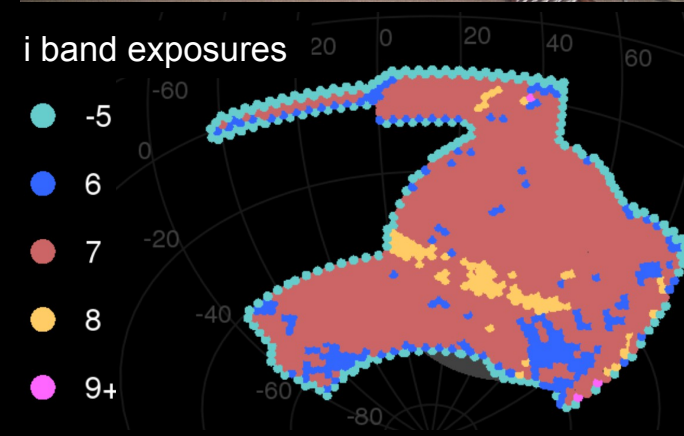
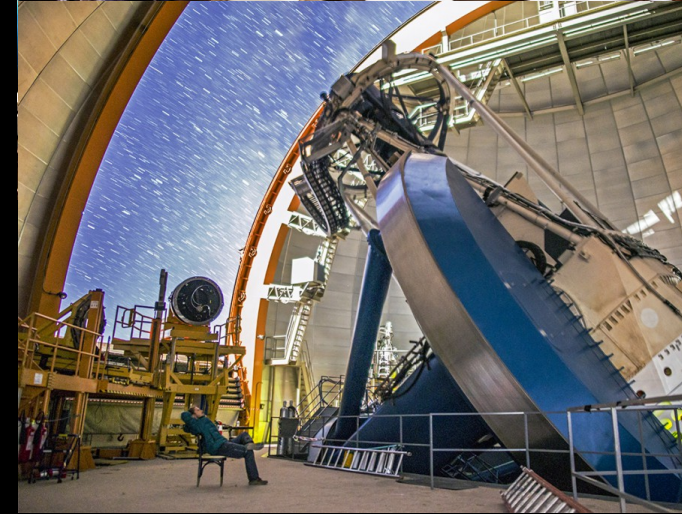
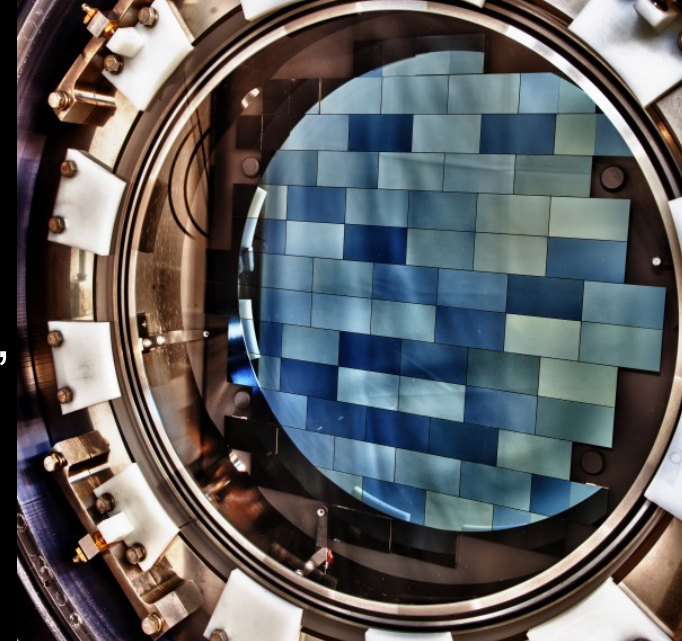
A crack in Λ CDM?

A systematic error?



The Dark Energy Survey

- 5000 sq. deg. survey in grizY from Blanco @ CTIO, 10 exposures, 5 years, >400 scientists
- Primary goal: dark energy equation of state
- Probes: Large scale structure, Supernovae, Cluster counts, Gravitational lensing
- Status:
 - SV (150 sq. deg, full depth): most science done, catalogs at <http://des.ncsa.illinois.edu>
 - Y1 (1500 sq. deg, 40% depth): data processed, results on cosmology today
 - Y3 (5000 sq. deg, 50% depth): data processed, vetting catalogs
 - Y4: data taking finished (70% depth)



Funded by:



U.S. DEPARTMENT OF
ENERGY

Office of
Science



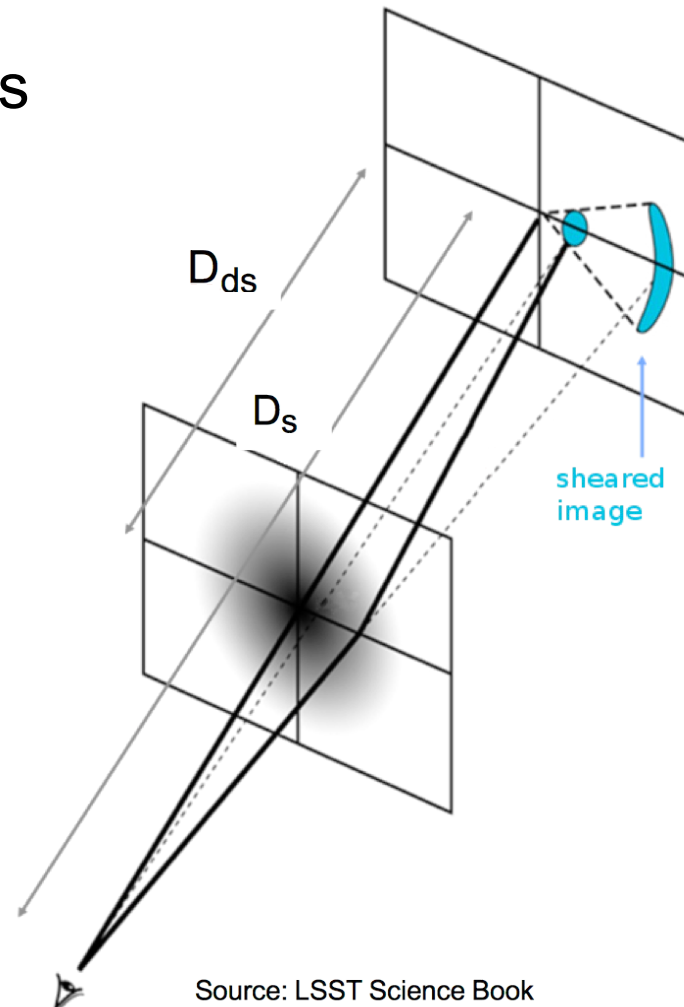
**Collaborating
institutions:**



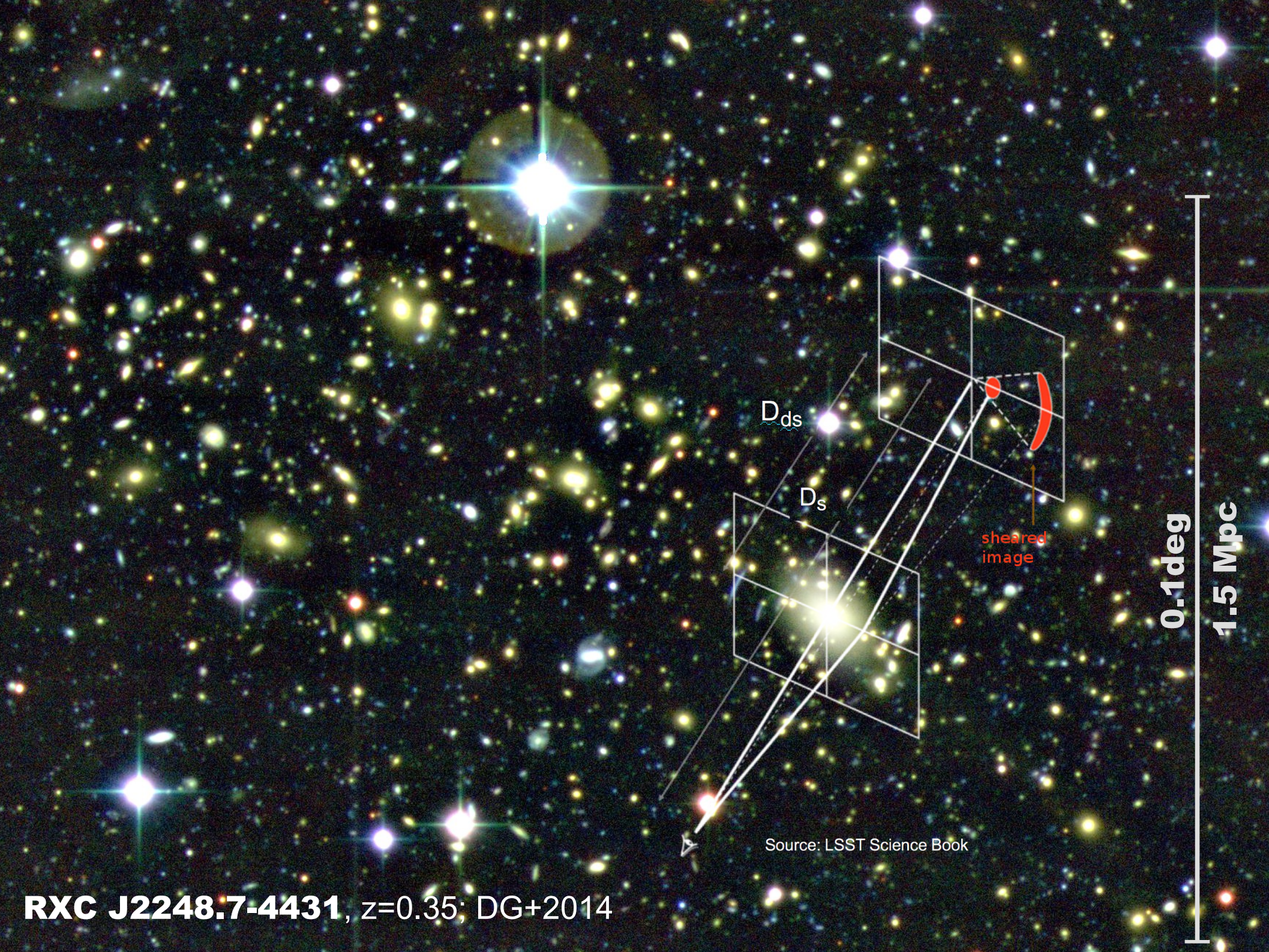
Gravitational lensing

- When light passes massive structures, it feels gravity and its path gets bent
- This causes shifting, and magnification, and shearing of the galaxy image

$$\gamma_t(\theta) = \langle \kappa(\theta') \rangle_{\theta' < \theta} - \kappa(\theta)$$
$$\kappa = \Sigma / \left[\frac{c^2}{4\pi G} \frac{D_s}{D_d D_{ds}} \right]$$



Source: LSST Science Book



D_{ds}

D_s

sheared
image

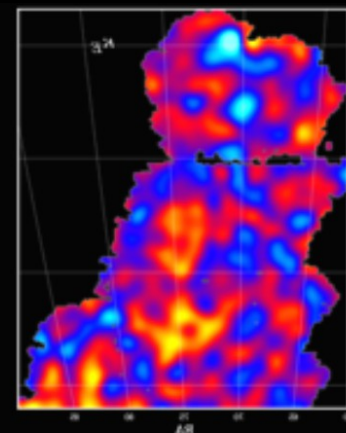
0.1deg

1.5 Mpc

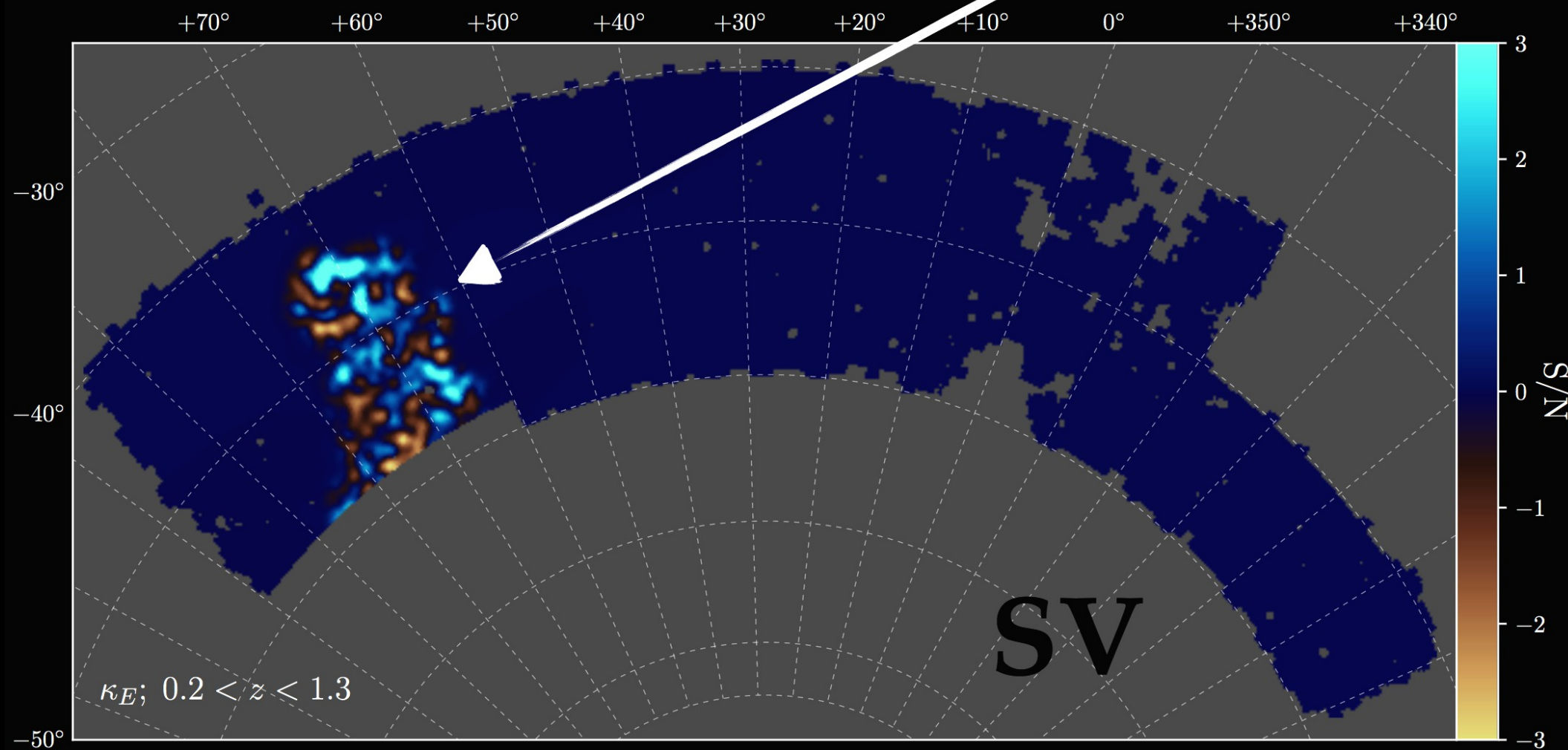
Source: LSST Science Book

RXC J2248.7-4431, $z=0.35$; DG+2014

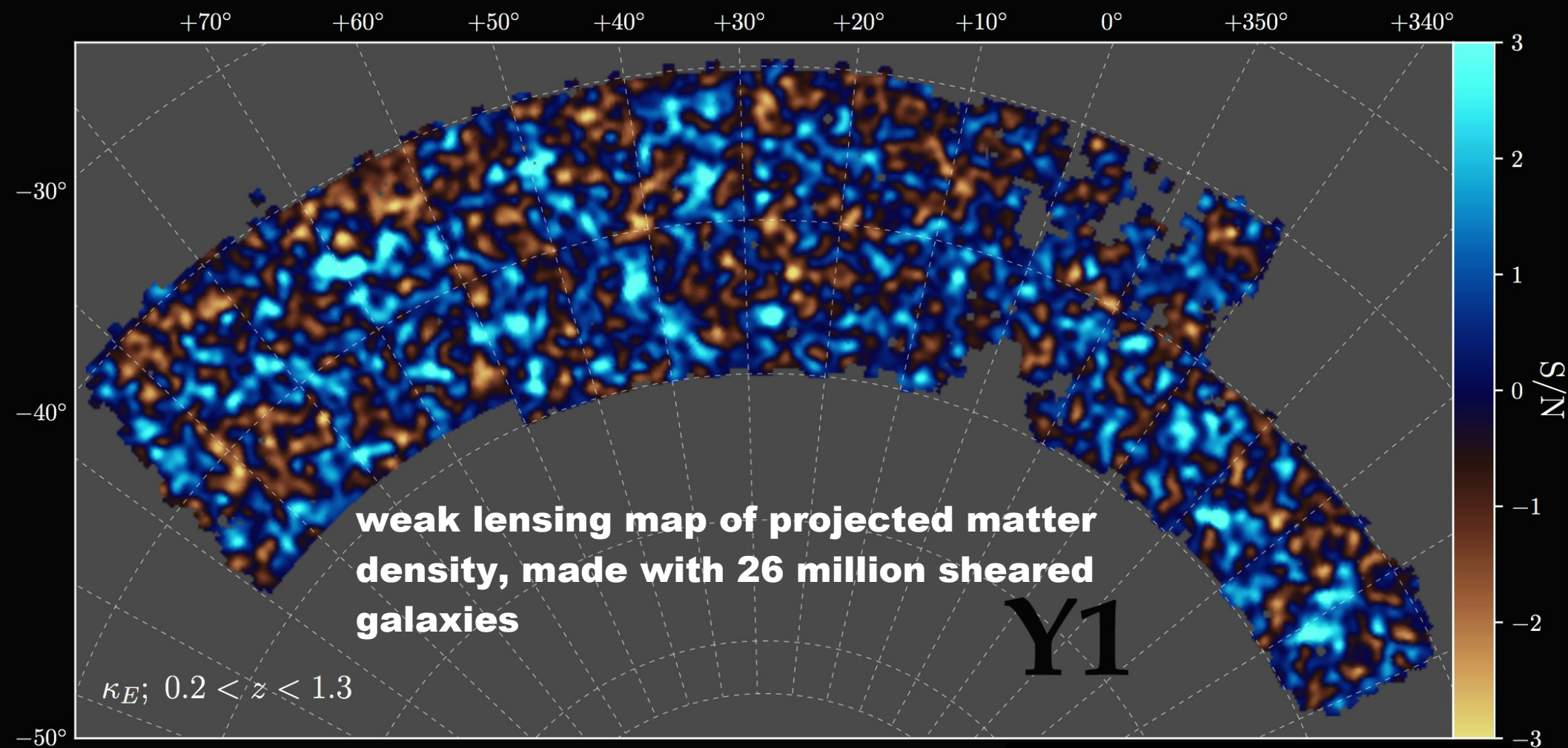
DES SV ...



Chang+;
Vikram+
2016



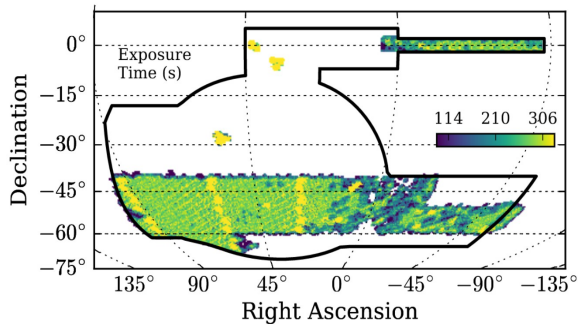
DES SV ... to Y1



Chang et al. (released today)

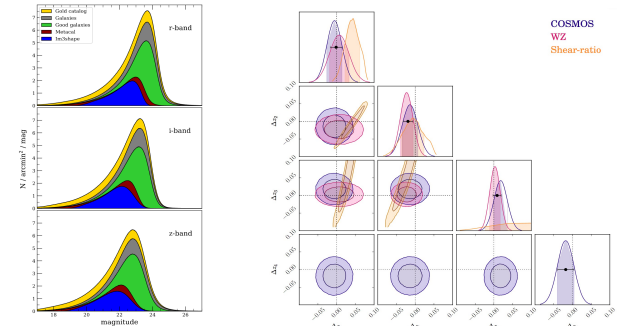
**With great statistical power comes
great systematic responsibility**

Unprecedented size and depth of photometric data



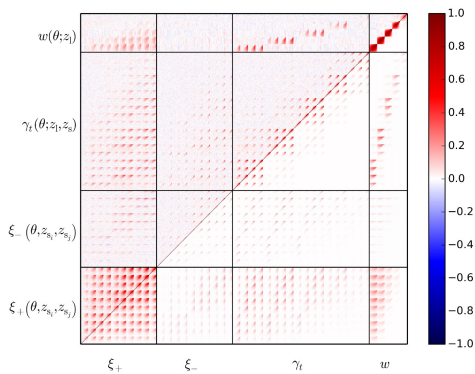
Drlica-Wagner, Rykoff, Sevilla+ released today

Two independent shape & photo-z catalogs and calibrations

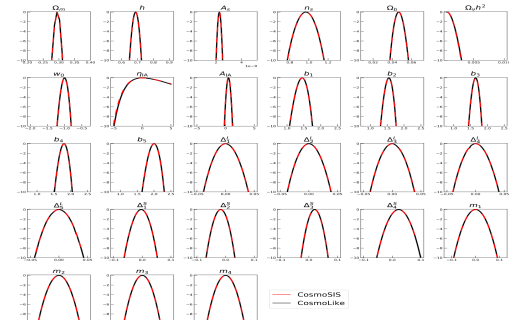


Zuntz, Sheldon+; Samuroff+; Hoyle, Gruen+ released today;
Davis+, Gatti, Vielzeuf+, Cawthon+ in prep.

Full, validated treatment of covariance and nuisance parameters (including v)



Theory and simulation tested, blind,
analysis with two independent codes,
CosmoLike and CosmoSIS

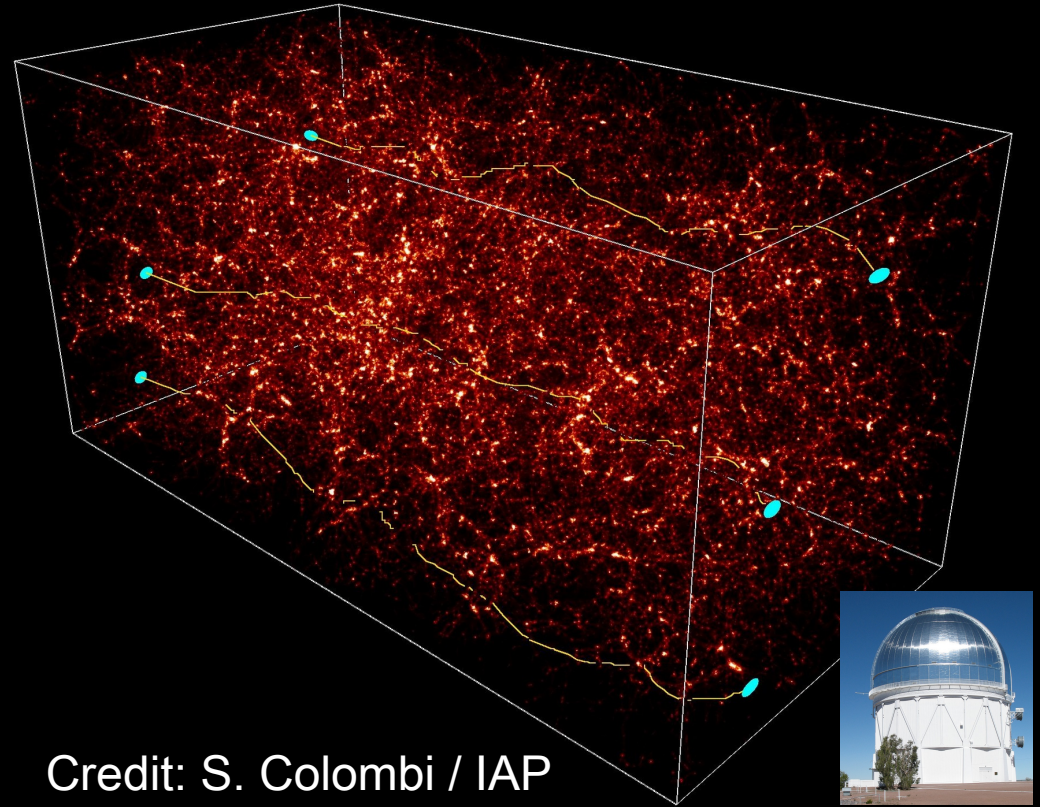


Krause, Eifler+2017; MacCrann, DeRose+ in prep

Measurements: cosmic shear

Troxel+ released today

- Light from distant galaxies passes the same foreground structure

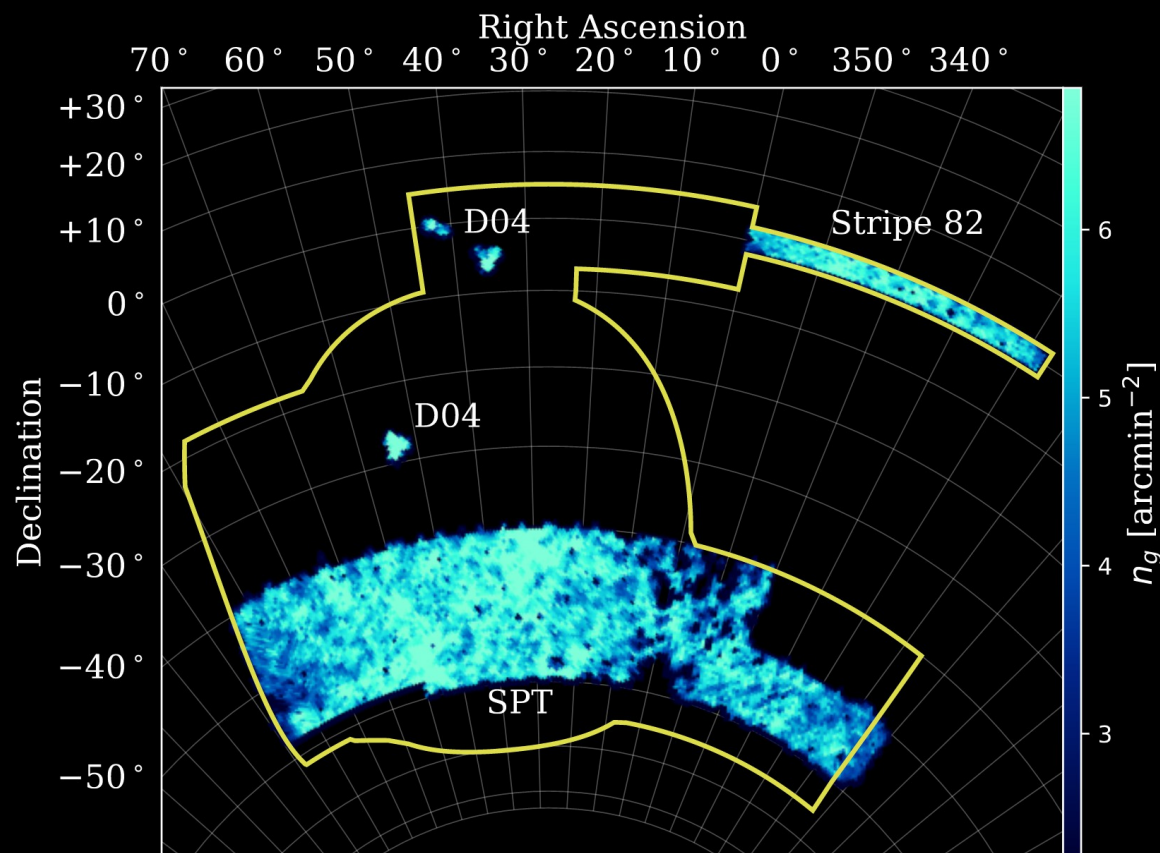


Credit: S. Colombi / IAP

Measurements: cosmic shear

Troxel+ released today

- Light from distant galaxies passes the same foreground structure
- We measure their shapes



35 million shapes in primary catalog with metacalibration
(Sheldon+2017; Huff+2017) with multiplicative bias $|m|$ below 1.3% (68% C.L.)

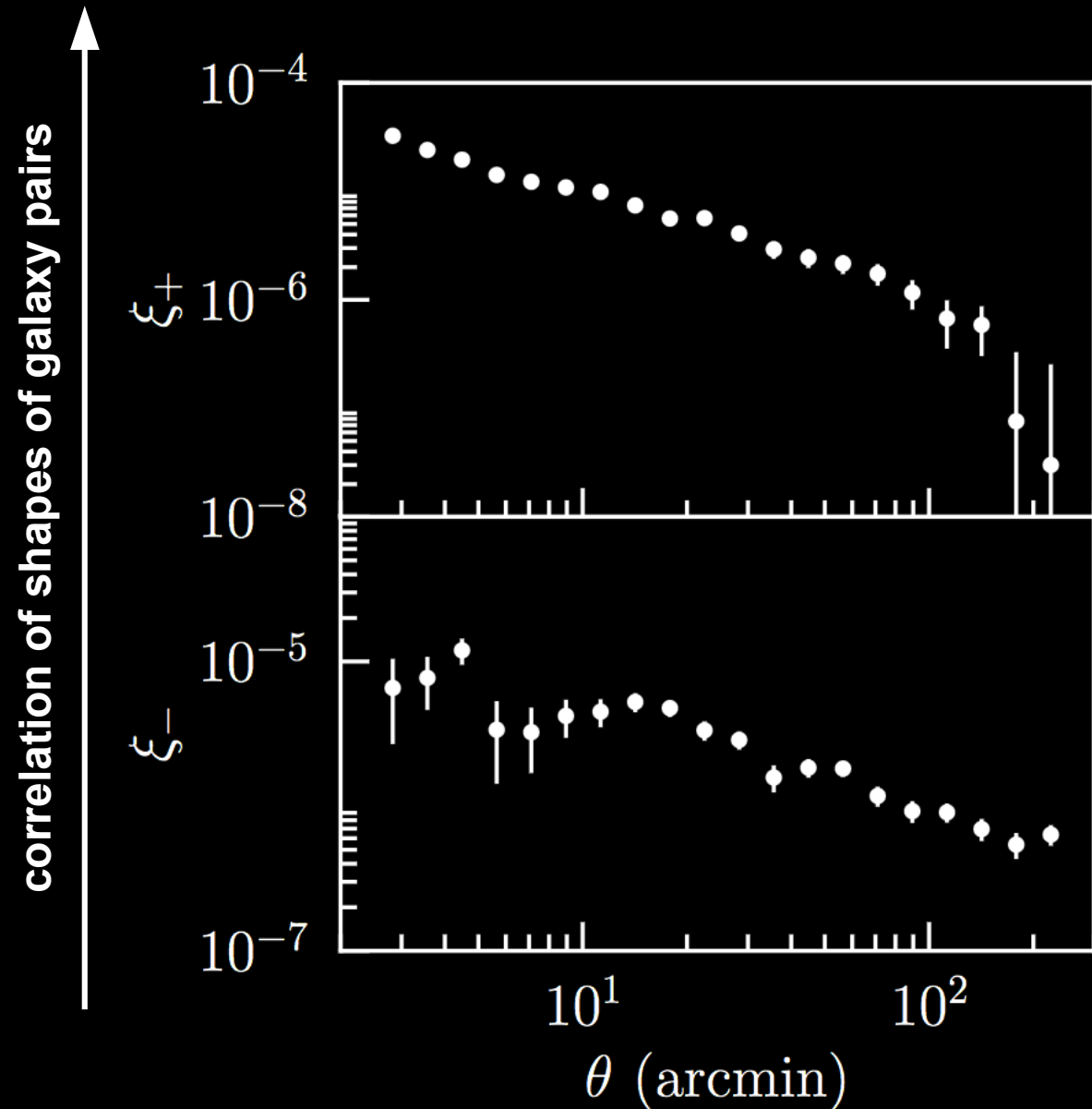
Independent im3shape (Zuntz+2013) catalog, calibrated with image simulations (Samuroff+ released today)

Suite of detailed tests: Zuntz, Sheldon+ released today

Measurements: cosmic shear

Troxel+ released today

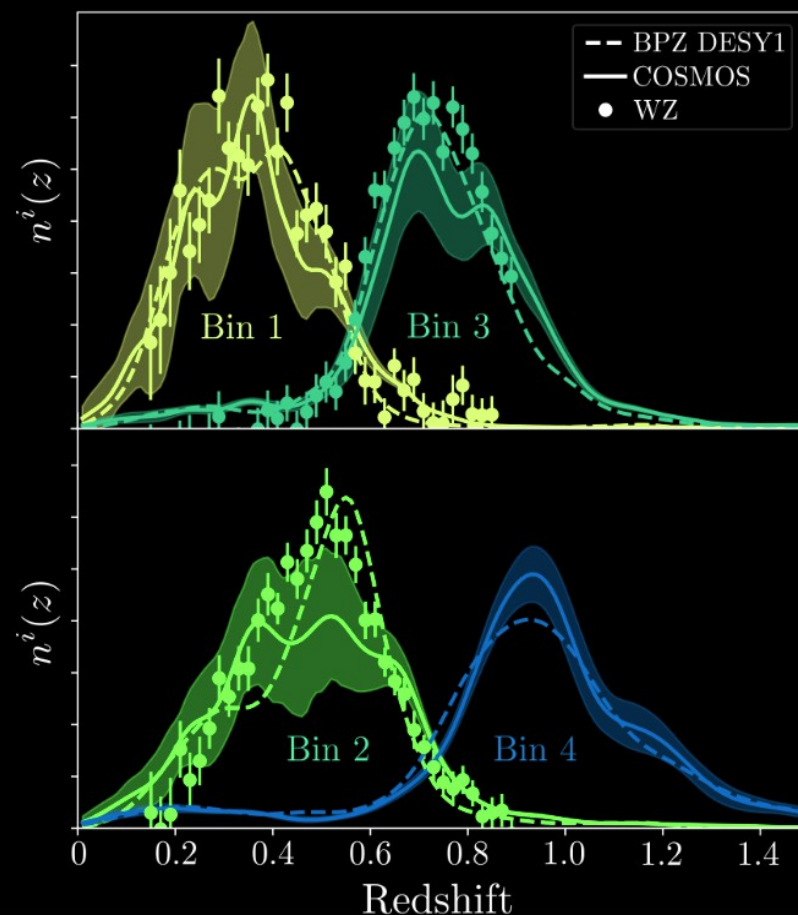
- Light from distant galaxies passes the same foreground structure
- We measure their shapes
- We measure the correlation of shapes of galaxy pairs



Measurements: cosmic shear

Troxel+ released today

- Light from distant galaxies passes the same foreground structure
- We measure their shapes
- We measure the correlation of shapes of galaxy pairs
- Using photometric redshifts, we do this tomographically



Redshift distributions $n(z)$ measured with independent methods [BPZ, Benitez+2000; DNF, de Vicente+2016; COSMOS, Laigle+2016]

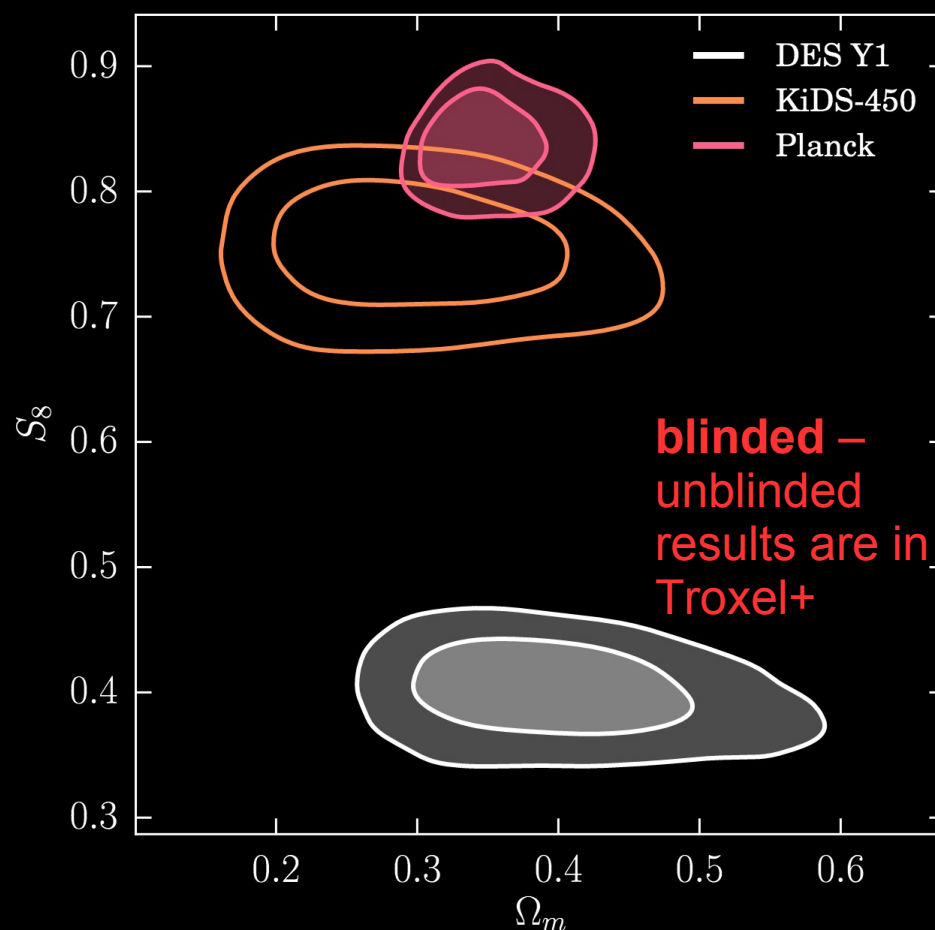
Bias in redshifts independently calibrated with COSMOS photometry and cross-correlations with LRGs (WZ).

Hoyle, Gruen+ released today; Cawthon+, Davis+, Gatti, Vielzeuf+ in preparation

Measurements: cosmic shear

Troxel+ released today

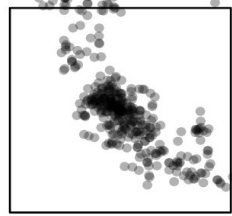
- Light from distant galaxies passes the same foreground structure
- We measure their shapes
- We measure the correlation of shapes of galaxy pairs
- Using photometric redshifts, we do this tomographically
- We constrain cosmological parameters – blindly first, before we pass all tests



most precise cosmic shear experiment to date!

galaxy field

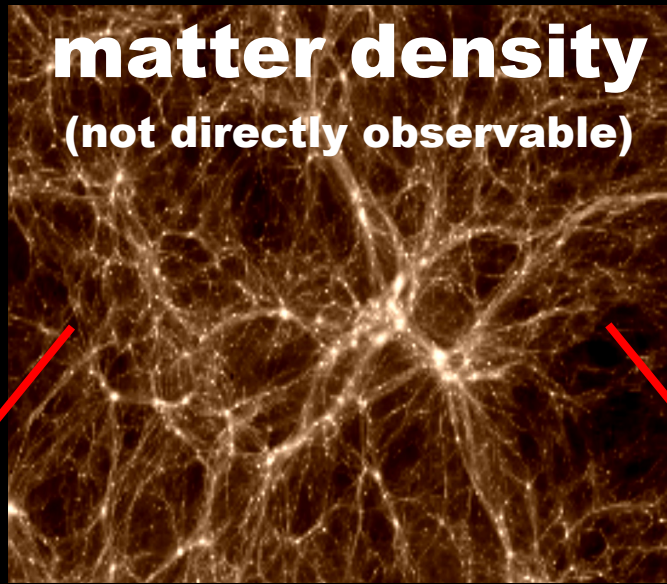
Melchior+2015



(1)

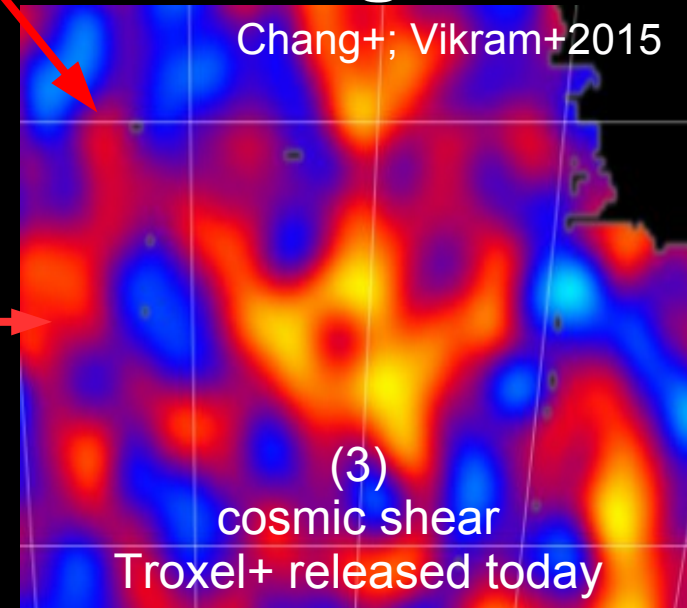
angular galaxy clustering
Elvin-Poole+ released today

matter density
(not directly observable)



lensing convergence

Chang+; Vikram+2015



(3)

cosmic shear
Troxel+ released today

(2)
galaxy-galaxy lensing
Prat, Sanchez+ released today

**combination of these three two-point functions maximizes use of information
and jointly and robustly constrains nuisance parameters**

[Hu&Jain 2004, Huterer+2006, Bernstein+2009, Joachimi&Bridle 2010, van Uitert+2017, Joudaki+2017]

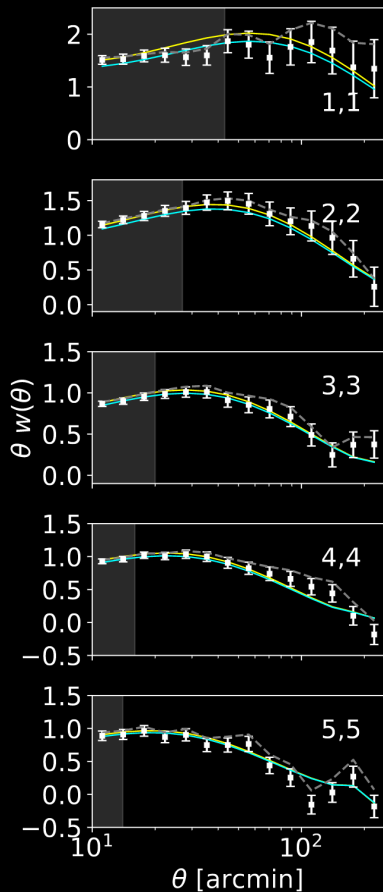
**largest individual data sets and joint constraints from these three probes for the first time:
DES Collaboration+ released today**

Measurements: galaxy clustering and galaxy-galaxy lensing

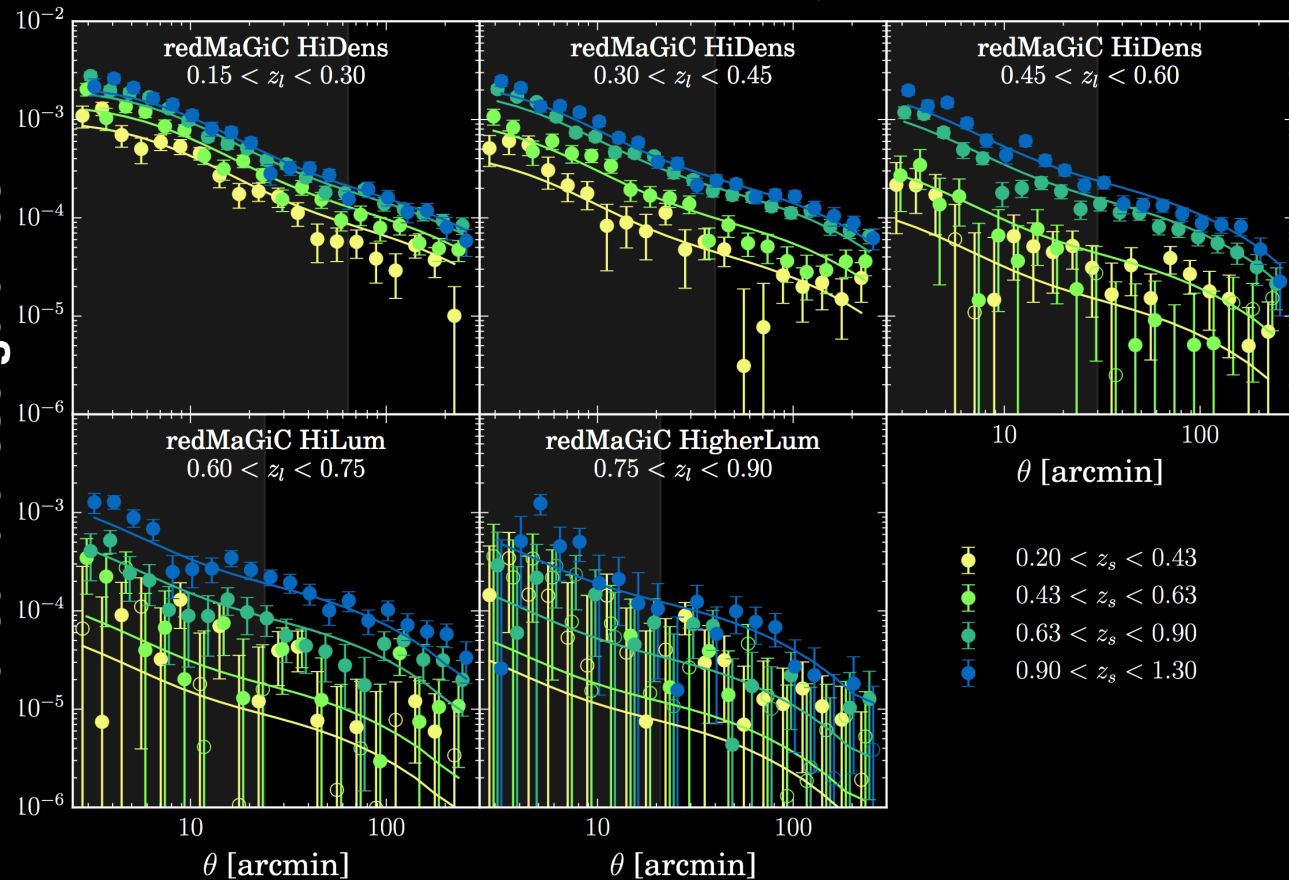
Elvin-Poole+; Prat, Sanchez+ released today

- Lens galaxies: redMaGiC LRGs with high-quality photometric redshift estimates (Rozo, Rykoff+2016)

clustering of galaxies in 5 redshift bins between $z=0.15 \dots 0.90$



tangential gravitational shear around these galaxies



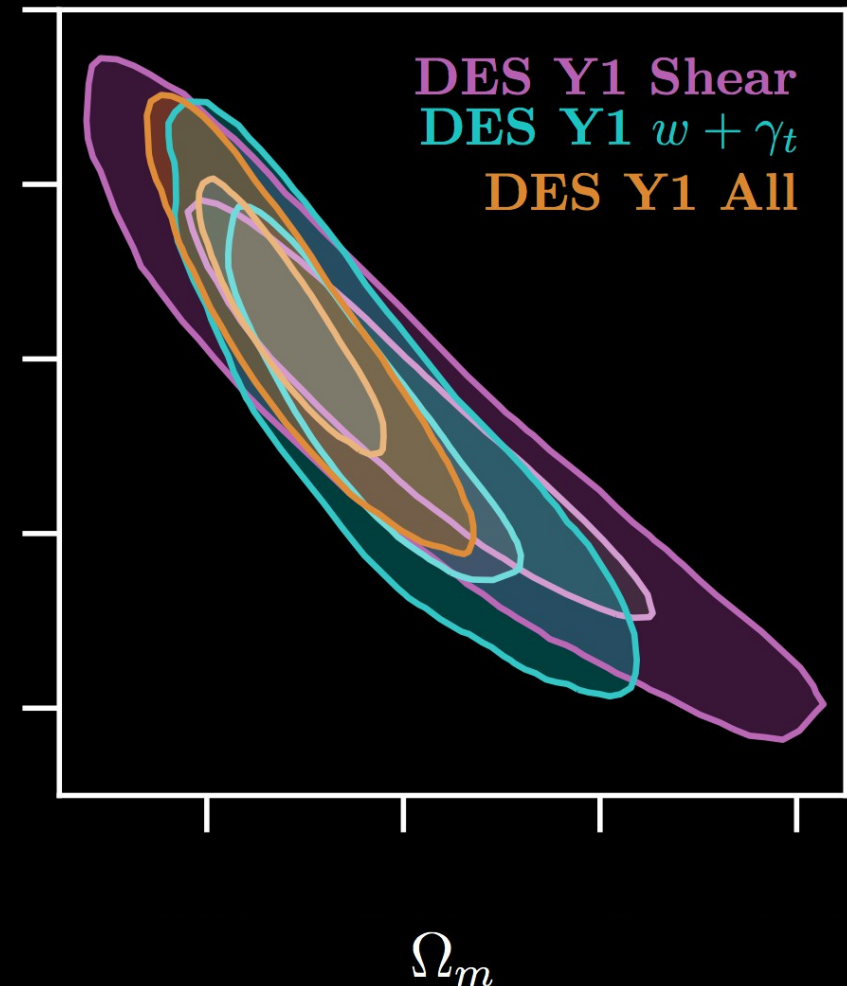
Consistency of the individual constraints in Λ CDM

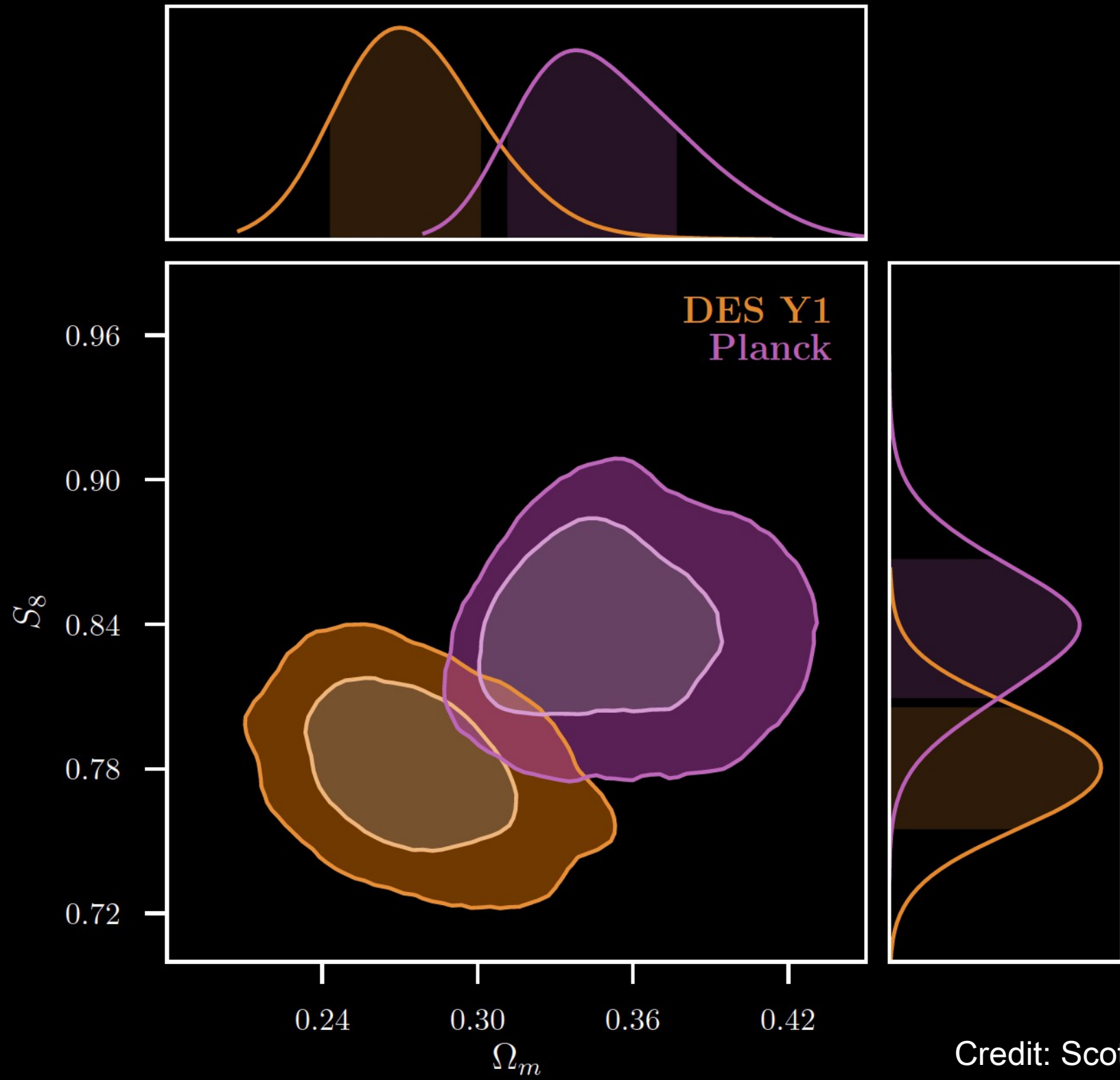
- Cosmic shear and redMaGiC clustering + lensing yield consistent cosmological constraints

- Criterion:
Bayes Factor

$$R = \frac{P(\vec{D}_1, \vec{D}_2 | M)}{P(\vec{D}_1 | M) P(\vec{D}_2 | M)} = 2.8 > 0.1$$

- passing 11 other null tests, we unblind

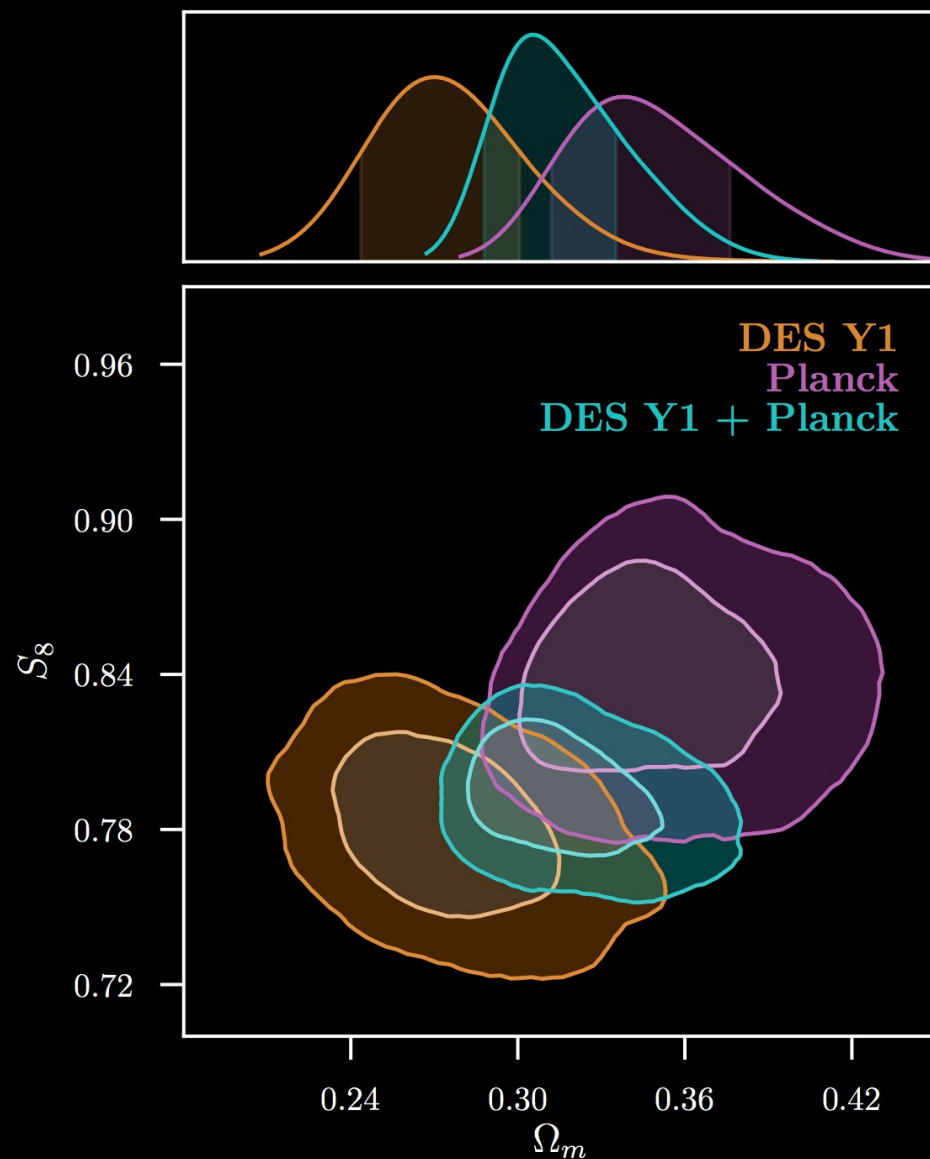




Credit: Scott Dodelson

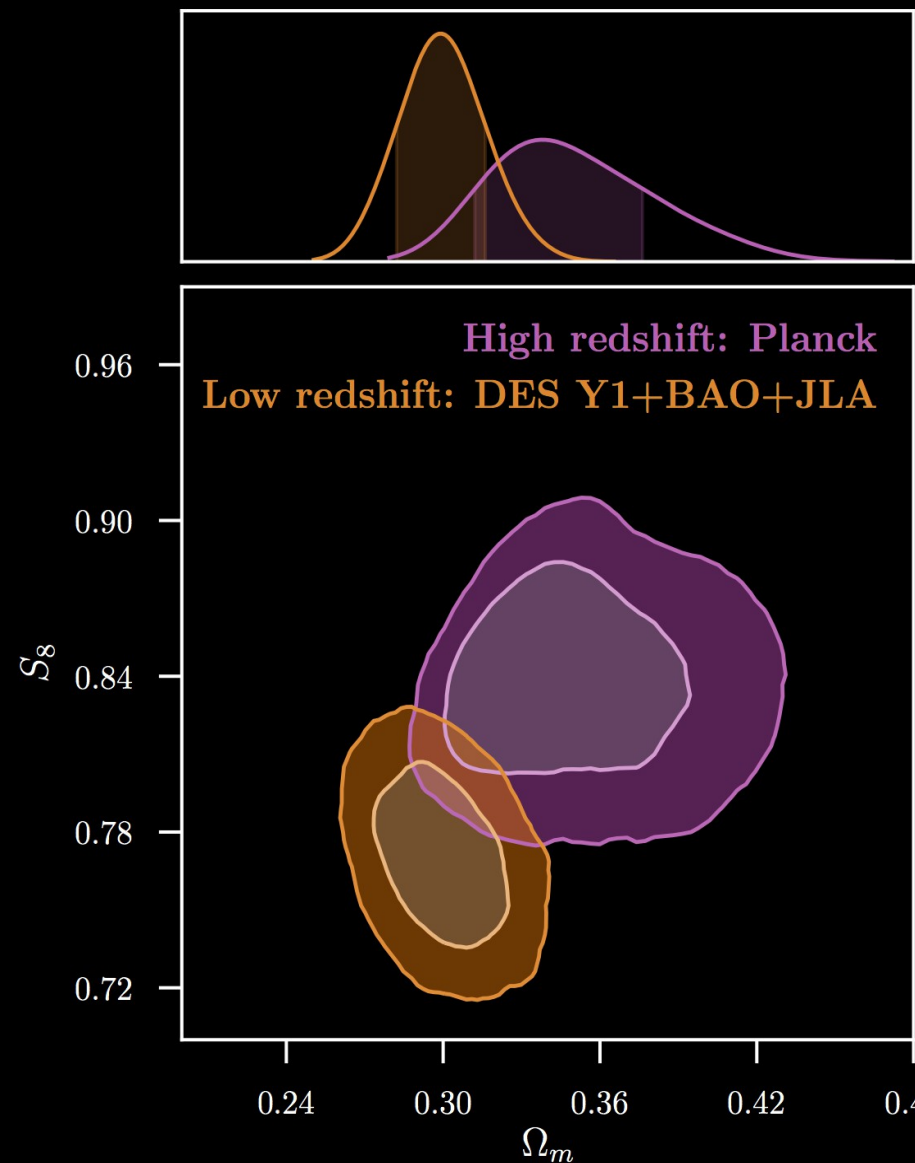
Key result: Consistency of late Universe with Planck in Λ CDM

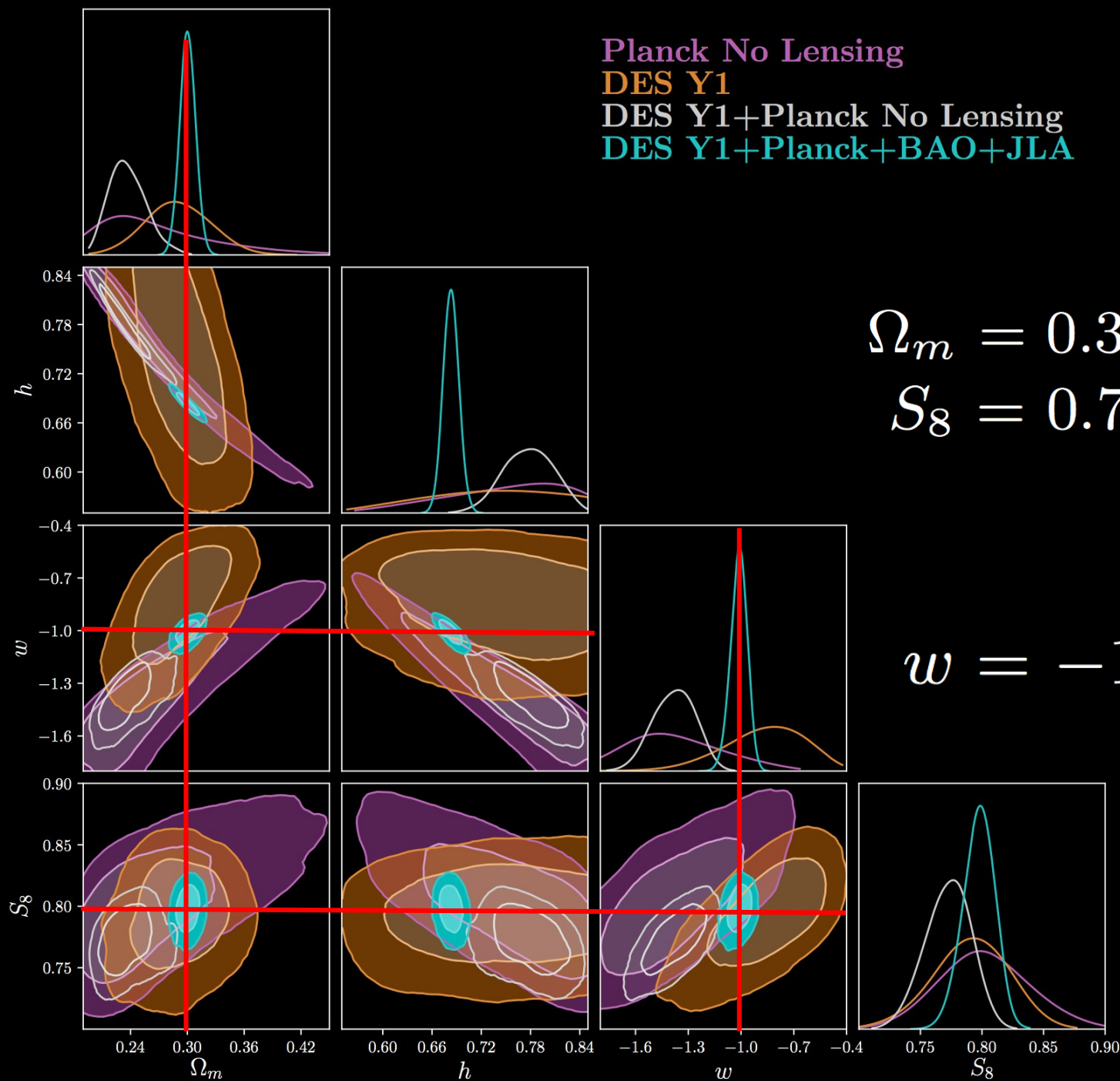
- DES and Planck constrain matter density and S_8 with equal strength
- Difference in central values $1-2\sigma$ in the same direction as earlier lensing results
- Bayes Factor 4.2 – no evidence for inconsistency



Key result: Consistency of late Universe with Planck in Λ CDM

- DES and Planck constrain matter density and S_8 with equal strength
- Difference in central values $1-2\sigma$ in the same direction as earlier lensing results
- Bayes Factor 4.2 – no evidence for inconsistency
- Still consistent ($R=9.0$) for joint low-z results + Planck, which is why we combine...





Key result: DES + geometry + CMB yields consistent, tightest constraints

- consistent constraints from geometric probes + DES (R=244)

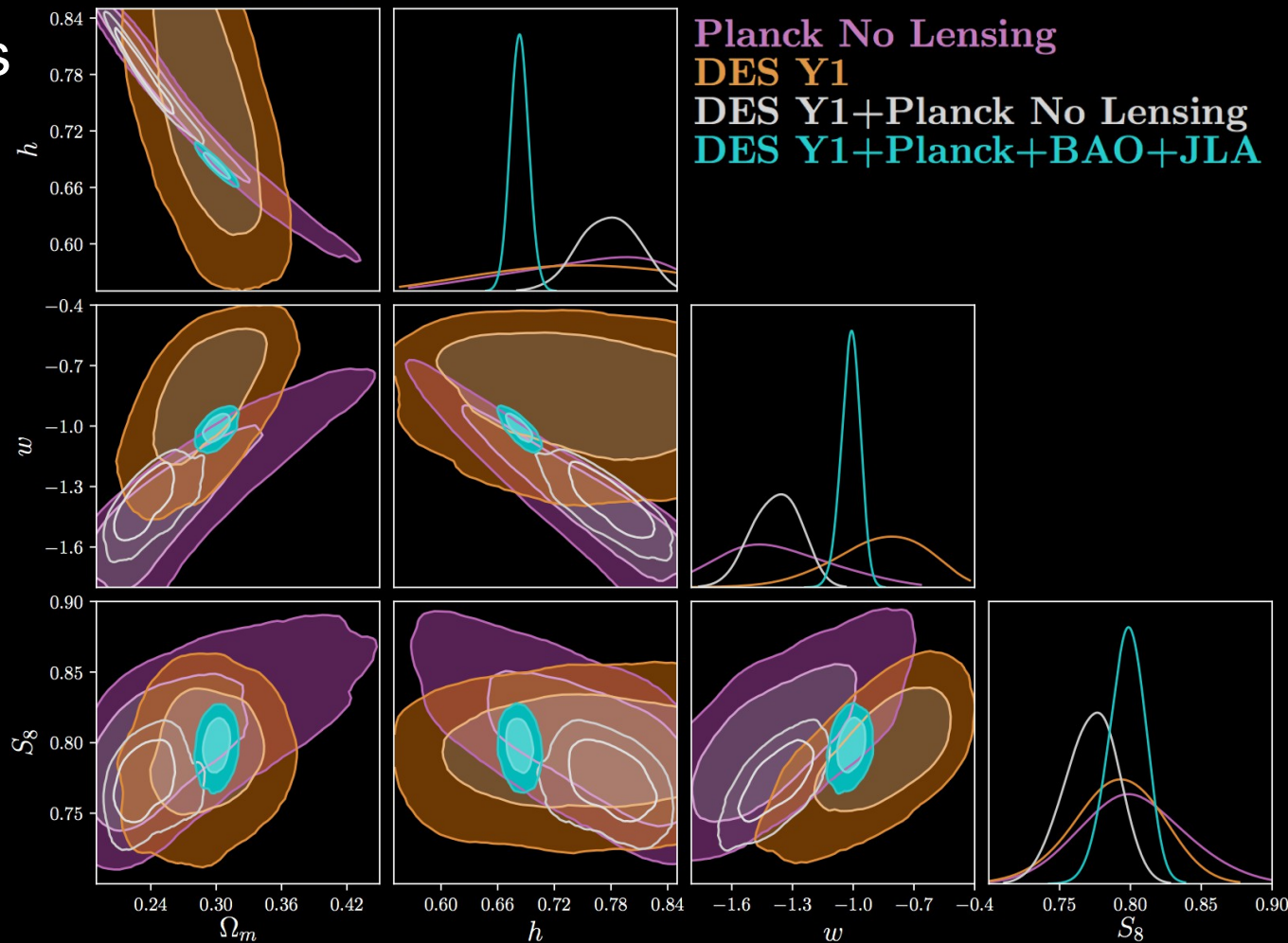
- most precise measurements in Λ CDM:

$$\Omega_m = 0.301^{+0.006}_{-0.008}$$

$$S_8 = 0.799^{+0.014}_{-0.009}$$

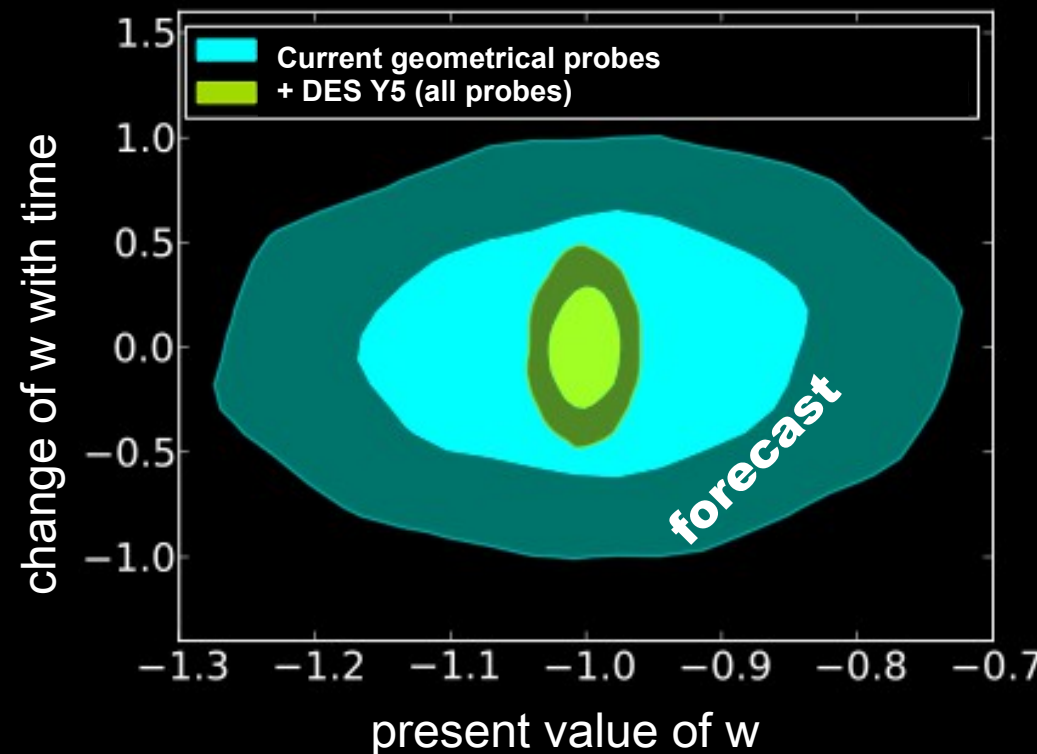
- no evidence for $w \neq -1$ in any combination

$$w = -1.00^{+0.04}_{-0.05}$$



Steps forward: more precise tests of broader range of models

- This is a precise test of Λ CDM, and it shows any potential discrepancies are smaller than its uncertainty
- It does not explain Λ CDM
- It is not very sensitive to models with time-varying Dark Energy equation of state (among others)
- Future joint analyses will be!



Credit:
T. Eifler, E. Krause, J. Frieman

Summary

- Wide range of probes from early & late Universe, geometry & structure, agree on fiducial Λ CDM cosmology
- DES has added the most precise measurement of structure in the evolved Universe
 - Competitiveness and consistency with Planck CMB in Λ CDM, insignificant offset in the direction of other lensing studies
 - Precise joint measurements close to $\Omega_m=0.30$, $\sigma_8=0.80$, $w=-1.0$
 - Papers released now on <http://www.darkenergysurvey.org/>
 - DES Collaborators in the audience happy to discuss later!
- Need even more precise & different model tests to understand Dark Energy – work in progress!